# Average Lifetime of Technological Civilization



# Average Lifetime of Technological Civilization

L = ?

- End of Communication Efforts (Civilization Survives)
   (Decades?)
- Civilization Evolves away from interest or capability (Post-technological Civilization)
   (Centuries - Millenia)

Civilization Collapses

 (Reversion to Pre-technological Culture)
 Exhaustion of resources
 Population explosion
 (~ 100 yrs - 1000 yrs)

Sudden, Catastrophic End of Civilization or Extinction of our Species Nuclear War leads to Nuclear Winter (10's - 100's of years) Natural Catastrophes (> 10<sup>5</sup> yr for most)

# **Resource Depletion**

Metals, Drinkable Water, Arable Land, ...

Energy is most fundamental

Energy is conserved

"Depletion" = conversion to less usable forms (entropy increases)

# **Resource Depletion**

Fossil Fuels (Stored Solar Energy) will eventually run out

~ 500 years for coal 200?

Nuclear Power?

Stopgap...

Ultimately Solar Power

Little Attempt to Plan Ahead

# World Energy Usage

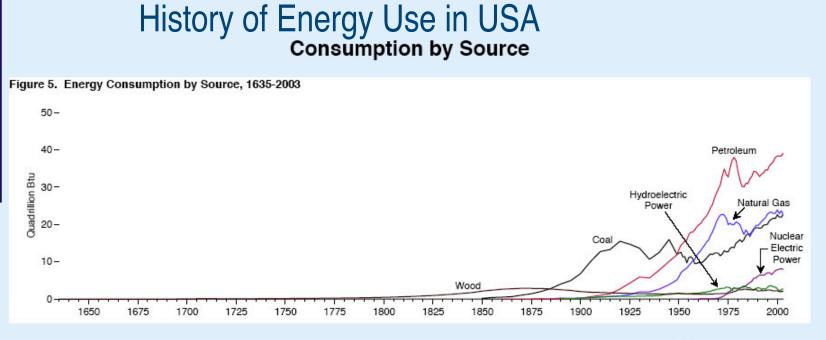
World 380 "Quads" Per year 1 Quad =  $10^{15}$  BTU  $\simeq 3 \times 10^{11}$  kw - H  $\simeq 10^{18}$  Joules (one exajoule)

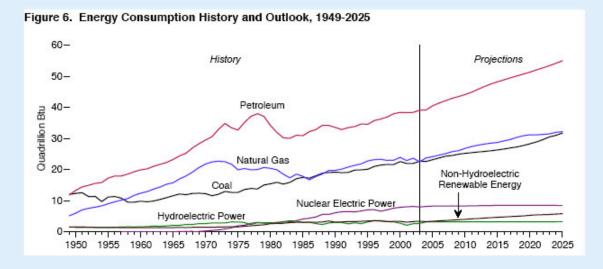
Average power is  $13 \times 10^6$  MW

U.S. uses 26% of this

Energy per capita ~ 6 metric tonnes of oil equivalent ~  $2 \times$  Europe ~  $5 \times$  World avg.

# Energy Perspectives





In the long view of American history, wood served as the preeminent form of energy for about half of the Nation's history. Around 1885, coal surpassed wood's usage. Despite its tremendous and rapid expansion, coal was, in turn, overtaken by petroleum in the middle of the 20th century. Natural gas, too, experienced rapid development into the second half of the 20th century, and coal began to expand again. Late in the 20th century still another form of energy, nuclear electric power, was developed and made significant contributions.

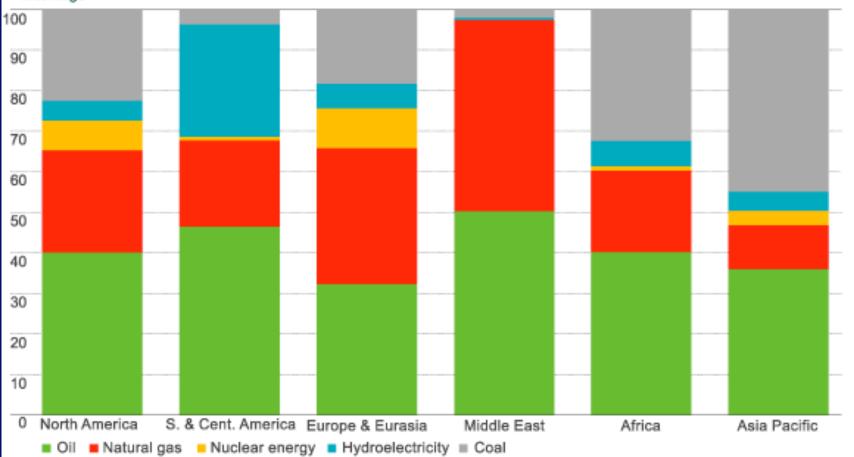
While the Nation's energy history is one of large-scale change as new forms of energy were developed, the outlook for the next couple of decades (assuming current laws, regulations, and policies) is for continued growth and reliance on the three major fossil fuels—petroleum, natural gas, and coal—modest expansion in renewable resources, and relatively flat generation from nuclear electric power.

Energy Information Administration / Annual Energy Review 2003

# Regional Primary Energy Consumption Pattern

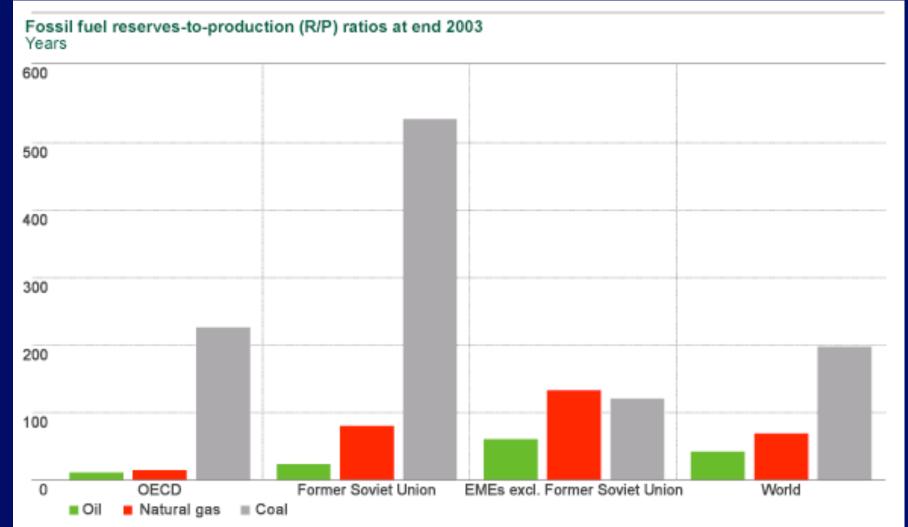
### Regional primary energy consumption pattern 2003

Percentage



Oil remains the largest single source of energy in most parts of the world. The exceptions are the Former Soviet Union, where gas dominates and Asia Pacific where coal is the dominant fuel.

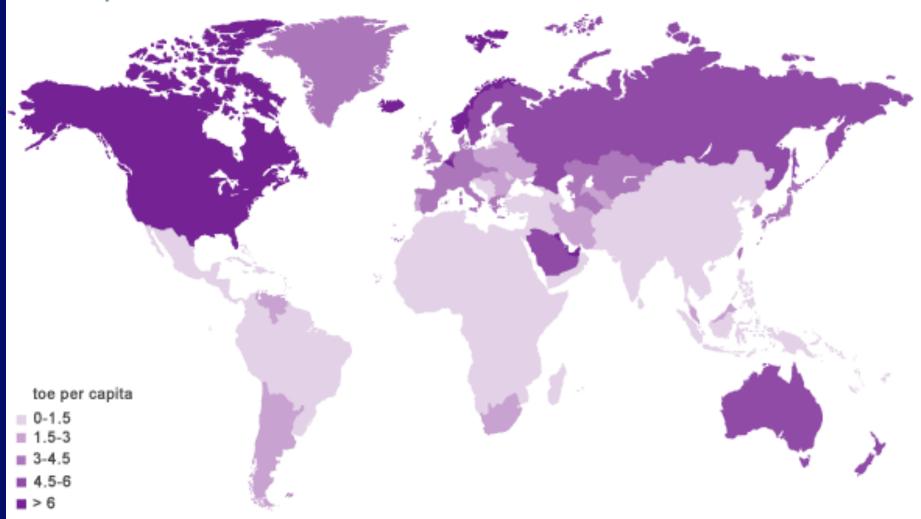
# Fossil Fuel R/P ratios



The world's reserves-to-production ratio for coal is around five times that for oil and more than three times that for natural gas. Coal's dominance in reserves-to-production ratio terms is particularly pronounced in the OECD and the Former Soviet Union.

# Energy Consumption per capita

Primary energy consumption per capita Tonnes oil equivalent



# Side Effects

- General Pollution of Air, Water, Land Makes resources less usable Air pollution, respiratory problems Undrinkable water
   Desertification of farm-lands
- Ozone Layer Destruction

   ⇒ UV reaches surface
   Skin Cancer, Cataracts, ...
   Crop Damage

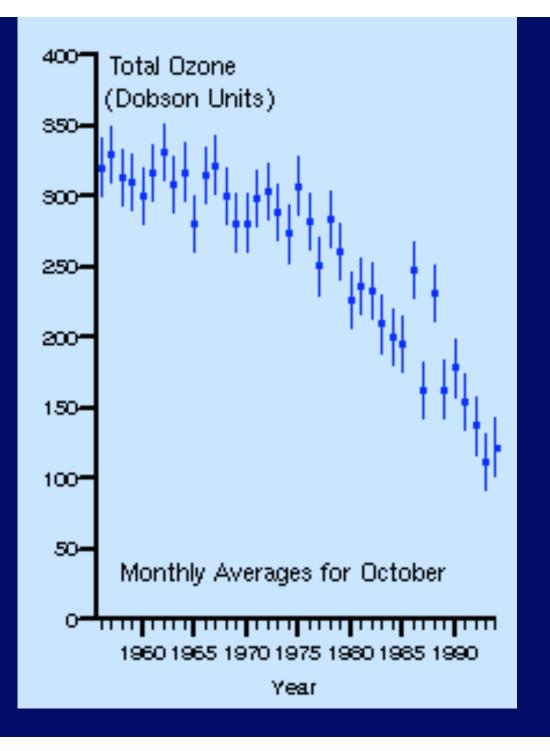
Caused by CFC's (refrigeration, styrofoam,...) other chemicals

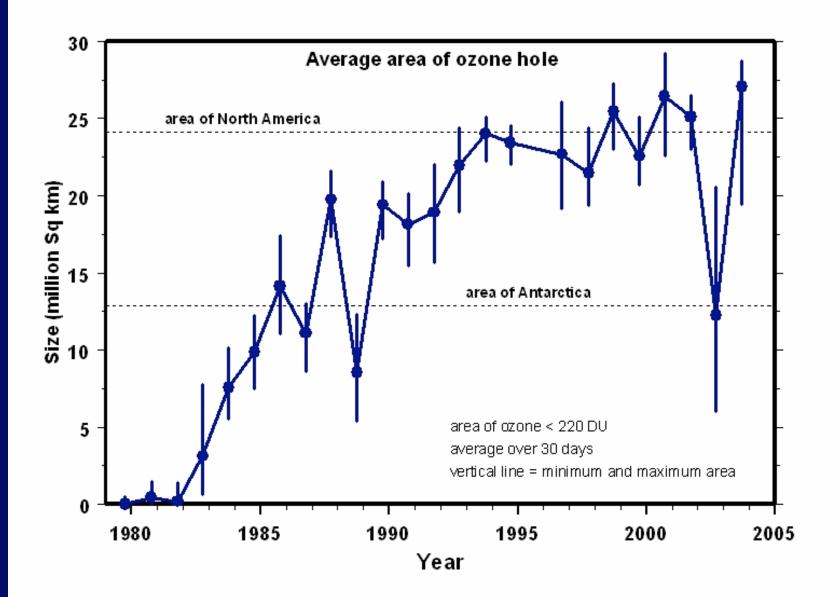
Stratosphere is very sensitive and poorly understood

Catalytic reactions: One CFC molecule leads to the destruction of **many** ozone molecules

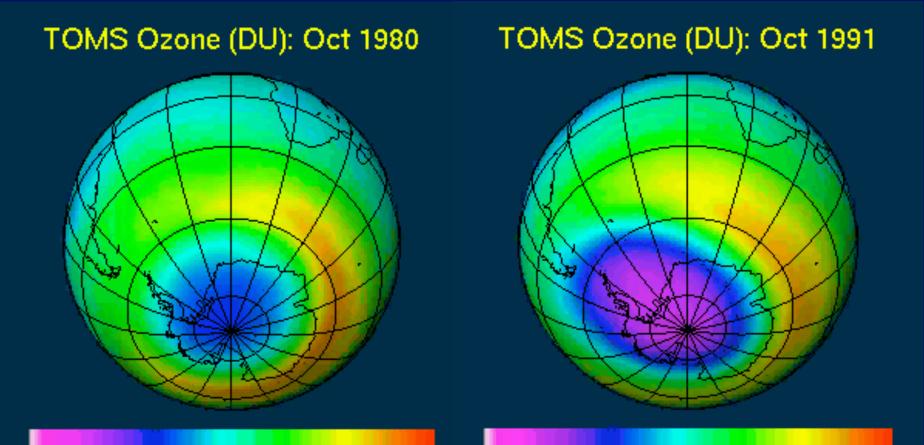
 $O_3 \longrightarrow O_2$ 

### Ozone over South Pole

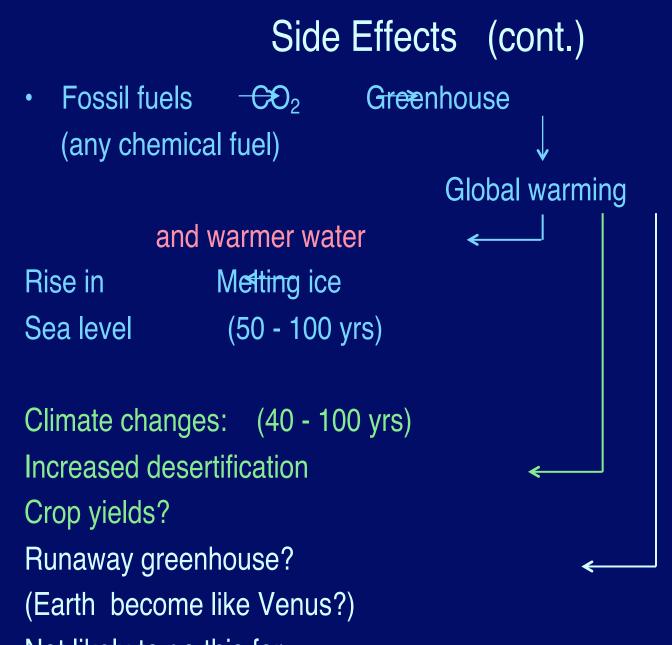




# Growth of ozone hole

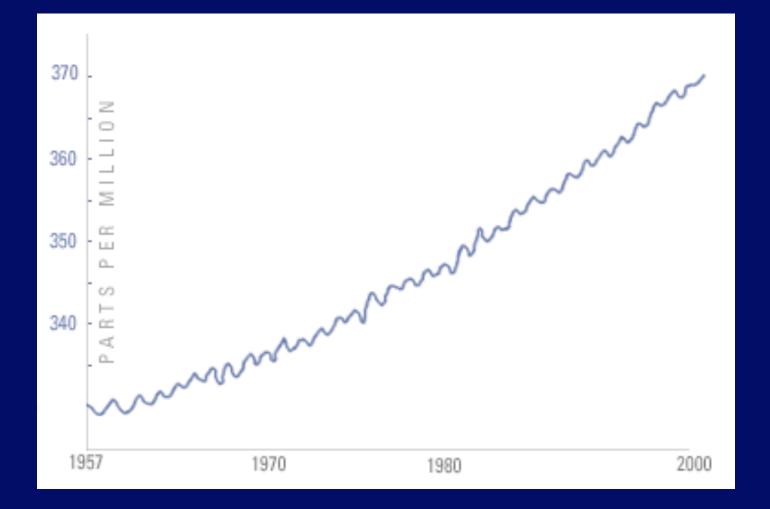


### 100 140 180 220 260 300 340 380 420 460 500 100 140 180 220 260 300 340 380 420 460 500

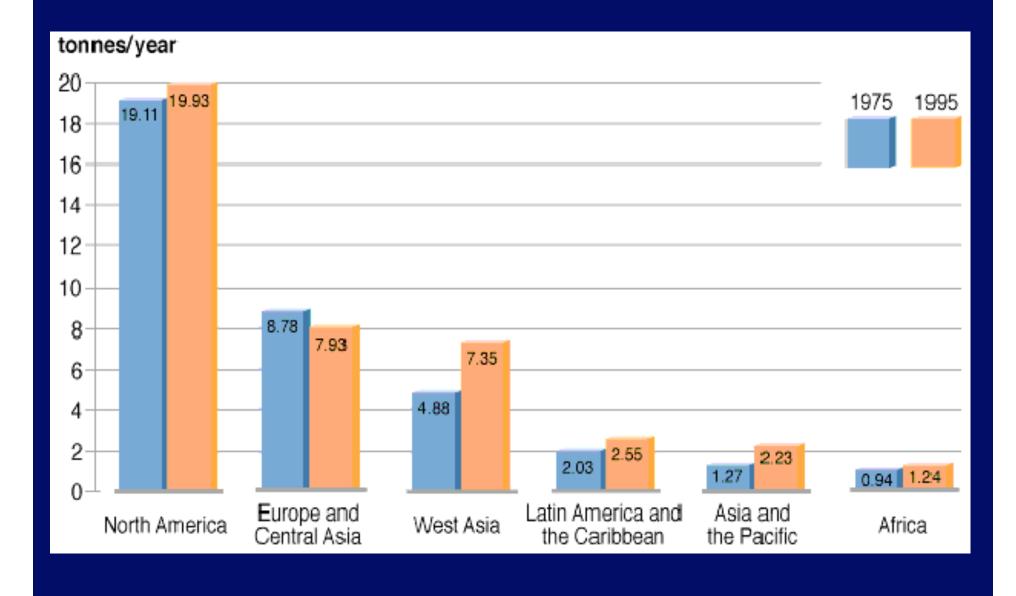


Not likely to go this far

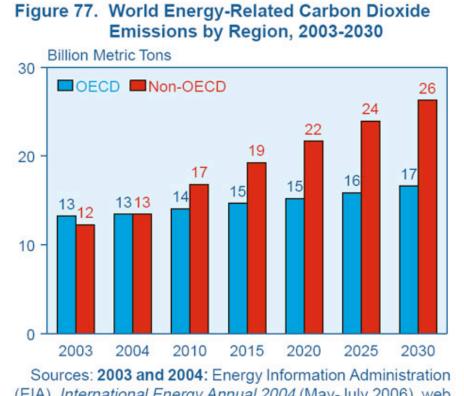
# Carbon Dioxide Increase



# CO<sub>2</sub> Production



# Update on production balance

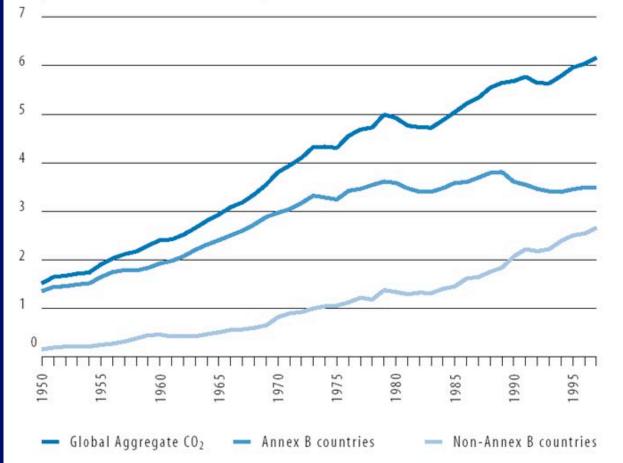


(EIA), *International Energy Annual 2004* (May-July 2006), web site www.eia.doe.gov/iea. **Projections:** EIA, System for the Analysis of Global Energy Markets (2007).

# OECD: Europe, US, Canada, Australia, Japan, Mexico Not Russia, India, China

### FIGURE 5: GLOBAL CO<sub>2</sub> EMISSIONS, 1950-1997

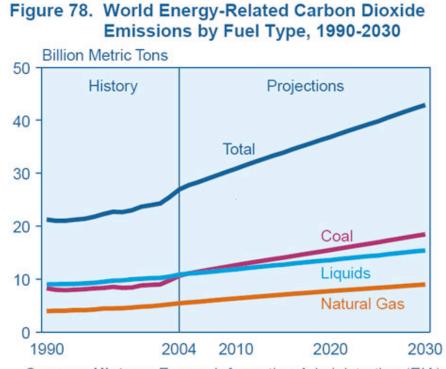
(billions of metric tons of carbon)



Source: Meyerson, F. A. B. 2001. "Population and Climate Change Policy." In: Climate Change Policy: A Survey, edited by S. Schneider, A. Rosencranz, and J. Niles. (Forthcoming.) Washington, D.C.: Island Press.

# In Billion Metric tonnes of CARBON (12/44 of CO<sub>2</sub>)

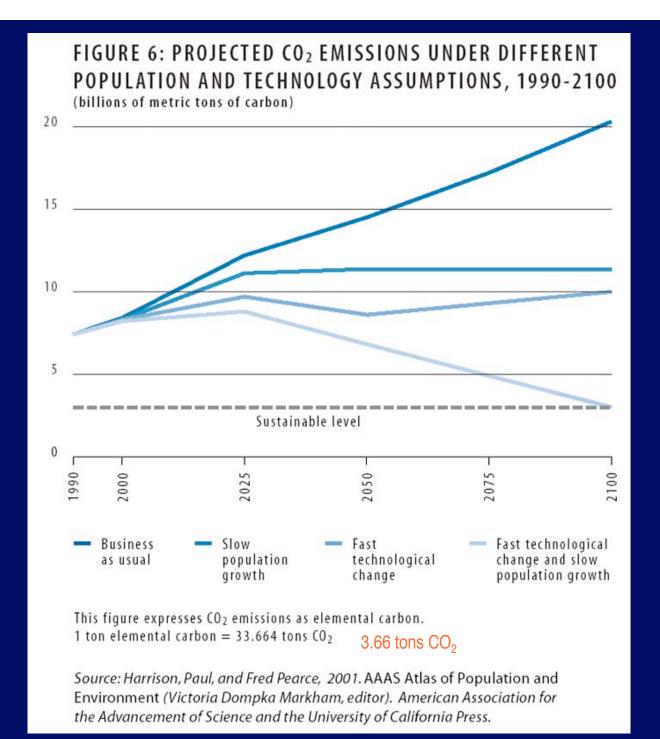
# CO<sub>2</sub> Production Continues to Increase



Sources: **History:** Energy Information Administration (EIA), International Energy Annual 2004 (May-July 2006), web site www.eia.doe.gov/iea. **Projections:** EIA, System for the Analysis of Global Energy Markets (2007).

# Note: this is for $CO_2$ rather than C. So 44/12 times next plot

# From Energy Information Administration 2008



# Other factors

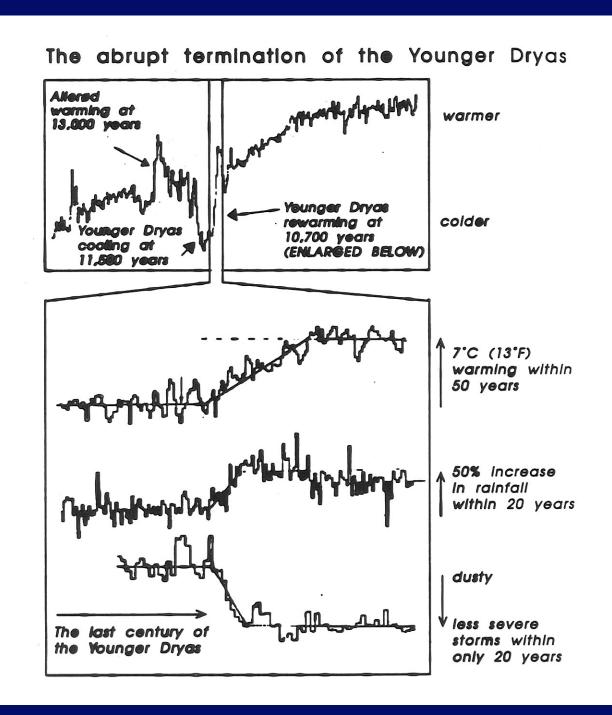
- New models include Sulfate emission leads to haze which leads to increase in albedo
- Cooling tends to balance warming from Greenhouse CO<sub>2</sub>
   Less temperature rise in short term

Ice core analysis shows strong correlation of temperature and astronomical cycles rotation axis, orbital variations, solar cycle

Also - we are still in last stages of "little ice age" In climate behavior, but not temperature Greenland ice cores *Nature*, 15 July 1993 Study temperature, climate... over 150,000 yr Last interglacial (Eemian) 115,000 - 130,000 yr ago

warmer 3 temp. states: like present colder Very rapid switches (up to 10° C)

Our current stable climate may not be typical of interglacials



# **Population Explosion**

(The revenge of Malthus?)

Agriculture - Population Growth - Disease Population Growth leads to more rapid depletion of resources More pollution More conflict?

Two "events" (transitions)10,000 yrs agoAgriculture250 yrs agoDisease lessened(demographic transition)

Time	Total Pop.	Growth Rate (per thousand per year)
Before Agriculture	~ 8 × 10 <sup>6</sup> (??)	0.015
~ 8000 BCE - 1 CE	~ 3 × 10 <sup>8</sup>	0.36
1 CE - 1750 CE	$\sim 8 \times 10^8$	0.56
1750-1800	~ 1 × 10 <sup>9</sup>	4.4
÷		
1950 - 1975	4 × 10 <sup>9</sup>	17.1
2000	6 × 10 <sup>9</sup>	~ 18

# Population Doubling in 55 years

# **Population Mathematics**

Rate of increase  $\propto$  Number  $\times$  (Birth - Death) leads to exponential growth if (Birth - Death) constant Pop (t) = Pop (Now)  $2^{(t/t_d)}$  $t_d$  = doubling time  $\sim$  55 years So doubles in 55 yrs Quadruples  $(2^2)$  in 110 yrs, ... 990 yr (18  $t_d$ ) Pop =  $1.3 \times 10^{15}$  $\sim$  fills land area 2530 yr (46  $t_d$ ) Mass >  $M_{(earth)}$  ! 12, 375 yr (225  $t_d$ ) Mass expands at c !! Current population growth is NOT sustainable

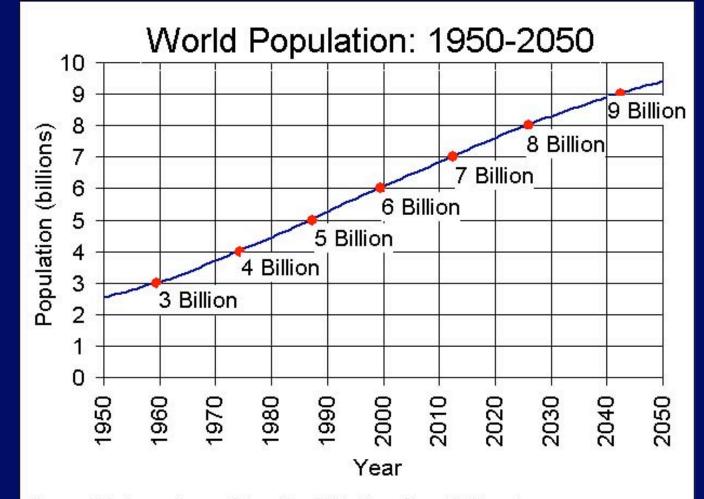
# **World Vital Events Per Time Unit: 2008**

(Figures may not add to totals due to rounding)

•	Natural						
•	Time uni	Births Deaths		increase			
•							
•	Year	133,398,951	55,503,922	2 77,895,029			
•	Month	11,116,579	4,625,327	6,491,252			
•	Day	364,478	151,650	212,828			
•	Hour	15,187	6,319	8,868			
•	Minute	253	105	148			
•	Second	4.2	1.8	2.5			

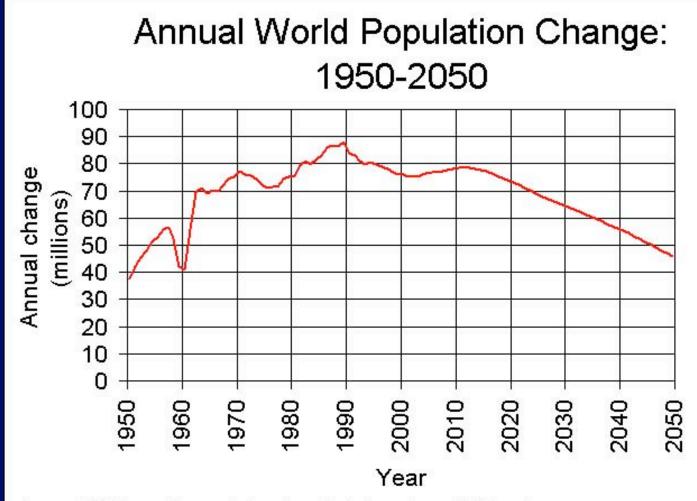
http://www.census.gov/main/www/popclock.html

# **Projected World Population Growth**

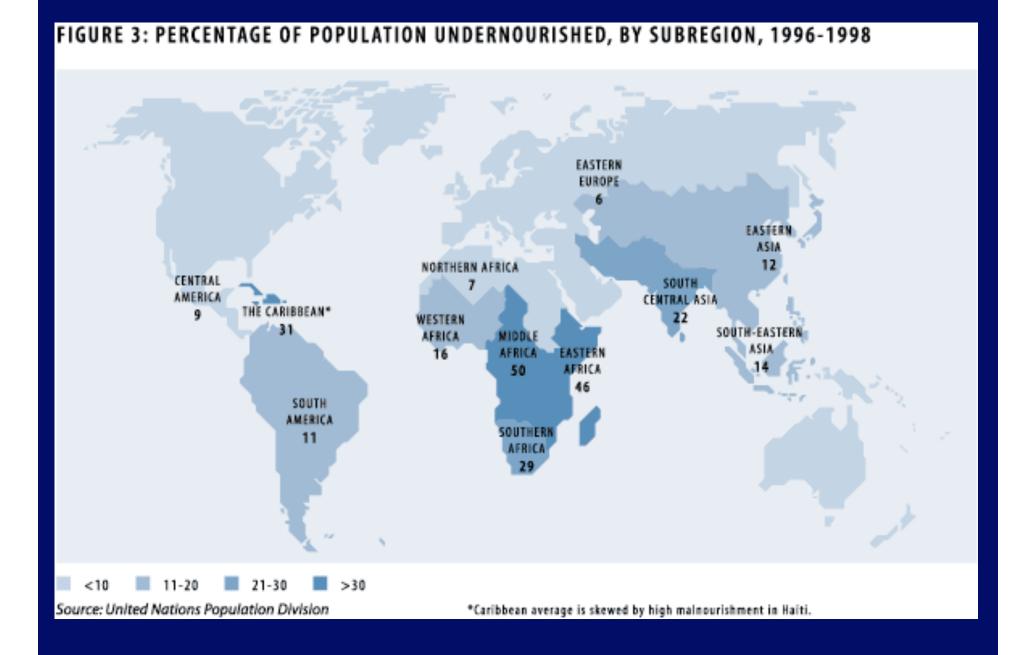


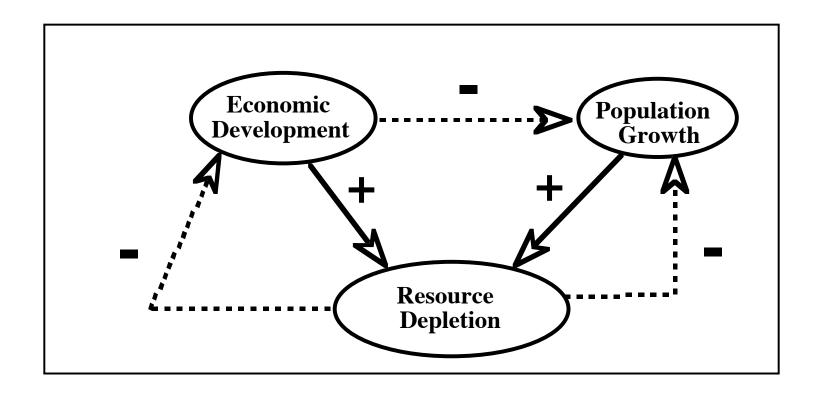
Source: U.S. Census Bureau, International Data Base, August 2006 version.

# Changes in Population









Does negative effect on population growth Beat positive effect on resource depletion?

Can we get to sustainable economy before We exhaust resources?

# Nuclear War

Total arsenal world-wide ~ 10,000 megatons

# Global effects of all-out war

- Depletion of ozone
- Radioactive fallout
- Dust and smoke in atmosphere would block sunlight and lead to cooling of the Earth "Nuclear Winter"

# The World's Nuclear Arsenals

			Country	Suspected Strategic Nuclear Weapons	Suspected Non- Strategic Nuclear Weapons	Suspected Total Nuclear Weapons
*	**		China	250	120	400
			France	350	0	350
	۲		India	60	?	60+
	✡		Israel	100-200	?	200+
	Ć	*	Pakistan	24-48	?	24-48

# The World's Nuclear Arsenals

Country	Suspected Strategic Nuclear Weapons	Suspected Non- Strategic Nuclear	Suspected Total Nuclear Weapons
Russia	~ 6,000	Weapons ~ 4,000	~ 10,000
United Kingdom	180	5	185
United States	8,646	2,010	10,656

http://www.cdi.org/issues/nukef&f/database/nukearsenals.cfm

### Natural Catastrophes

Collisions Stars? Negligible

Molecular Clouds? t ~  $10^8$  yr Likely, but the effects are unclear

Less dense clouds? More common but effects are probably less

Asteroids and other debris (comets, meteoroids, ...)

#### Effect of Asteroid Impact:

e.g. 1/4 km radius

 $V = 30 \text{ km s}^{-1}$  (65,000 miles/hour)

 $E_k = 1/2 \text{ Mv}^2 \simeq 7200 \text{ megatons of TNT}$  $\simeq \text{ all-out nuclear war}$ 

Crater ~ 10 km across, few km deep 10<sup>12</sup> tons of debris released into atmosphere If covers globe, leads to temperature drop and "asteroid winter" How Often do Large Asteroids Strike the Earth?

1937 Hermes	~ 500,000 miles
1989FC	Similar
1993BA	170,000 km (5 - 10 m diameter)

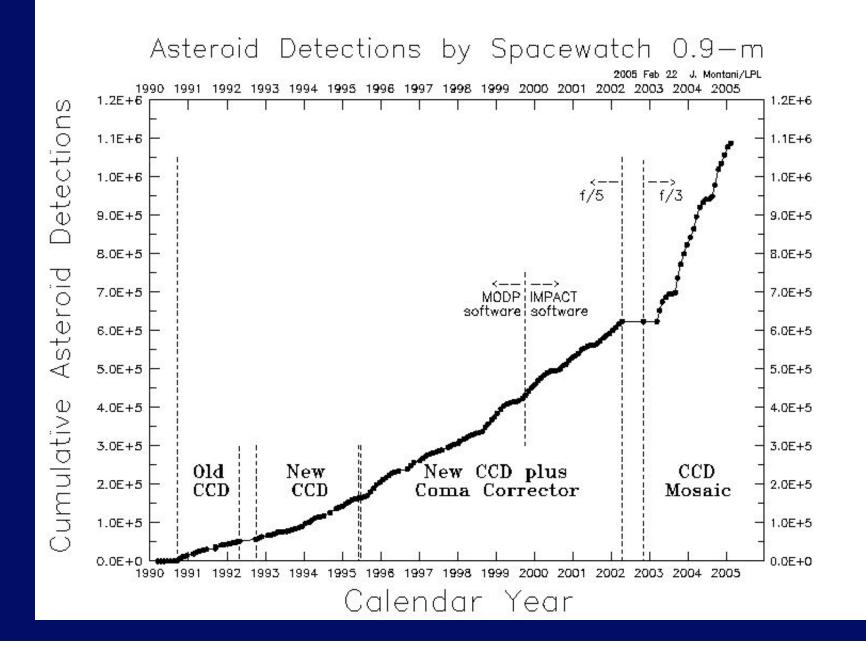
How often might we expect global catastrophe? "Substantial" Impacts (1 km or larger)  $t \sim 10^5$  yr -  $10^6$  yr Major Extinctions  $t \sim 30 \times 10^6$  yr Mass Extinctionst  $\sim 100 \times 10^6$  yr ? More massive asteroids more destructive, but also more rare, so collisions are less likely

Preventable by advanced civilization?

1991 BA ~ 40 kilotons TNT (3 × Hiroshima)
50 meter objects - once per century

April 1992 - proposal for project to search and identify - space watch underway

#### **Spacewatch Detections**

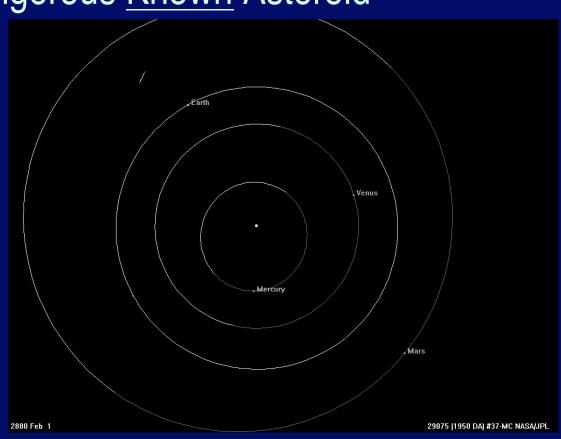


# Most Dangerous Known Asteroid

1950 DA

Radar used to map orbit ~ 1 km in diameter

Close approach in Yr 2880



Probability of collision ~ 0.33%

 $V \sim 14 \text{ km s}^{-1}$   $E \sim 10^5 \text{ Megatons}$ 

Exact orbit depends on small effects - tugs from Earth, Mars, light absorption + radiation, ...

# Another "Interesting" Asteroid

- Apophis (2004 MN4)
- d = 0.25 km, would release 400 Megatons
- 1/45000 chance of collision in 2036
- <u>http://neo.jpl.nasa.gov/risk/</u>
- Has data base of Near Earth Objects

### Solar variations

~ 10<sup>5</sup> yr

 Short term - cyclic variations in L, orbit of Earth -----> ice ages, climate change

~  $1-2 \times 10^9$  yr 2. Sun increases in L on main sequence -----> loss of oceans UV + H<sub>2</sub>O = 2H + O H lost to space ~  $5 \times 10^9$  yr

Off main sequence leads to Red Giant
 -----> atmosphere evaporates

Could advanced civilization delay loss of oceans?
(Decrease greenhouse, add dust)
Move to Mars? Mars will be in HZ by end of Sun's main sequence lifetime.
Red giants lose mass in winds: Earth's orbit moves out to 1.5 AU by 7 x 10<sup>9</sup> yr; but HZ is now 50-80 AU!
Sun's atmosphere engulfs Earth and it spirals in.

Other stars?

Nearby star leads to Supernova If within 30 *cy*, ozone is destroyed

 $\sim 2 \times 10^9 \text{ yr}$ 

# **Ultimate Limits**

If Universe Closed, recollapses

~ 10<sup>12</sup> Big Crunch (unlikely)



Very unlikely because evidence now indicates that expansion is accelerating (dark energy)

But, since we don't understand dark energy, it could reverse.

### If open, expands forever

About 5 x 10<sup>9</sup> years, Andromeda collides with MW

- 10<sup>11</sup> local galaxies collapse into a supergalaxy, if acceleration continues, all other galaxies have disappeared
- $10^{12} 10^{14}$  all stars die
- 10<sup>17</sup> planetary systems disrupted
- 10<sup>18</sup> 10<sup>20</sup> galaxies "evaporate"
- 10<sup>32</sup> 10<sup>34</sup> protons decay?
- 10<sup>100</sup> Black holes evaporate

#### What to choose for L?

•For number of civilizations now,  $L \le 5 \times 10^9$  yrs -[ age of galaxy – time to evolve]

Important to choose L consistent with what you think is the most likely way civilizations end.

#### Darkness

I had a dream, which was not all a dream. The bright sun was extinguish'd, and the stars Did wander darkling in the eternal space, Rayless, and pathless, and the icy earth Swung blind and blackening in the moonless air;

- Lord Byron, 1816

