Life in the Outer Solar System

### Jupiter



Big  $R = 11R_{\oplus}$ 

Massive  $M = 300 M_{\oplus}$ 

 $= 2.5 \times \text{all the rest}$ 

Day about 10 Earth hours

Year about 12 Earth years

Thick Atmosphere, mostly H<sub>2</sub>, He

But also more complex molecules

Colors, storms

Like Miller - Urey

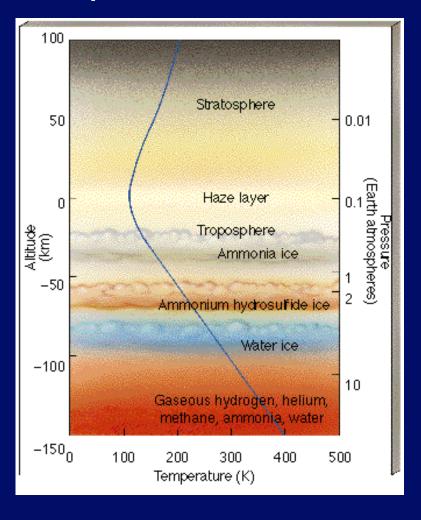
### Life in Jupiter Atmosphere?

Sagan-Salpeter, etc.

Sinkers (Plankton)

Floaters (Fish)

Hunters (Fish)



### Galileo Results on Jupiter

Reached Jupiter Dec. 1995 Sent probe into Jupiter's atmosphere at 100,000 mile/hour Decelerated at 230 g Lasted for 57 min.

Found: Strong winds

Turbulence, little lightning

Surprise: Little or no H<sub>2</sub>O May have entered in an unusual place (fewer clouds)

Life less likely?

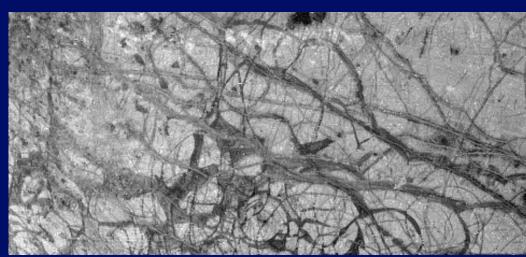




### Europa (Moon of Jupiter)

Surface: Fractured Ice Subsurface Oceans?

(Heated from Inside)

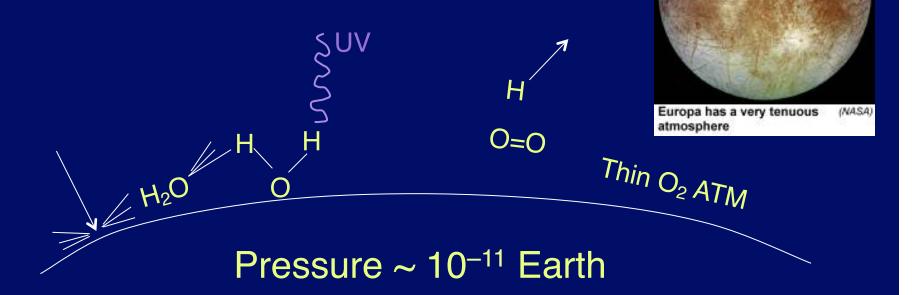


Close-up of "ice floes"

## Galileo - Jupiter's Moons

http://www.jpl.nasa.gov/galileo/index.html

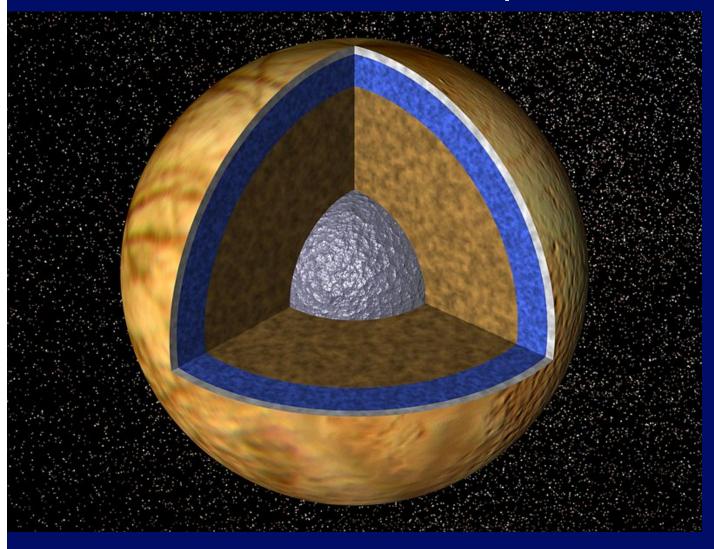
Europa has a (THIN!) atmosphere



More evidence for resurfacing along cracks by

Organic molecules on Callisto & Ganymede, maybe Europa?

# Model of Europa's Interior



Ice crust may be 10-30 km thick. Ocean may be 90 km deep.

#### Future Missions

- Juno, launched in 2011
  - Will go into polar orbit, map gravity field
  - Determine if Jupiter has a rocky core
- Jupiter Icy Moon Explorer (JUICE)
  - ESA Selected in 2012
  - Launch 2022, arrive 2030
  - Ganymede, Callisto, Europa orbiter
  - Look for evidence of organic molecules

### Saturn



- Big (9.4 R⊕)
- Massive (95 M⊕)
- Year 29.5 earth years
- Day 0.43 earth days
- Composition similar to Jupiter



#### Titan

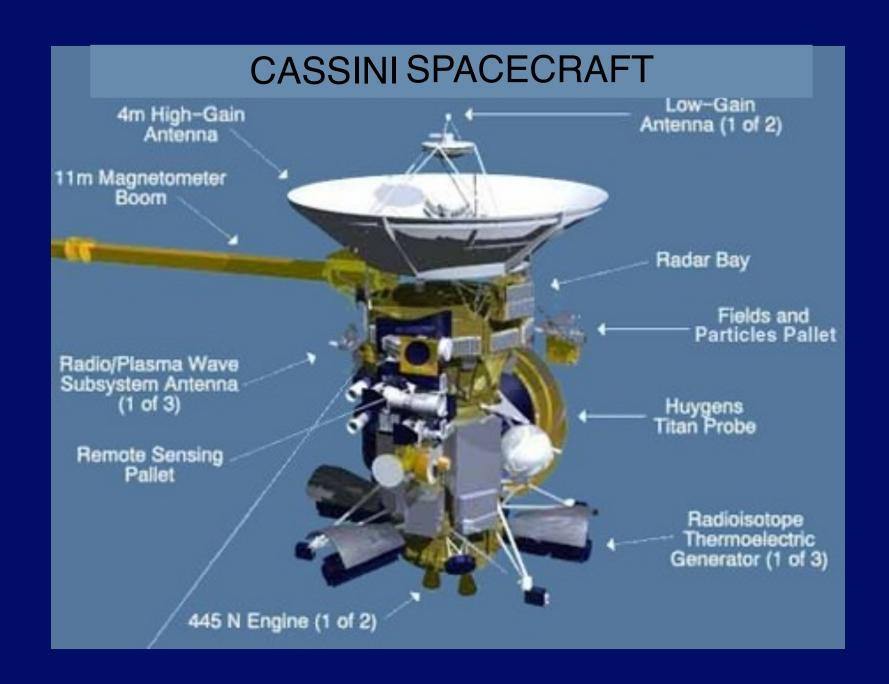
- Moon of Saturn
- Diameter ~0.4 Earth
- Atmospheric Pressure = 1.5 × Earth
- 85% Nitrogen BUT
- Cold (~90 K)
- Reducing atmosphere
- Haze
- Lab for prebiotic chemistry

### The Cassini-Huygens Mission

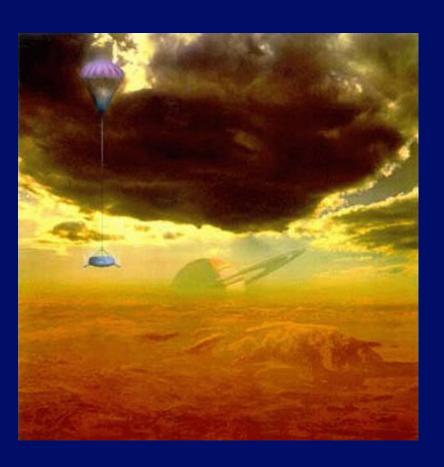


- Launched 10/13/97
- Arrived Saturn 7/2004
- Cassini studies
  - Saturn
  - Moons
- Huygens
  - Dropped onto Titan
  - Study atmosphere
  - Surface

http://saturn.jpl.nasa.gov/

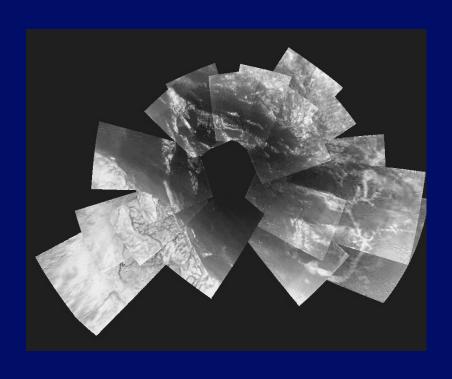


### Huygens Probe



- Released from Cassini
- Slowed by heat shield
- Parachute deployed
- Soft landing
- Sampled gases in atm.
- Results:
  - High winds
  - 430 km/hr at 120 km

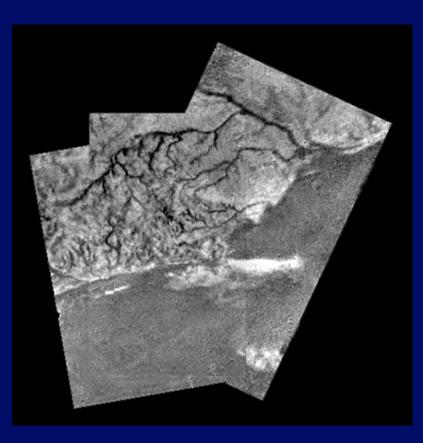
### Titan Surface 10km up



- Mosaic of images
- Taken during descent
- Clearly shows features

Photo: ESA

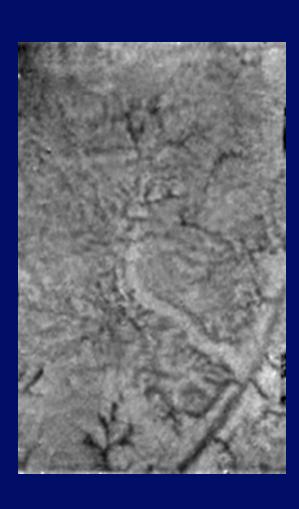
### Titan



- River channel
- Coastline
- Liquid is present
- Methane (CH<sub>4</sub>)

Photo: ESA

### Water Rift and Methane Springs?



- Straight feature:
- Water ice extruded?
- Stubby channels:
- Methane springs?

#### Lakes at northern latitudes

- Radar mapping of northern latitudes (2006)
- Strong evidence for liquid lakes
- And big cloud of ethane (C<sub>2</sub>H<sub>6</sub>)
- Ethane raining (or snowing) into lakes

### Lakes and Islands

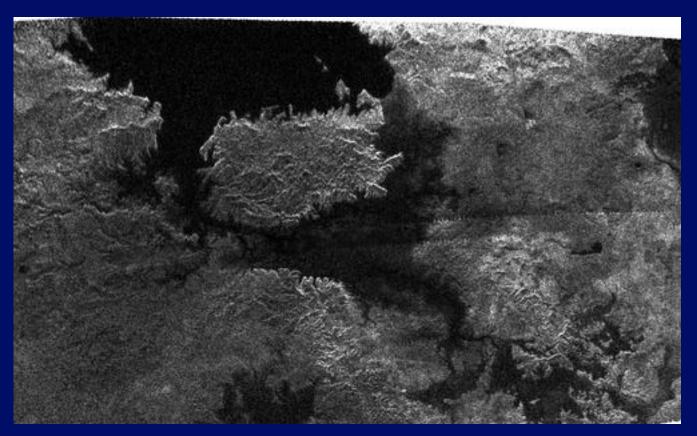


Image from Feb. 2007: based on radar.

Large lake and island (size of Big Island, Hawaii)

And smaller lakes

#### From the surface of Titan



- First view of surface
- "Rocks" of water ice
  - Pebble size (15 cm)
- Surface yielding
- Mixture of ices
  - Water
  - hydrocarbons

#### More Titan Results

- Hints of ammonia (NH<sub>3</sub>)/water (H<sub>2</sub>O) ocean
  - About 200 km under surface
  - Outgassing of NH<sub>3</sub> may supply N<sub>2</sub> atm.
- Mapping by radar reveals many lakes and seas of hydrocarbons
  - Seasonal changes in size, depth of a lake
  - Total hydrocarbons on surface about 100 times total oil and gas reserves on Earth

#### Possible Site for Life

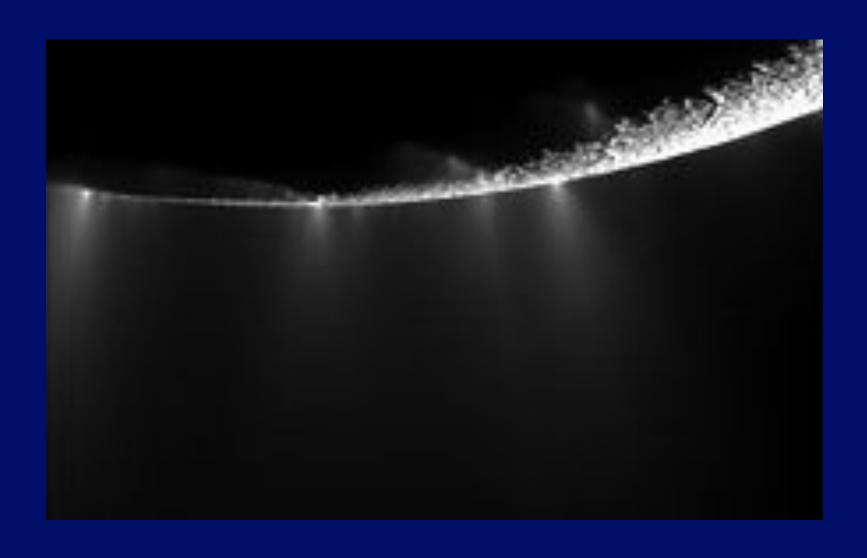
- Miller-Urey type experiments with Titan atm:
  - Formed amino acids and nucleotide bases
- Methane-based life?
- Metabolize with H<sub>2</sub> and C<sub>2</sub>H<sub>2</sub>, produce CH<sub>4</sub>
- Parallel to O<sub>2</sub> and glucose, produce CO<sub>2</sub>
- Also, could produce atmospheric nitrogen

#### Enceladus



- Moon of Saturn
- Very shiny
- Part of surface old (craters)
- Part is new, with cracks
- Cassini saw ice geysers (2006)
- Subsurface liquid water
- Source of heat unclear

## Geysers on Enceladus



#### How to search for life

Have to decide what test indicates life
Hard to anticipate conditions (recall Viking results)
What about finding "protolife"?

#### National Academy report - how to search for life

- 1. Delivery by comets, meteorites e.g. Mars meteorites
- 2. Sample return Mars possible
- 3. Experiments by landers -

Viking on Mars, ...

Future: Europa probe and return?

Titan?

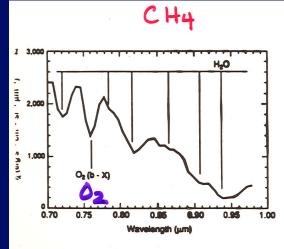
Issues of contamination

#### 4. Biomarkers

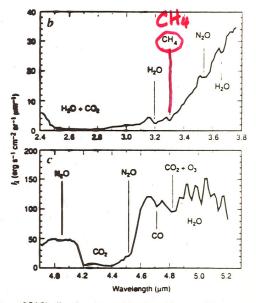
Presence of both O<sub>2</sub> and CH<sub>4</sub> in Earth atmosphere indicative of life How convincing?

### Detecting Life on Earth from Space

Galileo used during close Earth approach
Photographs (1 km resolution) No clear signs of intelligent life
Spectrometers - evidence of life Lots of O<sub>2</sub>



1 a, Galileo long-wavelength-visible and near-infrared spectra of larth over a relatively cloud-free region of the Pacific Ocean, north meo. The incidence and emission angles are 77° and 57° respective. The (b' $\sum_{k}^{+}$ — $X^{3}\sum_{k}^{-}$ ) 0–0 band of  $O_{0}$  at 0.76  $\mu$ m is evident, along a number of  $H_{2}O$  features. Using several cloud-free regions of gairmass, we estimate an  $O_{2}$  vertical column density of 1.5 km/st  $\pm$ 25%. b and c, Infrared spectra of the Earth in the 2.4–5.2  $\mu$ m n. The strong  $v_{3}$  CO $_{2}$  band is seen at the 4.3  $\mu$ m, and water vapour s are found, but not indicated, in the 3.0  $\mu$ m region. The  $v_{3}$  band rous oxide,  $N_{2}O$ , is apparent at the edge of the CO $_{2}$  band near m, and  $N_{2}O$  combination bands are also seen near 4.0  $\mu$ m. The

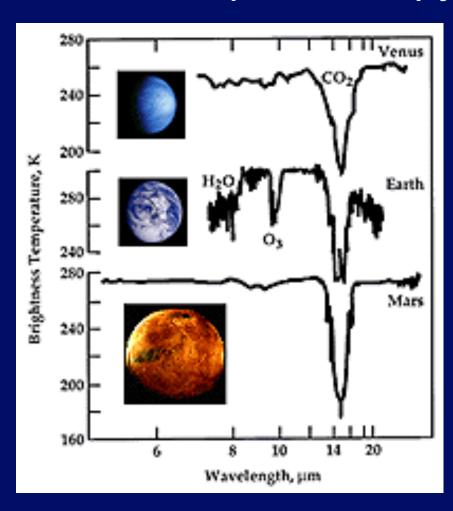


methane (0010) vibrational transition is evident at 3.31  $\mu$ m. A cr estimate <sup>10</sup> of the CH<sub>4</sub> and N<sub>2</sub>O column abundances is, for both spec of the order of 1 cm-amagate ( $\equiv$ 1 cm path at STP).

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RADIO EMISSION: CLEAR EVIDENCE OF

### Spectroscopy of atmosphere



Could be detected with future large space telescope, but very difficult
Need specialized capabilities