

AST 358 HWk 2 / Q 1

(a) $\frac{GM_R m}{R^2} = \frac{m v^2}{R}$ for a star of mass m , speed v

$$M_R = \frac{v^2 R}{G}$$

$$\left(\frac{M_R}{M_\odot}\right) = 2.2 \times 10^5 \left(\frac{v}{\text{km s}^{-1}}\right)^2 \left(\frac{R}{\text{kpc}}\right) \quad (1)$$

For $R = 10 \text{ kpc}$, $v = 300 \text{ km s}^{-1}$

$$M_R = 1.98 \times 10^{11} M_\odot$$

(b) $\left(\frac{M_R}{L}\right)_{\text{gal}}$ interior to radius $R = \frac{1.98 \times 10^{11}}{3.3 \times 10^9} = 60 \frac{M_\odot}{L_\odot}$

(c) $\left(\frac{M}{L}\right)_* = 6 \left(\frac{M_\odot}{L_\odot}\right)$

Fraction of total mass not in stars

$$= \frac{60 - 6}{60} = 90\%$$

(d) $M_R = (M_{\text{gas}} + M_{\text{dust}}) + M_* + M_{\text{dark}}$

$$\begin{aligned} M_{\text{dark}} &= 1.98 \times 10^{11} - 5 \times 10^9 - 2 \times 10^{10} \\ &= 1.73 \times 10^{11} M_\odot \end{aligned}$$

(e) $\frac{M_{\text{dark}}}{M_R} = \frac{1.73}{1.98} = 87.4\%$

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(a)
$$V_c^2(R)/R = GM(R)/R^2$$
 Centripetal acceleration = accel. due to force of gravity

$$M(R) = V_c^2(R) \frac{R}{G} \quad (1)$$

At $R = R_h \rightarrow M(R_h) = M_{Tot} = \frac{V_0^2 R_h}{G}$

(b) At $R > R_h$, $M(R)$ in (1) = M_{Tot}

$$\rightarrow V_c^2(R) = M(R) \frac{G}{R} = M_{Tot} \frac{G}{R} = V_0^2 \frac{R_h}{R}$$

$$V_c(R) = V_0 \sqrt{\frac{R_h}{R}} \quad (2)$$

(c)
$$\begin{aligned} \bar{\Phi}(R) &= \text{Work done in bringing unit mass from } \infty \text{ to } R \\ &= \int_{\infty}^R F(r) dr = \int_{\infty}^R \frac{GM(r)}{r^2} dr \\ &= \int_{\infty}^{R_h} \frac{GM(r)}{r^2} dr + \int_{R_h}^R \frac{GM(r)}{r^2} dr \quad \text{for } R_h \leq R \leq R_h \\ &= \int_{\infty}^{R_h} \frac{GM_{Tot}}{r^2} dr + \int_{R_h}^R \frac{V_c^2(r)}{r} dr \end{aligned}$$

Use (1) and (2)

$$\begin{aligned} &= V_0^2 R_h \int_{\infty}^{R_h} \frac{1}{r^2} dr + V_0^2 \int_{R_h}^R \frac{1}{r} dr \\ &= -V_0^2 R_h \left[\frac{1}{r} \right]_{\infty}^{R_h} + V_0^2 \left[\ln r \right]_{R_h}^R \end{aligned}$$

$$\rightarrow \bar{\Phi}(R) = -V_0^2 \left[1 + \ln(R_h/R) \right] \quad \text{for } R_h \leq R \leq R_h \quad (3)$$

④ Escape speed v_e satisfies

$$\frac{1}{2} v_e^2 = |\bar{\Phi}(R)| \quad \text{from conservation of energy} \quad (4)$$

For $R_e \leq R \leq R_h$ use $\bar{\Phi}(R)$ from (3)

$$v_e^2(R) = 2V_0^2 \left[1 + \ln(R_h/R) \right]$$

$$\ln(R_h/R) = \frac{1}{2} \left(\frac{v_e(R)}{V_0} \right)^2 - 1 \quad (5)$$

$$\text{At } R = R_0, \quad v_e(R_0) \geq (500 \text{ km s}^{-1})^2$$

$$\rightarrow \ln(R_h/R_0) \geq \frac{1}{2} \left(\frac{500}{220} \right)^2 - 1$$

$$R_h \geq 4.9 R_0 \approx 41.3 \text{ kpc}$$