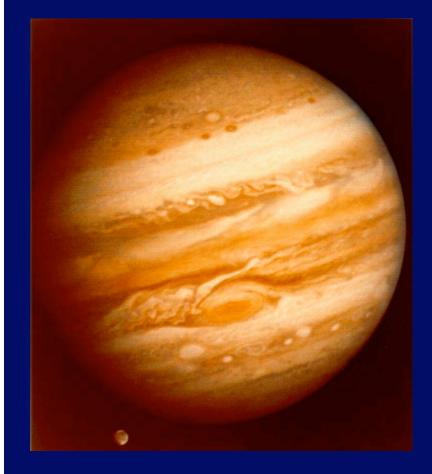
Life in the Outer Solar System

Jupiter



Big $R = 11R_{\oplus}$ Massive $M = 300 M_{\oplus}$ = 2.5 all the rest

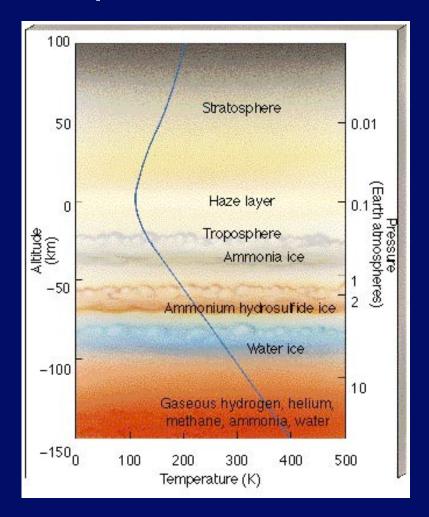
Mostly H₂, He Thick Atmosphere But also more complex molecules Colors, storms

Like Miller - Urey

Life in Jupiter Atmosphere?

Sagan-Salpeter, etc.

Sinkers Floaters Hunters (Plankton) (Fish) (Fish)



Galileo Results on Jupiter Reached Jupiter Dec. 1995 Sent probe into Jupiter's atmosphere at 100,000 mile/hour Decelerated at 230 g Lasted for 57 min.

Found: Strong winds Turbulence, little lightning _

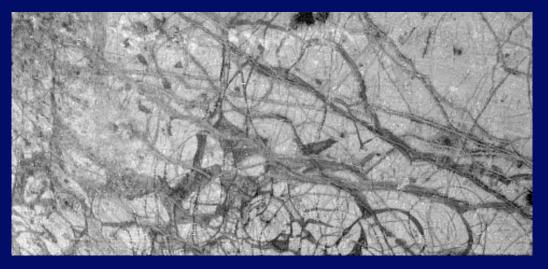
Life less likely?

Surprise: Little or no H₂O May have entered in an unusual place (fewer clouds)

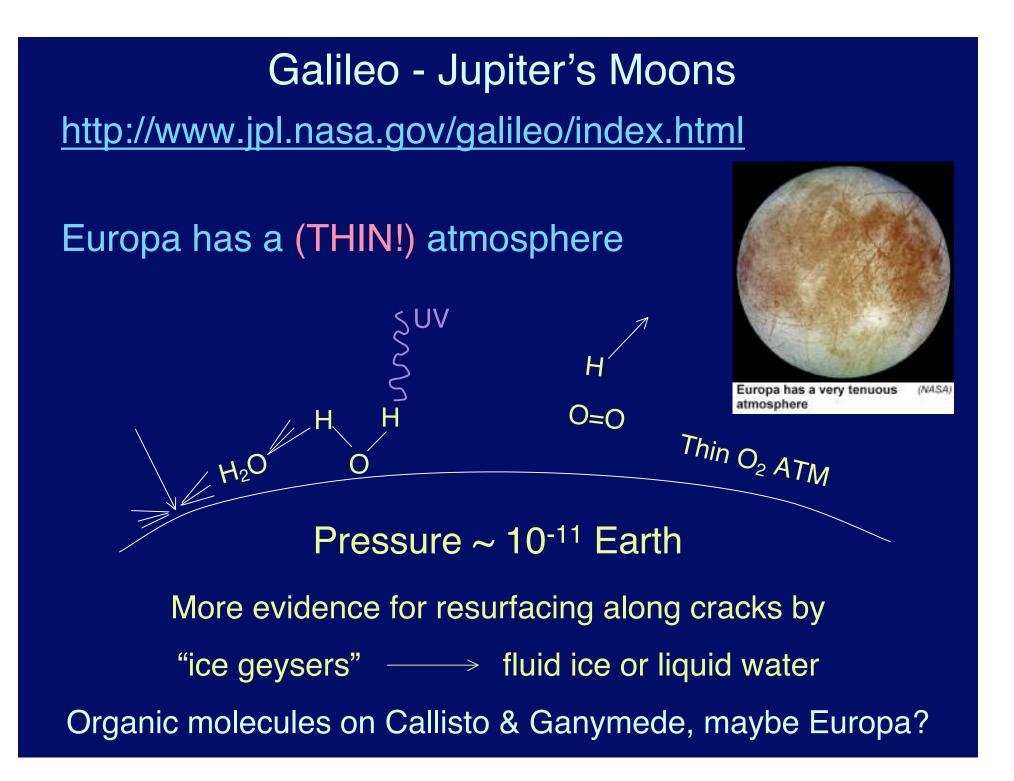




Europa (Moon of Jupiter) Surface: Fractured Ice Subsurface Oceans? (Heated from Inside)



Close-up of "ice floes"



Saturn



- Big (9.4 R⊕)
- Massive (95 M_⊕)
- Year 29.5 years
- Day 0.43 days
- Composition similar to Jupiter



Titan

- Moon of Saturn
- Diameter ~0.4 Earth
- Atmospheric Pressure = 1.5 × Earth
- 85% Nitrogen BUT
- Cold (~90 K)
- Reducing atmosphere
- Haze
- Lab for prebiotic chemistry

The Cassini-Huygens Mission



- Launched 10/13/97
- Arrived Saturn 7/2004
- Cassini studies
 - Saturn
 - Moons
- Huygens
 - Dropped onto Titan
 - Study atmosphere
 - Surface



2005 Saturn Tour Highlights:

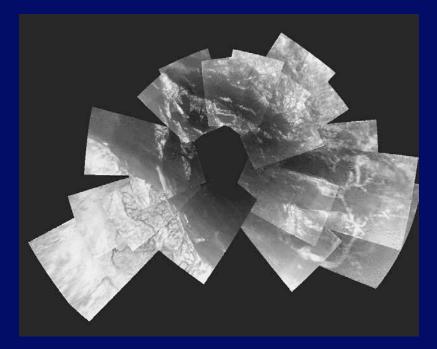
- Jan. 14, 2005: The European Space Agency's Huygens probe descends through Titan's cloudy atmosphere, touching down on the surface about two and half hours later. Cassini will send the data back to Earth.
- Feb. 15, 2005: Cassini makes another pass by Titan. In 2005, the spacecraft will have six chances to study Titan at altitudes ranging from 1,025 kilometers (637 miles) to 60,000 kilometers (37,290 miles).
- Mar. 9, 2005: Cassini flies within 500 kilometers (311 miles) of icy Enceladus. Cassini will visit Enceladus five times in 2005.
- Sep. 26, 2005: Cassini studies Hyperion at a range of 1,010 kilometers (628 miles), the closest approach ever to the tiny moon. It will be Cassini's only visit to the moon during the primary mission.
- Oct. 11, 2005: Cassini turns its instruments on Dione from a distance of 500 kilometers (311 miles).
- Nov. 26, 2005: Cassini passes within 500 kilometers (311 miles) of Rhea.

Huygens Probe



- Released from Cassini
- Slowed by heat shield
- Parachute deploys
- Goal of soft landing
- Sample gases in atm.
- Results so far:
 - High winds
 - 430 km/hr at 120 km

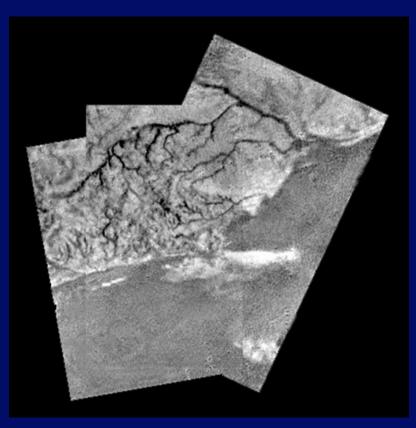
Titan Surface 10km up



- Mosaic of images
- Taken during descent
- Clearly shows features

Photo: ESA

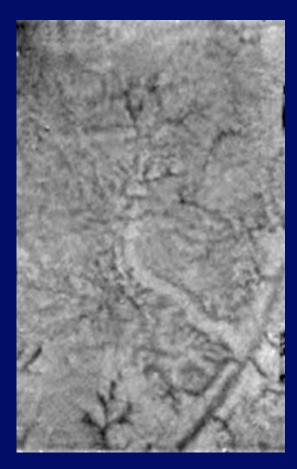
Titan



- River channel
- Coastline
- Liquid is present
- Methane (CH_4)

Photo: ESA

Water Rift and Methane Springs?



- Straight feature:
- Water ice extruded?
- Stubby channels:
- Methane springs?

From the surface of Titan



- First view of surface
- "Rocks" of water ice
 Pebble size (15 cm)
- Surface yielding
- Mixture of ices
 - Water
 - hydrocarbons

More Titan Results

- Hints of ammonia (NH₃)/water (H₂O) ocean
 Under surface
 Output of NUL means the Number of Null Means the Nu
 - Outgassing of NH_3 may supply N_2 atm.

How to search for life

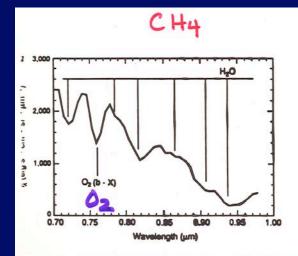
Have to decide what test indicates life Hard to anticipate conditions (recall Viking results) What about finding "protolife"?

National Academy report - how to search for life

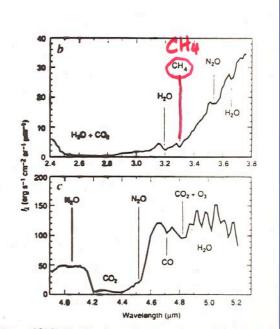
- 1. Delivery by comets, meteorites e.g. Mars meteorites
- 2. Sample return Mars possible
- 3. Experiments by landers -Viking on Mars, ...
 Future: Europa probe and return? Titan?
 Issues of contamination
- 4. Biomarkers

Presence of both O_2 and CH_4 in Earth atmosphere indicative of life How convincing?

Detecting Life on Earth from Space Galileo used during close Earth approach Photographs (1 km resolution) No clear signs of intelligent life Spectrometers - evidence of life Lots of O₂



1. a, Galileo long-wavelength-visible and near-infrared spectra of arth over a relatively cloud-free region of the Pacific Ocean, north meo. The incidence and emission angles are 77° and 57° respective. The (b' $\sum_{\alpha}^{*} - X_{a}^{*}$) 0-0 band of O₂ at 0.76 µm is evident, along a number of H₂O features. Using several cloud-free regions of the annuber of H₂O features. Using several cloud-free regions of the gairmass, we estimate an O₂ vertical column density of 1.5 km-(at ± 25%. b and c, infrared spectra of the Earth in the 2.4-5.2 µm n. The strong v₃ CO₂ band is seen at the 4.3 µm, and water vapour s are found, but not indicated, in the 3.0 µm region. The v₃ band rous oxide, N₂O, is apparent at the edge of the CO₂ band near m, and N₂O combination bands are also seen near 4.0 µm. The



methane (0010) vibrational transition is evident at $3.31 \,\mu$ m. A cr estimate¹⁰ of the CH₄ and N₂O column abundances is, for both spec of the order of 1 cm-amagate (=1 cm path at STP).

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RADIO EMISSION: CLEAR EVIDENCE OF INTELLIGENT LIFE

TPF Concepts

TPF-I Infrared Interferometer (2020)



TPF-C Visible light coronagraph (2014)

Spectroscopy of atmosphere

