Future of Life in the Solar System

Future of Life in Solar System

Terraform other planets (Mars most likely)

Space Colonies

Solar Power from space

Dyson spheres

Robots

Von Neumann Devices

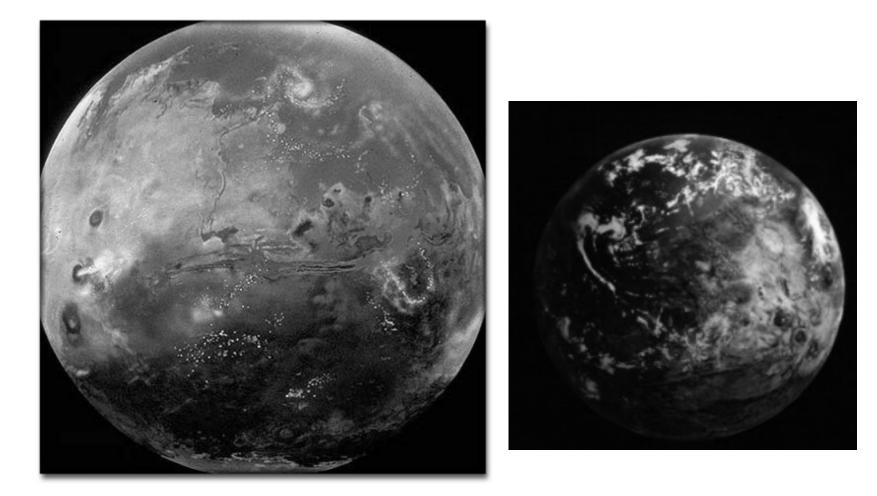
Future of Life in the Solar System

Seed other planets with "bio-engineered organisms"

Use these to make more habitable for humans

Terraform (need H₂O, O₂, O₃) e.g. Melt polar caps on Mars (10¹⁴ tons of ice) 2500 to 10000 years to build up atm. pressure, get liquid water

Terraformed Mars



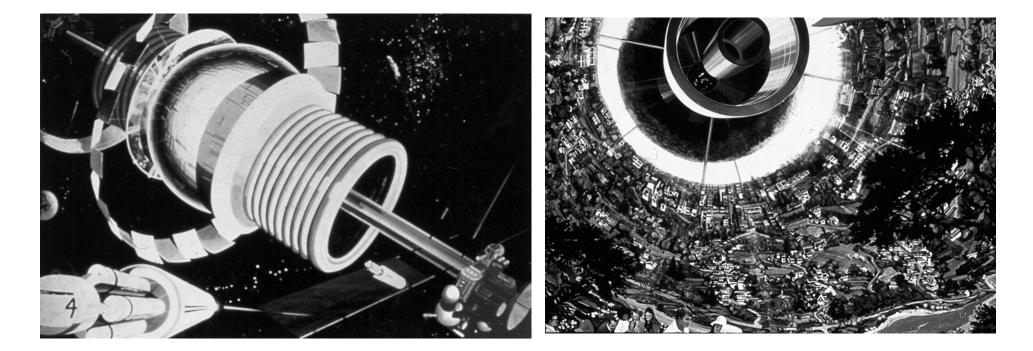
Ocean in northern lowlands covers 25% of planet

Living in Space to Robots...

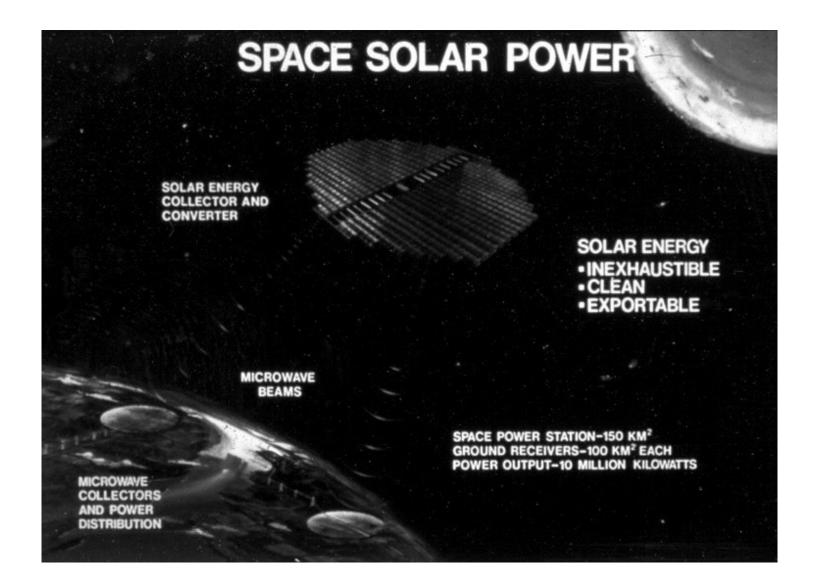
Space colonies Solar Power satellites U Dyson sphere (Type II Civilization)

Role of Robots

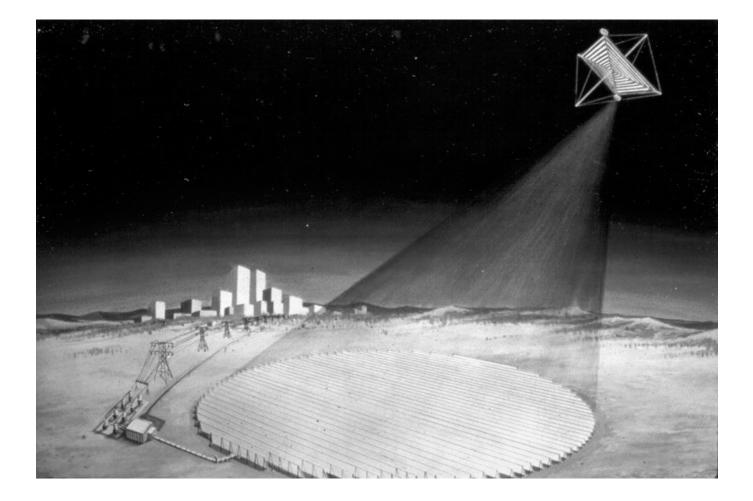
Space Colony (Island One)



Solar Power Satellites



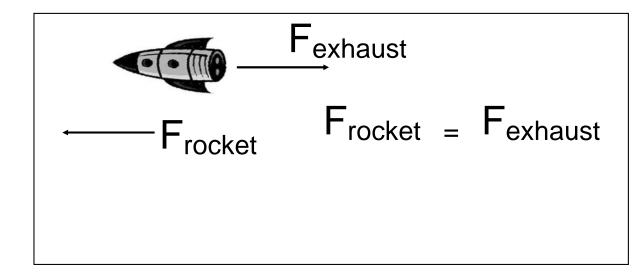
Solar Power Satellite



Rockets

Principle:

Newton's Third Law



1. Exhaust velocity $V_e \ (\text{km s}^{-1})$ $V_e \ \propto \ \sqrt{\frac{T}{M}}$

2. Thrust (Force) $F = \stackrel{\circ}{M} V_e$ (Newtons, Pounds)

$$\stackrel{\circ}{M}$$
 = rate of mass ejected)

3. Mass ratio

 $R_M = Total Mass at Takeoff$ Mass After Fuel Used Up

4. Specific impulse (s.i.)

Thrust(Newtons/kg/sec,Rate of Fuel UsePounds/Pounds/sec = "sec")

A measure of efficiency.

Highest possible s.i. with chemical fuels is < 500

RocketsFgravFthrustTo take off: Thrust > Weight

To escape gravity $v > v_{esc} = 11.2 \text{ km s}^{-1}$ (7 miles/sec)

Rockets

Multi-stage Rockets

- Space Shuttle: Mass = 2×10^6 kg
- $F_{thrust} = 29 \times 10^6$ Newtons $R_M = 68$
- s.i. = 455 sec. ~ best possible with chemical fuel

For more adventurous exploitation of Solar System Probably want Nuclear Propulsion Fission could give s.i. = 1.5×10^6 sec (in principle, more likely to get 20,000 sec)

Current Initiative

- Human mission to Mars
- Several attempts to get started in past
- Exploration Vision in 2004
 - First return to Moon
 - Then Mars
 - Long-term program needed
 - http://www.nasa.gov/missions/solarsystem/explore_main.html

Evaluating your Drake Equation

Basic Ideas

- Number of Civilizations in our Galaxy
 - Product of rate of emergence and L
 - Running product gives rate for each step
 - Until L, we have rates
 - Through f_c, we get "communicable" civilizations
 - Multiplying by L gives the number (N)
 - Assumes "steady state" between birth and death of civilizations

Drake Equation:

$N = R * f_p n_e f_l f_i f_c L$

- N = number of communicable civilizations in our galaxy
- R = Rate at which stars form

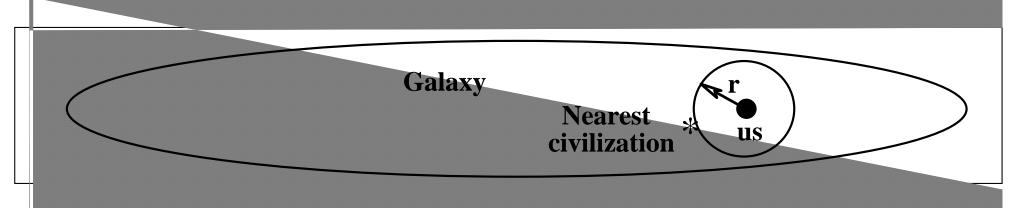
f

f_c

r

- f_p = Fraction of stars which have planetary systems
- n_e = Number of planets, per planetary system, which are suitable for life
 - = Fraction of life bearing planets where intelligence develops
 - Fraction of planets with intelligent life which develop a technological phase during which there is a capacity for and interest in interstellar communication
 - = Average of lifetime of communicable civilizations
 - = Average distance to nearest civilization

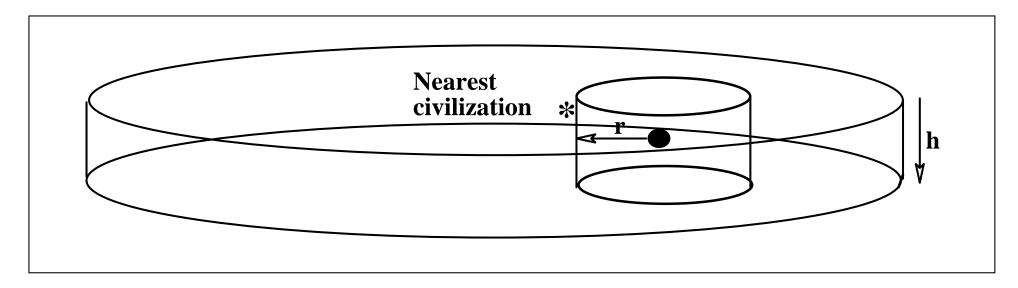
Distance to Nearest Neighbor 1. Assume civilizations spread uniformly but randomly through galaxy



r = radius of imaginary sphere centered on us that touches nearest civilizaztion search vol ∝ r^3

$$\Rightarrow r = \frac{10^4 \text{ ly}}{\text{N}^{1/3}}$$

Distance to Nearest Neighbor



If N < 8000, r from previous formula is 500 ly About equal to thickness of Galaxy

```
Use cylinder for search vol \propto r^2 h
so r = \frac{5 \times 10^4 \text{ ly}}{\text{N}^{1/2}}
```





	R	f p	n _e	f _l	f _i	f _c	L	Ν	r
Estimate	50	1	1	1	1	1	$5 imes 10^9$	$2.5 imes 10^{11}$	1.6 İ y
Birthrate	50	50	50	50	50	50	│ │ │		

2.5 out of 4 stars

If N > 8000, $r = \frac{10^4 \text{ light years}}{N^{1/3}}$ If N <8000, $r = \frac{5 \times 10^4 \text{ light years}}{N^{1/2}}$

Angela Angst



	R	f p	n _e	f _l	f _i	f _c	L	Ν	r
Estimate	5	0.1	0.1	0.01	0.01	0.01	100	5×10 ⁻⁶	
Birthrate	5	0.5	0.05	5 x 10 ⁻⁴	5 × 10 ⁻⁶	5×10 ⁻⁸			

Never two civilizations at same time

If N > 8000, $r = \frac{10^4 \text{ light years}}{N^{1/3}}$ If N < 8000, $r = \frac{5 \times 10^4 \text{ light years}}{N^{1/2}}$

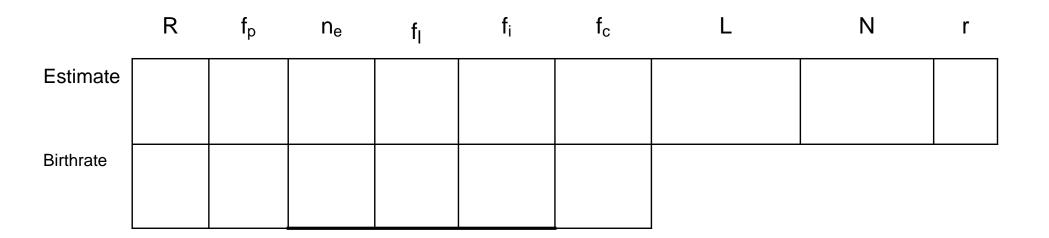


	R	f p	n _e	f _l	f _i	f_{c}	L	Ν	r
Estimate	10	0.5	0.89	0.5	0.7	0.6	1 × 10 ⁶	$9.4 imes10^5$	100
Birthrate	10	5	4.45	2.23	1.56	0.94			
				1 out of × 10 ⁵ sta 0 × 10 ⁵ :					
If N > 8000,			r =		light yo N ^{1/3}		I	0 ~ 10 .	- 10
If N < 8000,			r =	<u>5 ×</u>	10 ⁴ lig N ^{1/2}	ht yea	rs -		

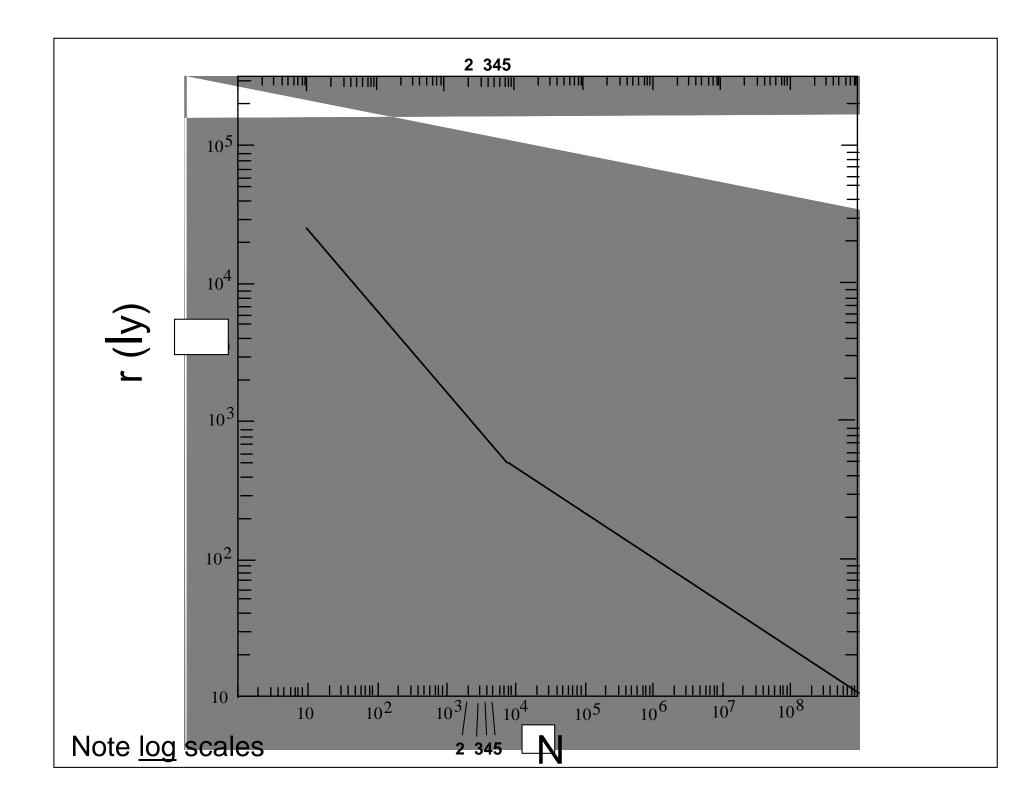
Evaluating YOUR Drake Equation

- Almost no answers are wrong
 - It must be possible for us to exist
 - N must be no greater than the number of stars in the Galaxy
 - May imply limit on L
- Ways to evaluate:
 - Plug into equations
 - Use calculator on web
 - <u>http://www.as.utexas.edu/astronomy/education/drake/drake.html</u>
 - Ask us for help

Your Drake Equation



If N > 8000, $r = \frac{10^4 \text{ light years}}{N^{1/3}}$ If N < 8000, $r = \frac{5 \times 10^4 \text{ light years}}{N^{1/2}}$



Points to bear in mind

- r is based on assuming spread uniformly
 Could be less in closer to center of MW
- r is based on averages
 - Could be closer but unlikely
- r is less uncertain than N
- Since signals travel at c, time = distance in ly
- If L < 2r, no two way messages