



Astro 301/ Fall 2005 (48310)



Introduction to Astronomy

Instructor: Professor Shardha Jogee

TAs: David Fisher, Donghui Jeong, and Miranda Nordhaus

Lecture 6 + 7: Tu Sep 20, Th Sep 22

Recent and upcoming topics in class

The upcoming topics are the most math-oriented chapters of this entire course, but we will go through them slowly and use plenty of examples in class.

--- The Four Fundamental Forces

- The Force of Gravity and Newton's Universal Law of Gravitation (end of Lec 5)
- Electromagnetic Forces
- Strong and Weak forces

--- Relating Motion to Forces using Newton's Three Laws of Motion

- Mass, speed, velocity, acceleration

- Newton's 1st and 2nd Law

- Applying Newton's 1st and 2nd laws to objects moving along a circular orbit

--> Centripetal acceleration and centripetal force (end of Lec 7)

--> Centripetal acceleration of electrons orbiting the nucleus in an atom

--> Centripetal acceleration of planets orbiting about the Sun in Solar system

--> Centripetal acceleration of stars orbiting the center of a galaxy

- Newton's 3rd law

- Using Newton's 3rd law to understand weight and weightlessness

--- Motion of galaxies on very large scales : Hubble's Law and the Expansion of the Universe

Lecture 6

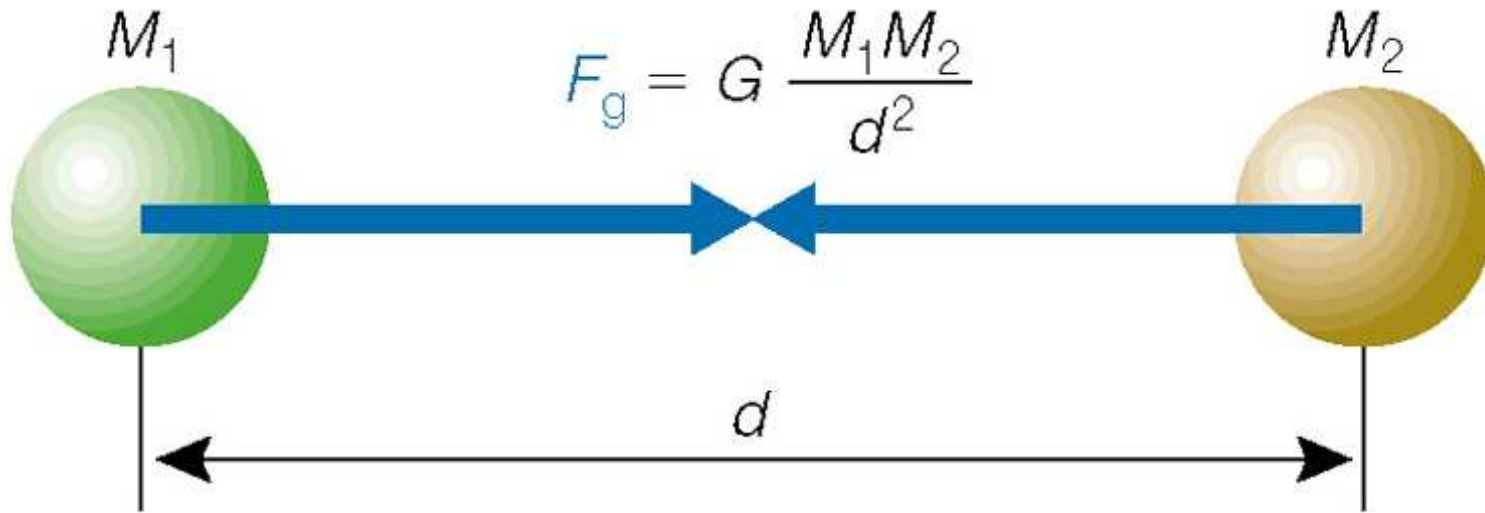
Announcements

- Hwk 1 due today at start of class.
- ALWAYS CHECK THE CLASS WEBSITE for most current posted notes, announcements, and help sessions

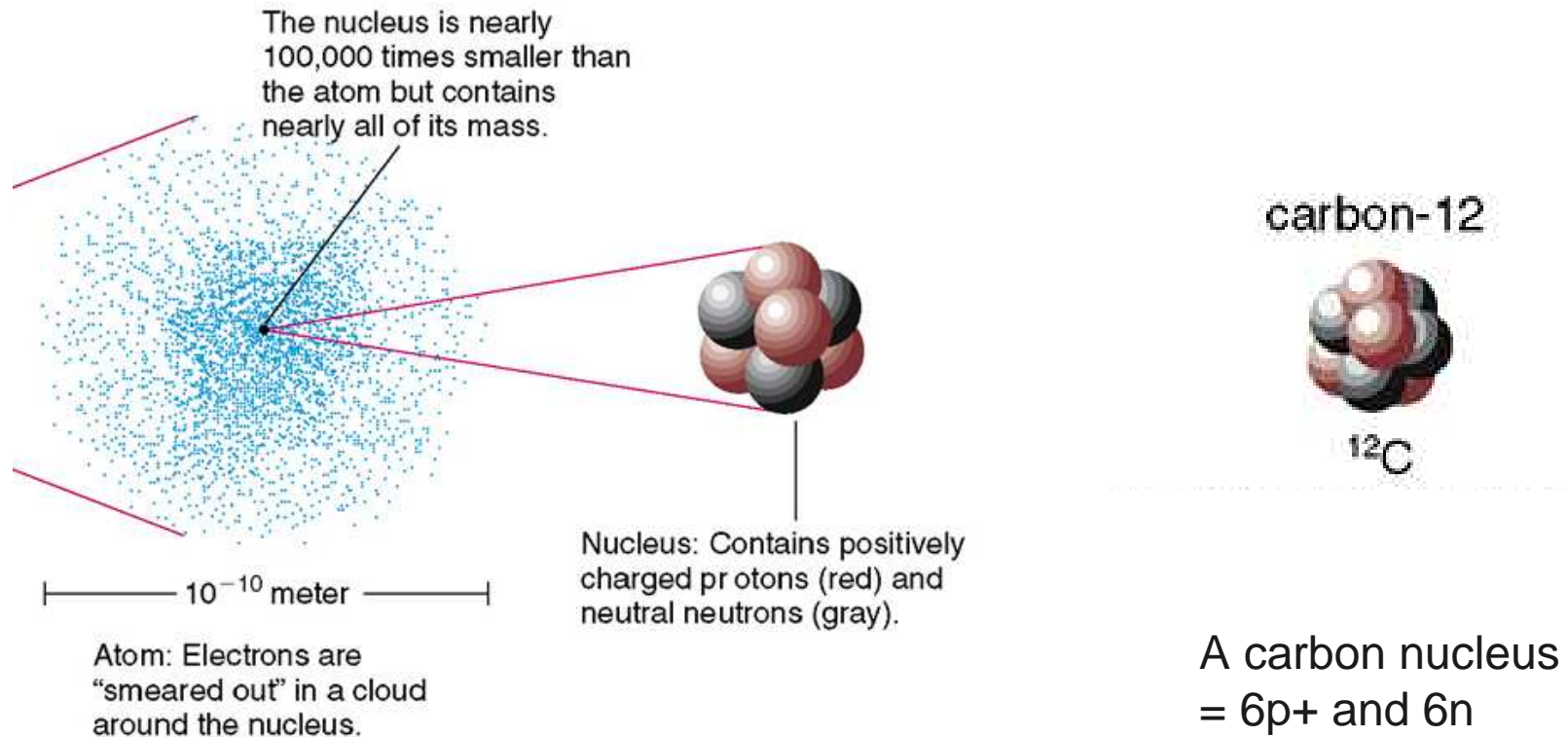
<http://www.as.utexas.edu/~sj/a301-fa05/>

The Four Fundamental Forces

Newton's Universal Law of Gravitation



Forces within the nucleus of an atom

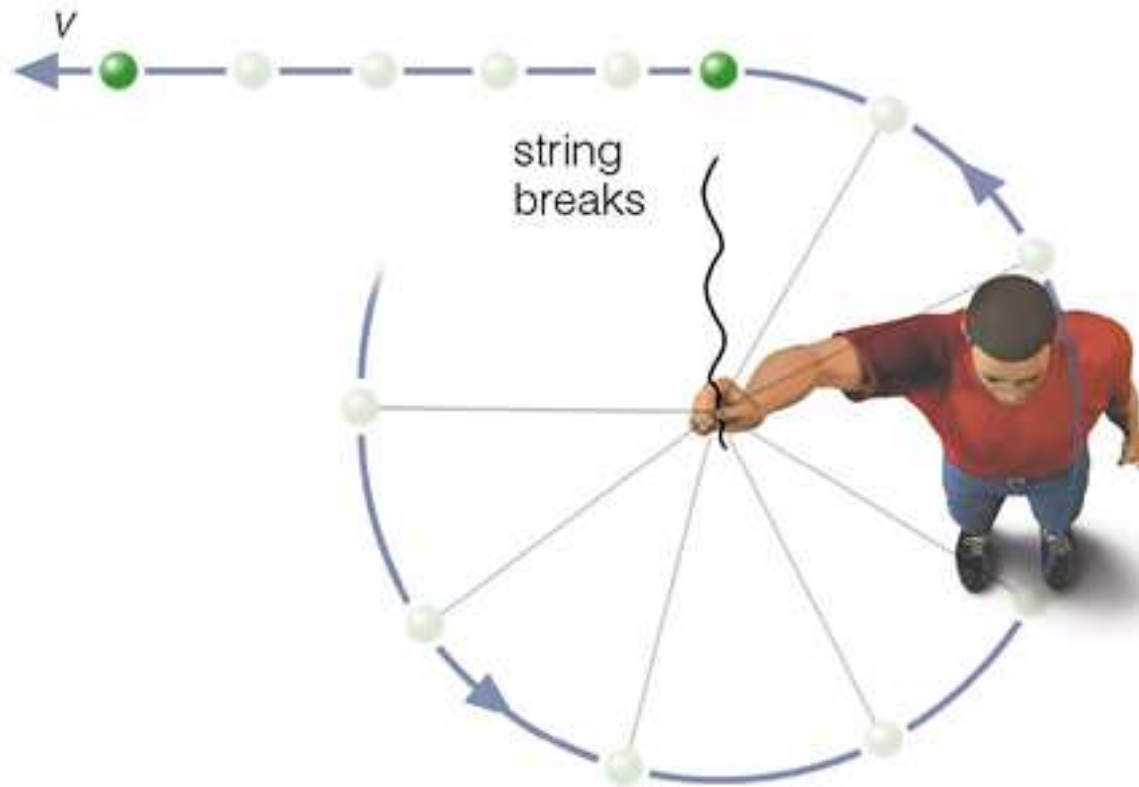


A carbon atom is made of 6 e- orbiting a tiny carbon nucleus

- See in-class notes: Forces acting are gravity, EM between e- p+, EM between p+ p+, Strong Forces, Weak Forces

Relating Motion to Forces: Newton's Three Laws of Motion

Speed, Velocity, Acceleration



In class notes

Lecture 7

Announcements

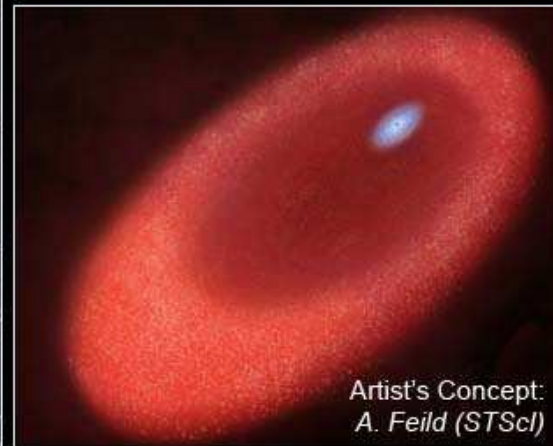
- Quiz 1 back today : 75% Awell done!
- Quiz 2 on Th Sep 29, based on lectures 4,5,6,7,8
- Homeworks handed after class on due day: 80% credit
Homeworks handed 1 or more days late: 50% or lower credit
- Class website : <http://www.as.utexas.edu/~sj/a301-fa05/>

Picture of the day

Andromeda Galaxy Nucleus ■ M31



Hubble Space Telescope ■ WFPC2



NASA, ESA, R. Gendler, T. Lauer (NOAO/AURA/NSF), and A. Feild (STScI)

STScI-PRC05-26

**HST
and UT
press
release
this
week
Sep
2005**

**Stars close to BH experience tremendous force of gravity and orbit very fast:
1000 kilometers a second.... Would take 40 s to circle the Earth !**

Newton's Law of Gravity and Laws of Motion



Sir Isaac Newton (1642-1727)
Born in England



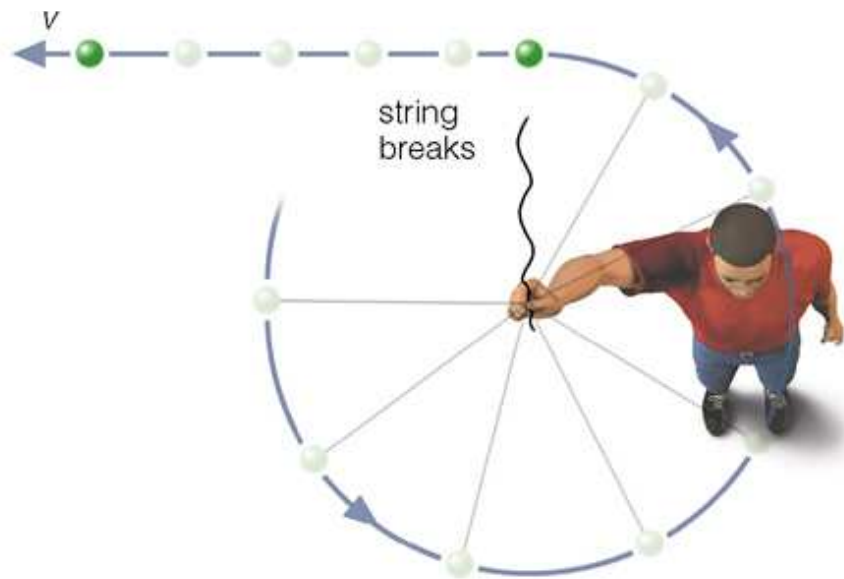
- Student (1661); Lucasian Professor of Mathematics (1669) at Trinity College, Cambridge Univ
- Unified “the Earth and the Heavens” with his laws of gravity and motion (1665-1666)
- Published “Principia” (Mathematical Principles of Natural Philosophy) in 1687

Newton's 2nd Law of Motion

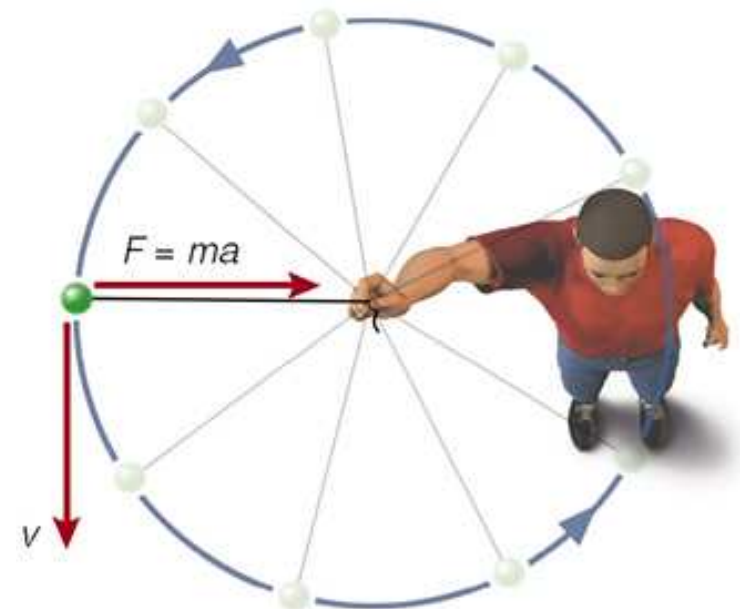


Force $F = m a$

Applying Newton's 1st and 2nd laws to objects moving on a circular orbit: centripetal acceleration



By Newton's 1st law: Without a net force (e.g., if string breaks), the rotating object would move at a **constant velocity** (i.e. at constant speed along a straight line).



An object moving in a circle of radius R at constant speed v experiences
à a continuously changing velocity
à an acceleration \mathbf{a} called a centripetal (center-seeking) acceleration

By Newton's 2nd law a force $\mathbf{F} = M\mathbf{a}$ must act on it to provide this acceleration.

Centripetal acceleration of different orbiting objects

- > Centripetal acceleration of electrons orbiting the nucleus in an atoms
- >.Centripetal acceleration of planets orbiting about the Sun in Solar system
- > Centripetal acceleration of stars orbiting the center of a galaxy
- > Centripetal acceleration of stars orbiting about a central black hole