

Cosmic Evolution

Part 1: Protons to heavy elements

Big Bang occurred 13.7 Billion yrs ago (13.7×10^9 yr)

Only fundamental particles existed for first few minutes

<u>Name</u>	<u>Symbol</u>	<u>Charge</u>	<u>Mass</u>
Proton	p	+	1.7×10^{-24} g
Neutron	n	0	1.7×10^{-24} g
Electron	e	-	1×10^{-27} g
Photon	γ	0	0
Neutrino	ν	0	~ 0 (?)

Building blocks of nuclei but only one kind of nucleus

Proton = nucleus of Hydrogen

Energy of Motion (Kinetic Energy)

$$E = \frac{1}{2} mv^2 \quad (\text{if } v \text{ not close to } c)$$

Gas at Temperature T ,

Avg. Energy $E = \frac{3}{2} kT$

So avg. v : $\frac{1}{2} mv^2 = \frac{3}{2} kT$

$$v = \sqrt{\frac{3kT}{m}} \equiv \left(\frac{3kT}{m} \right)^{\frac{1}{2}}$$

Higher $T \rightarrow$ Higher v , E on avg.

Early Universe so hot that collisions broke apart any complex things that might have formed

As Universe expanded, T dropped

at ~ 3 min, $T \sim 10^9$ K

A few nuclei form (nucleosynthesis)

at ~ 30 min, $T \sim 3 \times 10^8$ K end of nucleosynthesis

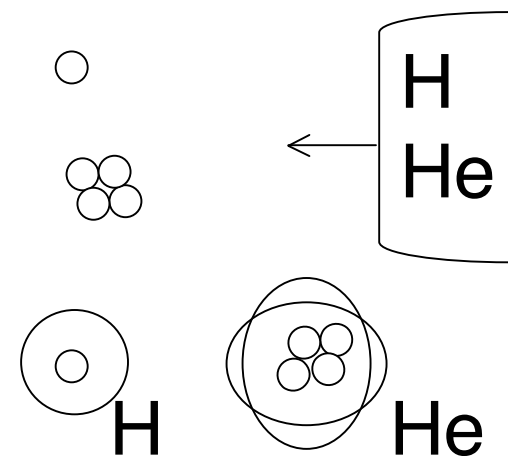
Composition of Universe at 30 min.

~ 94 % proton

~ 6 % alpha part.
(and electrons)

At $3 - 7 \times 10^5$ years $T \sim 3000$

Nuclei + electron \rightarrow Atoms



First Generation Stars

Expanding Universe

But, Gravity \rightarrow Galaxies ($\sim 10^{11}$ to $10^{12} M_{\odot}$)

\rightarrow Stars (0.1 to $100 M_{\odot}$)

Oldest stars in disk $\sim 10^{10}$ y old

First generation stars \rightarrow No C, O, N, ...

\Rightarrow No life No Si, Fe \Rightarrow No Earthlike planets

But they made “heavy” elements

So later stars could have solid planets, life

Gravitational Potential Energy

For example: Reservoir of water behind dam



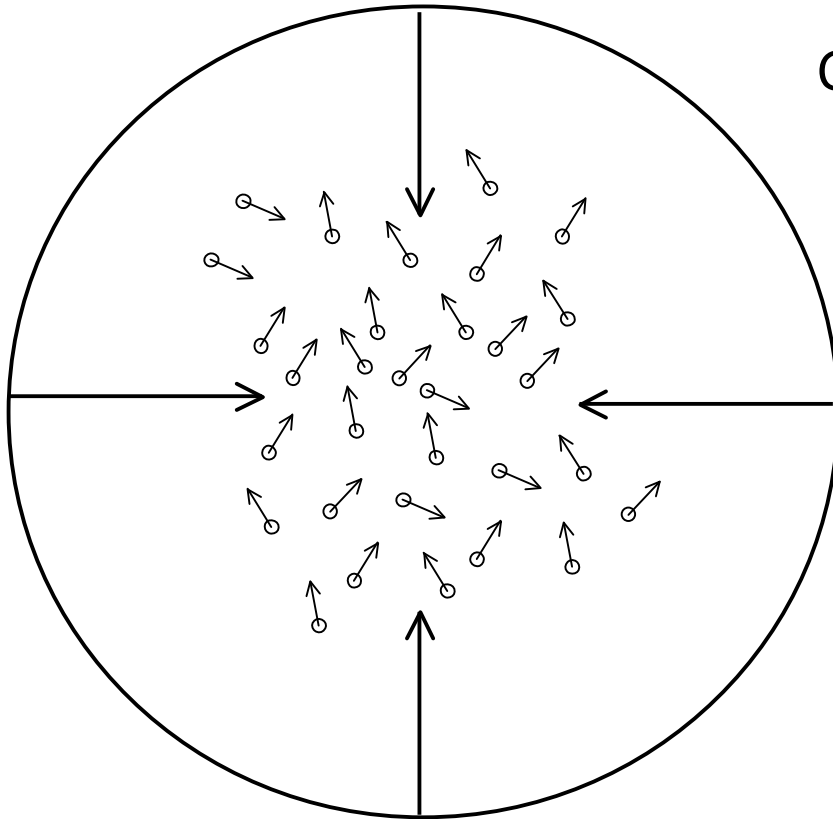
Higher
GPE

Energy released
as water falls

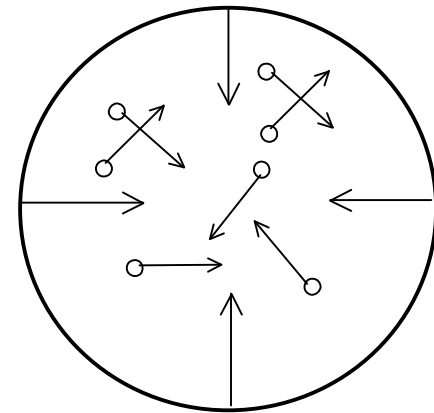
Lower GPE

Consider a clump of Gas collapsing to form a star

Apply to collapsing gas



GPE \rightarrow Heat



Atoms move faster
Temperature is higher

Temperature in core reaches 10^7 k

Nuclear reactions begin

Collapse stops

Why?

Nuclear Potential Energy

Four basic forces: gravity, electromagnetic, weak and strong nuclear force

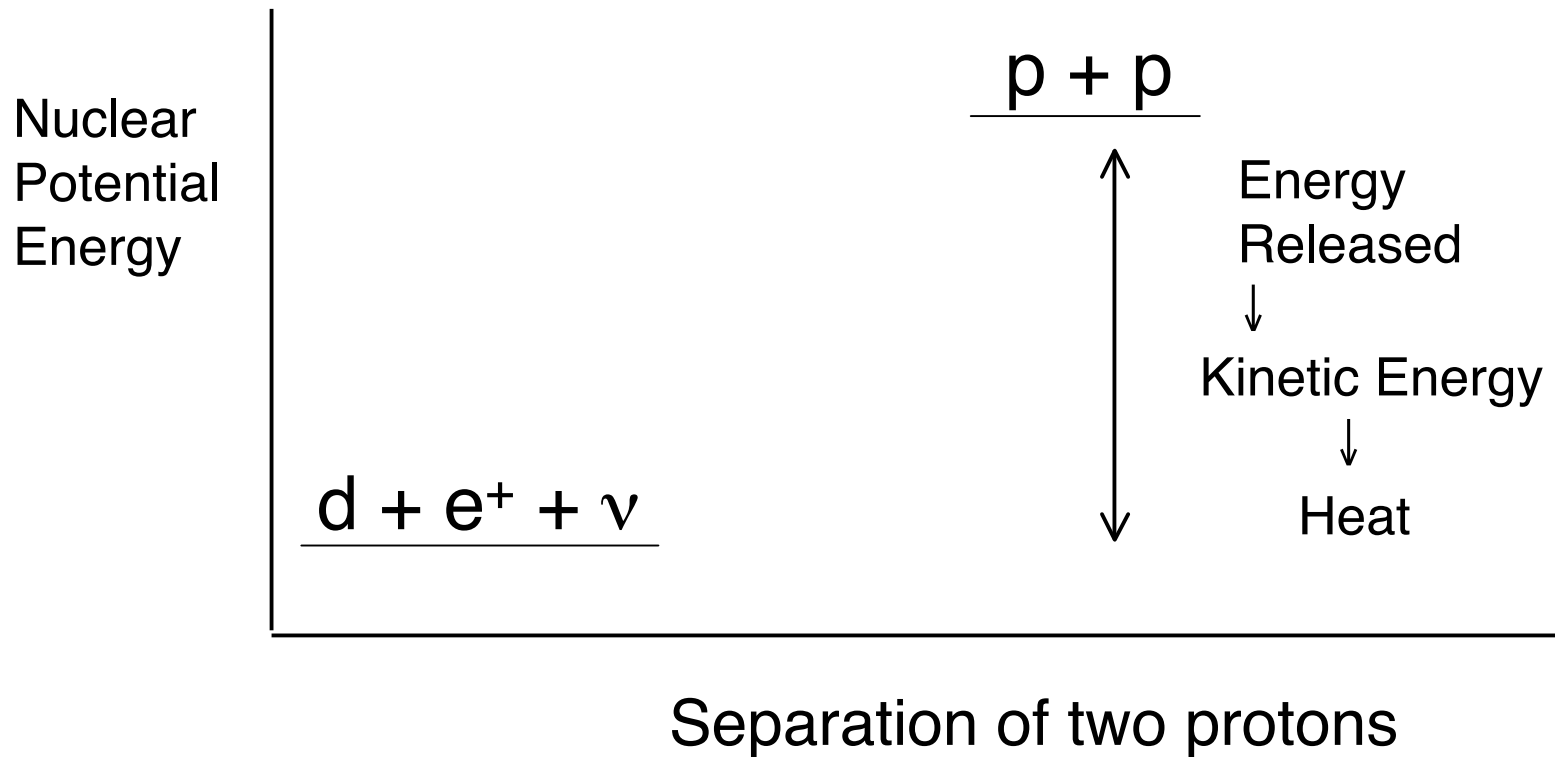
Each has potential energy. Nuclear potential energy can be released by nuclear reactions.

e.g. 1st step: $p + p \rightarrow d + e^+ + \nu$

d = deuteron = proton + neutron

e^+ = positron (antiparticle of electron)

Nuclear Potential Energy



The energy released by nuclear reactions
supplies heat \rightarrow pressure

Resists gravity \Rightarrow stable star

Electromagnetic Barrier

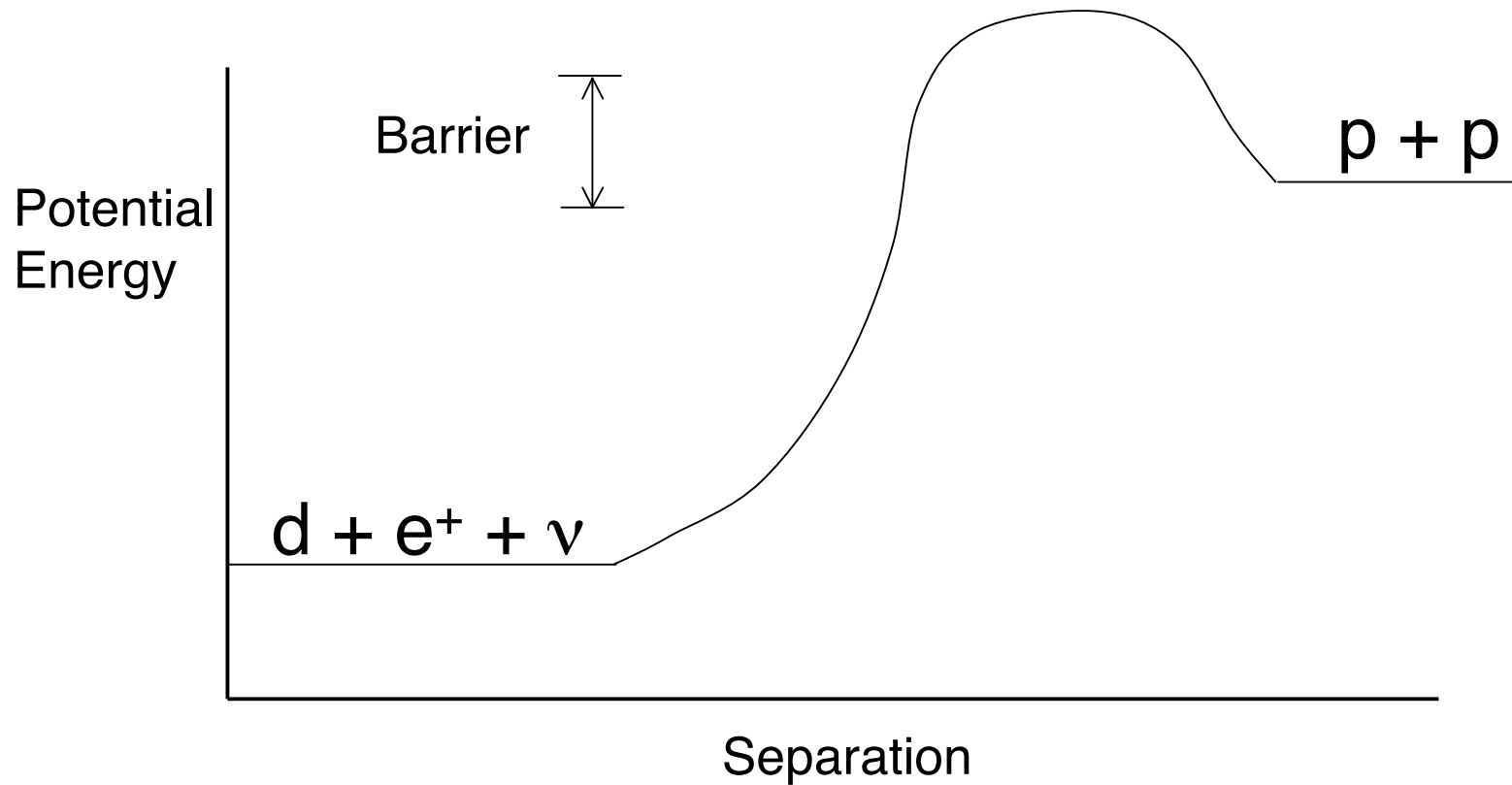
Why do we need high T ($\sim 10^7$ K)?

Protons have positive electric charge

Like Charges Repel

As protons approach, repulsion grows
corresponds to climbing hill of electromagnetic
potential energy

Electromagnetic Barrier



Barrier is really much higher than $\frac{3}{2} k \cdot 10^7 \text{ K}$

Very few can get over barrier

⇒ Stars live a long time rather than exploding

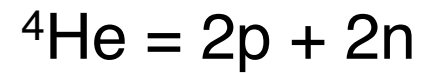
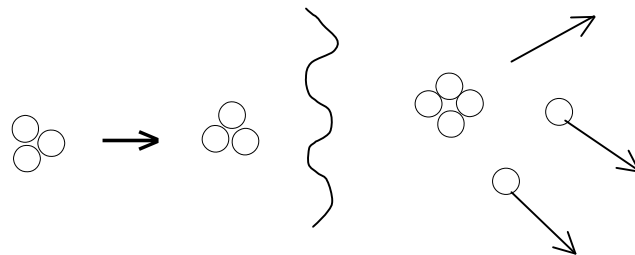
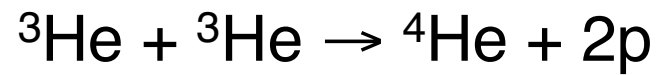
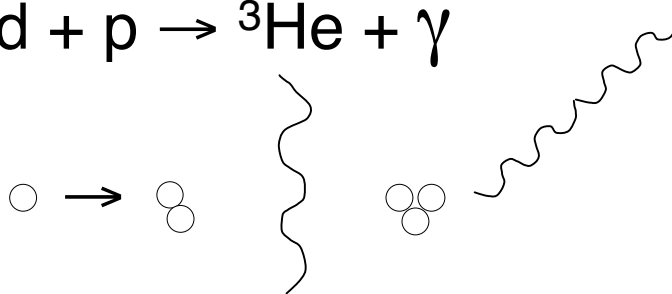
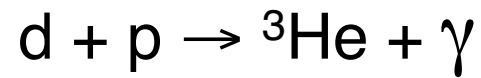
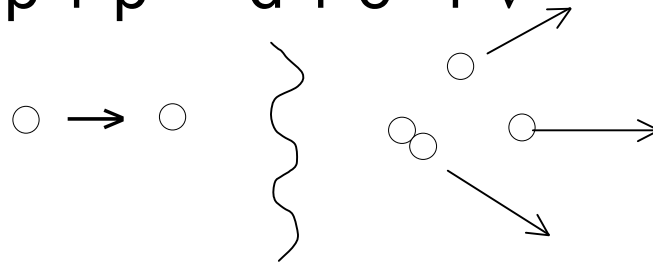


Questions

- Why do nuclear reactions produce a long-lived system in star, but an explosion in a bomb?
- What will happen when a star's fuel runs out?

○	proton	p
○	neutron	n
○	positron	e ⁺
○	neutrino	ν
~~~~~	photon	γ

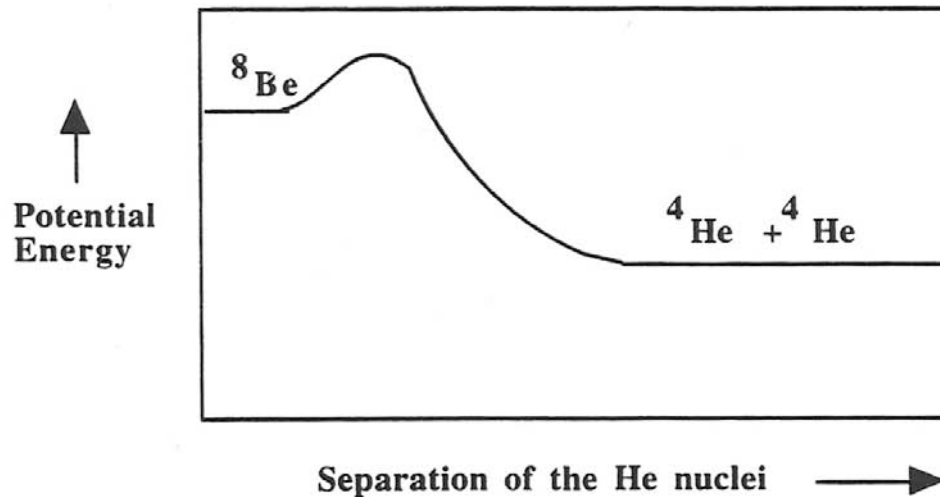
# Nucleosynthesis Again



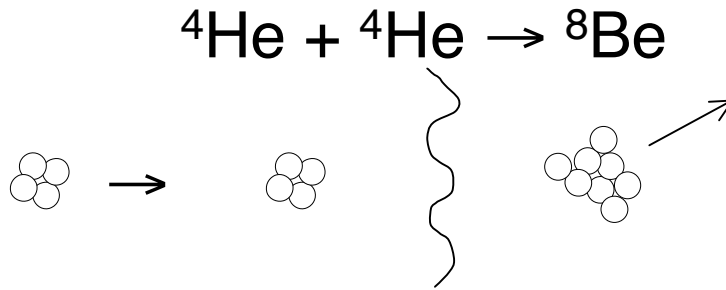
How to get past helium? We need C, O, N, P, S, ...



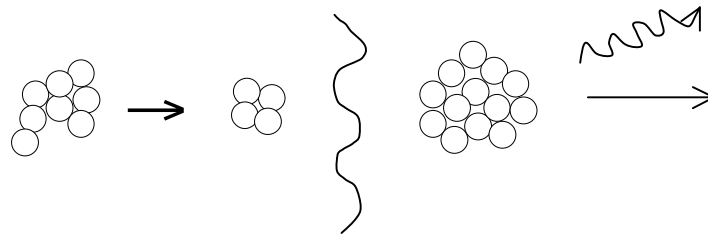
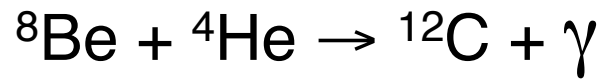
Problem:  ${}^8\text{Be}$  has more nuclear potential energy than parts  
It is unstable (radioactive).



Need another  ${}^4\text{He}$  to hit  ${}^8\text{Be}$  before it falls apart



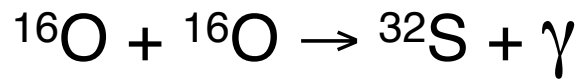
$${}^8\text{Be} = 4p + 4n$$



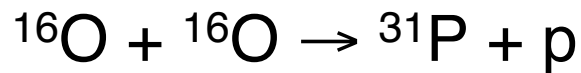
$${}^{12}\text{C} = 6p + 6n$$



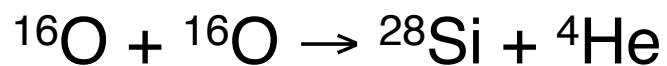
$${}^{16}\text{O} = 8p + 8n$$



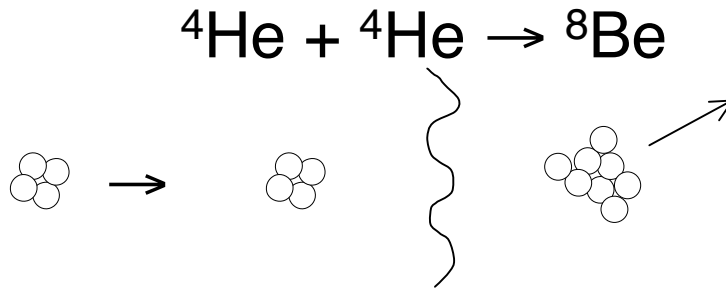
Sulfur



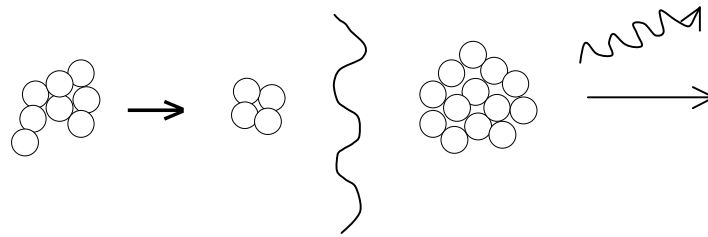
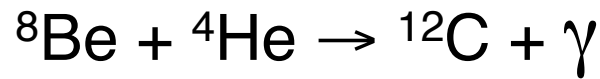
Phosphorus



Silicon



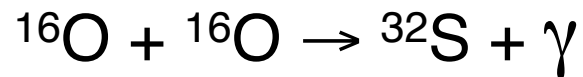
$${}^8\text{Be} = 4p + 4n$$



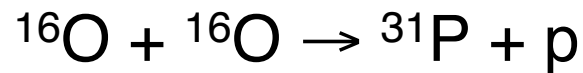
$${}^{12}\text{C} = 6p + 6n$$



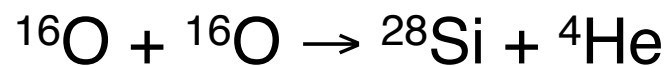
$${}^{16}\text{O} = 8p + 8n$$



Sulfur



Phosphorus

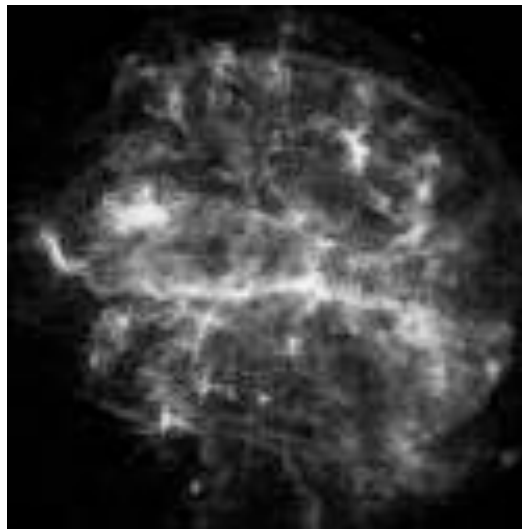


Silicon



# Questions

- What was needed to make the bioelements?
- Are any missing?
- How do the bioelements get out of the star?



# Summary

Heavy elements needed for life were created  
by  
early generations of massive stars.

Except for H, we are made of star debris

Natural forces (Gravity, EM, Nuclear) produced  
first evolution of matter from simple to complex  
(p  $\rightarrow$  heavy elements)