

In previous lectures, we answered the first two questions in structure formation: 02-18-2016

① what are the initial fluctuations?

and ② How do the perturbations grow?

Note: Growth in

• static medium : exponential
 $\delta \propto \exp(-t/t_{\text{ff}})$

→ Initial conditions don't matter!

• Expanding medium : power-law
 $\delta \propto t^{2/3}$ (for flat universe).

→ Initial conditions matter !!

In this lecture, we will start answering the last questions we posted about structure formation —

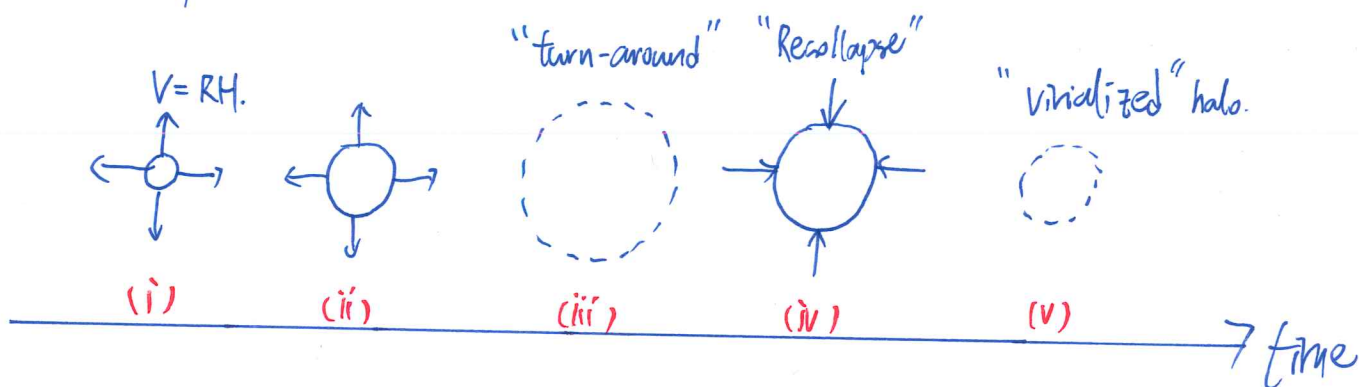
③ what happens when $\delta \gtrsim 1$?

→ "halo" formation

OR, → "Virialization"

First, let's look at a cartoon picture of the virialization process:

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- (i). Initially, the overdense region expands with the background universe according to the Hubble's law.
- (ii). The expansion slows down, "detaching" from the background universe.
- (iii). At "turn-around", the overdensity stops expanding.
- (iv). It starts to recollapse under its own gravity.
- (v). "Virialization" — the recollapsing dark matter particles hit the central point ($R=0 \rightarrow$ "singularity") because of random motion.
 $\rightarrow$ establish the state of "Virial Equilibrium" V.E.

where

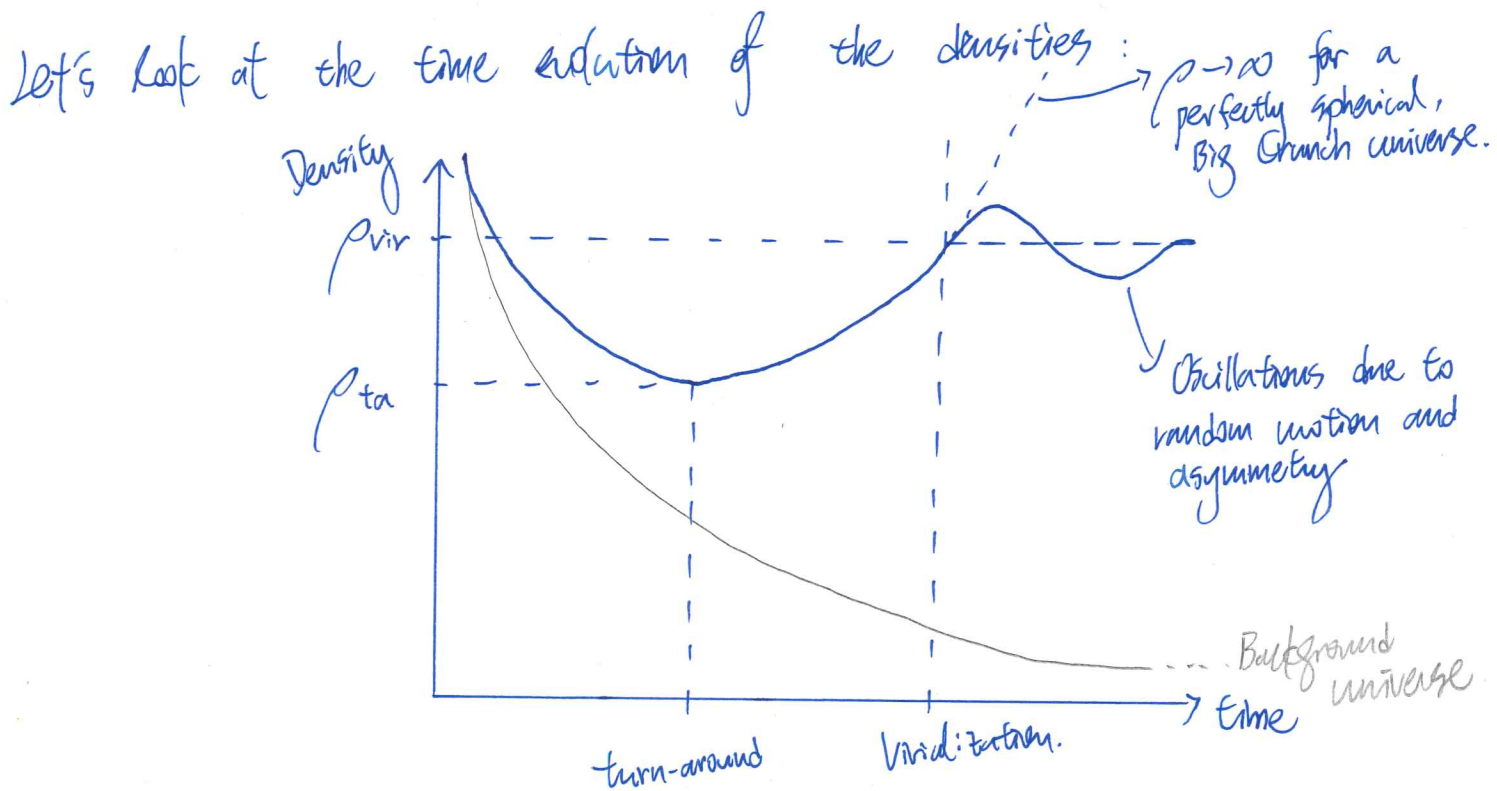
$$\boxed{2 E_{\text{kin}} \simeq - E_{\text{pot}}} \quad \text{"Virial Theorem"}$$

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Q: Why is Virial Theorem so important in astrophysics?

A: For spherical system, VT reads:

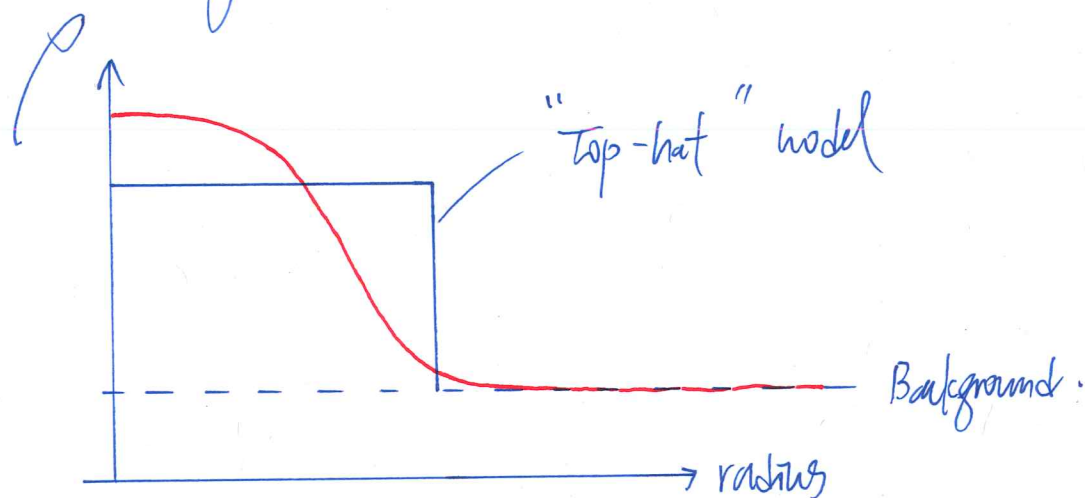
$$V^2 = \frac{GM}{R}$$

It helps us to get order of magnitude estimation of observed systems as the quantities are connected to observables.



Note: radial density profile in
overdensity

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- Idea: embed mini-Big Crunch (PS1, #2) into
simple background universe.

↳ Einstein-de-Sitter [E-ds] model.

Review: Basics of E-ds

→ flat (non-curved) universe with
only matter.

$$\rho_{\text{tot}} = \rho_m$$

$$\tilde{E}_{\text{tot}} = \tilde{E}_{\text{kin}} + \tilde{E}_{\text{pot}} = 0.$$

Friedmann Equation:

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3} \rho_{\text{tot}} \quad (k=0)$$

$$\rho_{\text{tot}} = \rho_m = \rho_{m,0} a^{-3}$$

$$a \propto t^{2/3}$$

$$\Rightarrow H(t) = \frac{\dot{a}}{a} = \frac{2}{3t}$$

$$\rho(t) = \frac{3H^2(t)}{8\pi G} = \frac{1}{6\pi G t^2}.$$

We will continue to derive ρ_{vir} in the next lecture.

