

# THE GALAXIES AND COSMOS EXPLORER TOOL

## CHARTING GALAXIES OVER COSMIC TIMES IN THE CLASSROOM

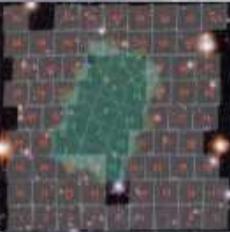
Shardha Jogee, Mary Kay Hemenway, Sarah Miller, Aaron Smith,  
Achal Augustine, Kyle Fricke, Randi Worthach, Sandi Preston, Dan Lester  
University of Texas at Austin

[WWW.AS.UTEXAS.EDU/GCET](http://www.as.utexas.edu/gcet/)

### SUMMARY

Recent large galaxy surveys conducted with NASA's Hubble Space Telescope Advanced Camera for Surveys (ACS) have provided unprecedented legacy datasets which allow astronomers to chart the evolution of galaxies over a large fraction of the age of the Universe. The Galaxies and Cosmos Explorer Tool (GCET; <http://www.as.utexas.edu/gcet/>) is an online web-based tool that allows the general public and students to actively participate in this exciting adventure through quantitative analyses of HST images from the Galaxy Evolution from Morphology and SEDs (GEMS) survey, one of the widest-area galaxy surveys conducted in two filters with ACS to date. The tool allows users to sort the vast cosmos and access ACS images of over 8,000 galaxies over the last eight billion years. For galaxies of interest, users can measure the size, determine the lookback time for concordance cosmology, perform morphological classification on images at two rest-frame wavelengths, and gauge the different stellar populations present. Users can record their measurements as well as reference information such as coordinates and redshift of each galaxy into Excel spreadsheets for further analysis. The celestial coordinates can be used to extract further multiwavelength data from existing archives and upcoming virtual observatories. For undergraduate classes, more advanced IDL or C-based analyses that employ the full samples can be combined with the visualization capabilities of GCET in order to explore the nature of interesting objects, such as the most massive galaxies, star-forming systems, as well as interacting and merging galaxies. GCET provides a powerful tool for discovery learning in undergraduate introductory science classes as well as high schools.

We thank the GEMS collaboration, and acknowledge support from NASA grants NAG5-13043 and NASA NNG06GBR00G, NSF grant AST-0507748, and the Faculty And Student Teams for Technology (FAST Tex) award from the University of Texas Division of Instructional Innovation and Assessment (DIIA).

(TO FIG. 1)  from the home page - an overview of the 30 by 30' area mapped at high resolution and in two filters by the GEMS survey (de Jong et al. 2004) using the ACS camera aboard the Hubble Space Telescope. The 25-arcsecond mosaic of GEMS overlaps with the central 15 filters covered by the GOODS (Giavalisco et al. 2004) survey. Results are from the COMBO-17 gravitationally lensed data (Wolf et al. 2004).  The 37-HST image (the superior spatial resolution (0.07") resolves components of galaxies.  The 37-ground-based image shows students can observe how the low resolution (1-10') ground-based galaxies in comparison with the space-based HST.

### EXAMPLE OF GCET APPLICATION

Advanced students are able to use GCET as a visual aid in their analysis of GEMS data. While learning to create plots in programming languages such as IDL, students can use GCET to view some of the more interesting galaxies.

For example, if a student were to make a plot of the star formation rate vs. mass in a specific redshift regime of interest (Fig. 2), the student could then extract the most massive, star-forming galaxies to view with GCET (Fig. 10). This allows the student not only to take their analysis further with visual classification, but also to view these galaxies and systems in their environmental context on the GEMS tile. Students can see for themselves if the most massive star-forming galaxies occur in clusters or in relative isolation. Using the GCET measuring tools, students will be able to calculate the sizes of certain characteristics of interest within the galaxy, for instance a star forming region.

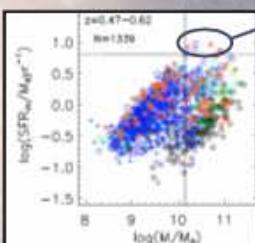


FIGURE 1 - Detailed structure elements are resolved in the GEMS data, and students are able to measure features like bulges, disks, spiral arms, and color-color plots for interacting pairs with sizes of up to 10'' on a elliptical galaxy with  $\approx 0.7$   $M_{\odot}$  at an interacting system with  $\approx 0.3$ . The images are a composite of both the F606W and F814W bands.

FIGURE 2 - The individual galaxy window enables students with measuring tools for images from both the F606W and F814W bands. Comparing these images allows students to introduce band-pass shifting and its effects. Students are able to click the "Measure Age" button which converts the galaxy's width to lookback-time, which is then recorded in the "Analysis" window (Fig. 10).

FIGURE 3 - Students are able to save their measurements and visual classification for each galaxy image, and all of that data can be downloaded later into a spreadsheet for further analysis.

