The History and Philosophy of Astronomy

(Lecture 24: Hubble II)

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The University of Texas at Austin
Edwin P. Hubble: Mariner of the Nebulae

- 1889 (Missouri) — 1953 (Pasadena)

- leading observational astronomer of 20th century:
  - discovers galaxies (1924):
    Milky Way but one of innumerable “island universe”
  - expansion of the universe (1929):
    Hubble’s Law:
    \[ v = H_0 \times d \]
Mount Wilson Observatory

- 1919 onwards: world’s largest telescope
Classification of Nebulae (=galaxies)

- 1923: Tuning-fork diagram
Big Q: How do the Spiral Nebulae Move?

- figure out radial velocities by taking spectra, and measure the shift of spectral lines using the Doppler effect!
The Doppler Effect (1842)

Christian Doppler (1803-53)
The Doppler Effect for Spectral Lines

- Calculate radial speed:

\[
\frac{\text{wavelength shift}}{\text{rest wavelength}} = \frac{\text{speed in line-of-sight}}{\text{speed of wave}}
\]

\[
\frac{\Delta \lambda}{\lambda} = \frac{v}{c}
\]

- By measuring the shift in wavelength of spectral lines, one can figure out radial velocity of source with high precision!
The Flight of the Spiral Nebulae (1912)

- Important discovery: most spiral nebulae display redshifts in their spectra!

- From Doppler effect: Do spiral nebulae move away from us???
Solving Einstein’s Equations of GR

• 1917: Einstein constructs model of the universe that is eternal and static

• balance between attractive gravity and repulsive cosmological constant ("anti-gravity")

• finite but without boundary (spatially closed)
Solving Einstein’s Equations of GR

• 1917: de Sitter constructs a model of the universe that contains no matter, but predicts motion!

Willem de Sitter
(1872-1934)

• an empty universe!
Solving Einstein’s Equations of GR

• 1917: de Sitter constructs a model of the universe that contains no matter, but predicts motion!

• particles are “scattered away” from origin ("de Sitter effect")
  -- the larger the distance, the larger the apparent speed!
Hubble and the Distance to Andromeda (M31)

- October 1923: He obtains photograph of M31

- Hubble discovers a Cepheid variable in Andromeda!
What are the Cepheids?

- A: Pulsating stars (periods of ~ few days)

_Cepheid_ variables: outward pressure (P) and inward gravity compression are out of sync, so star changes size and temperature: it **pulsates**. _RR-Lyrae_ variables are smaller and have pulsation periods of less than 24 hours. Also, their light curve looks different from the Cepheid light curve.
Cepheids as Standard Candles

Standard Candle
As a car approaches on a highway, its lights appear to get brighter. However, the amount of light or energy produced by the lights is constant. The apparent brightness of the headlights is a function of how far away the car is. The closer the car, the brighter the headlights. Astronomical objects that produce specific amounts of energy can be used to determine the distance to that object if the apparent brightness is measured. These objects are known as Standard Candles.
Cepheids as Standard Candles

1) Measure Period

2) Calculate luminosity

3) Calculate distance

- inverse-square law:
  \[ \text{flux} = \frac{\text{luminosity}}{\text{distance}^2} \]
Hubble teams up with Humason

- Hubble and Humason become partners in nebular research

Milton L. Humason (1891-1972)
Discovery of the Hubble Law (1929)

• Great Discovery: Redshift is proportional to distance ("Hubble Law")
• greatly improve accuracy by reaching larger distances!
Meaning of the Hubble Law

- recession speed = (Hubble) constant x distance
  
  \[ v = H_0 \times d \]

- \( (H_0=500 \text{ km s}^{-1} \text{ Mpc}^{-1}) \)

- Hubble initially did not interpret his law as implying an expanding universe! (He only referred to “de Sitter effect”)}
Solving Einstein’s Equations of GR

- 1922: an expanding universe (GR without cosmological constant)

Alexander Friedmann
(1888-1925)
Solving Einstein’s Equations of GR

• 1927: Lemaitre independently (re-) discovers the expanding-universe solutions of GR

Georges Lemaitre
(1894-1966)
The Expanding Universe (early 1930s)

• Lemaitre interprets Hubble’s Law as indicating that space itself expands!

• redshifts due to stretching of expanding space!
The Expanding Universe (early 1930s)

• Lemaitre interprets Hubble’s Law as indicating that space itself expands!

• redshfits due to stretching of expanding space!
The Age of the Expanding Universe

• Estimate expansion age of the universe:

\[
\text{time since you were together} = \frac{\text{distance}}{\text{velocity}} = \frac{d}{v}
\]

Later, how much time has passed?

\[
\text{time} = \frac{\text{distance to a given galaxy}}{\text{its velocity of recession}} = \text{age of the universe}
\]

\[
\text{age of the universe} = \frac{1}{H_0}
\]
History of $H_0$

Compilation by John Huchra

Baade identifies Pop. I and II Cepheids

“Brightest stars” identified as H II regions

Jan Oort
Hubble Wars

reasonable convergence only in last decade – see later
Big Q: What happened in the past, when all the galaxies where close together???

• an initial “Big Bang”???
Hubble (part 2)

• Discovery of Recession of Spiral Nebulae (1929-31)
  - already known: most spirals show redshifts in their spectra
  - Hubble and Humason use Cepheids and other standard candles to determine distances to spiral nebulae
  - “Hubble’s Law”: \( v = H_0 \times d \)
  - Initial estimate of “Hubble’s constant”: \( H_0 = 500 \text{ km s}^{-1} \text{ Mpc}^{-1} \)

• Interpretation of Hubble’s Law (1930s)
  - Lemaitre (rediscovering Friedmann’s work) proposes solutions to Einstein’s General Relativity for an expanding universe!
  - Naturally explains “Hubble’s Law” \( (v = H_0 \times d) \) with expansion of space itself!