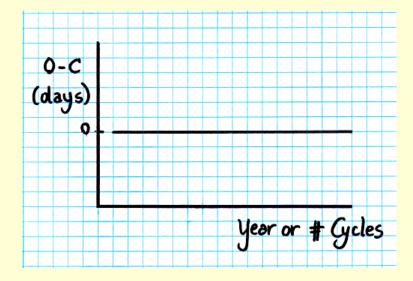
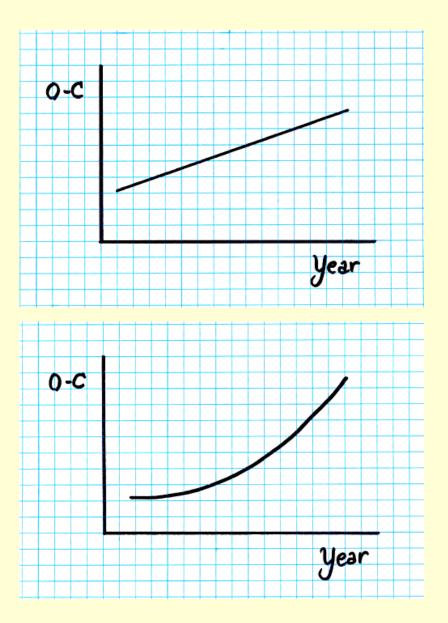
### The O-C diagram

O-C: Observed time – Calculated time



Period is correct and constant.

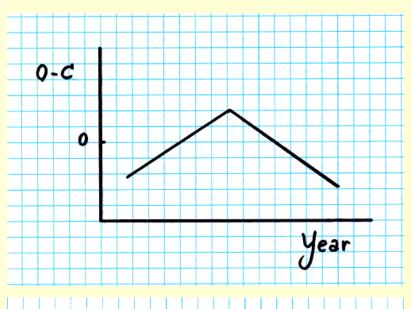
### The O-C diagram



Period is incorrect but constant.

Period is changing (increasing), e.g., due to evolution.

### The O-C diagram

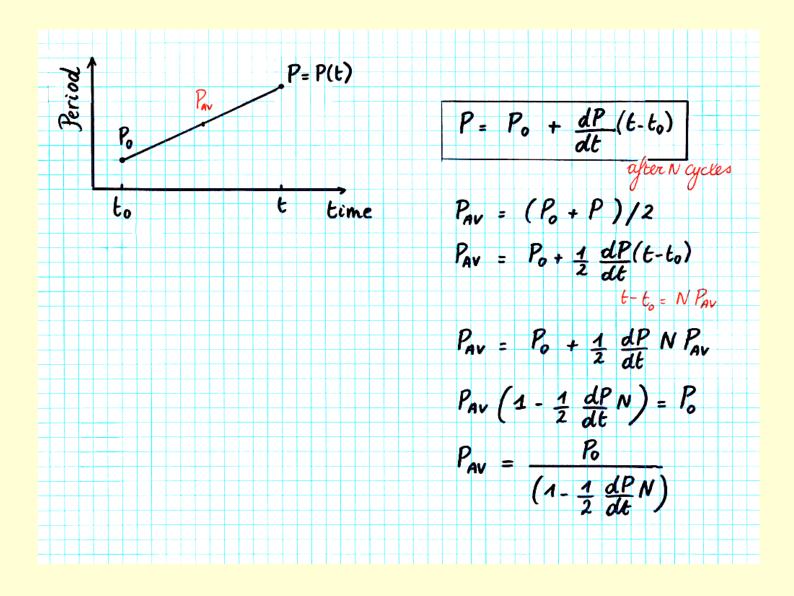


Abrupt change from one constant period to another,

e.g., in the RR Lyrae star SZ Hya.



Who knows ???



But we don't know 
$$P_0$$
,  $P_{AV}$ , ... We only know the trial period  $P_V$ 

$$(O-C) = NP_{AV} - NP_V$$

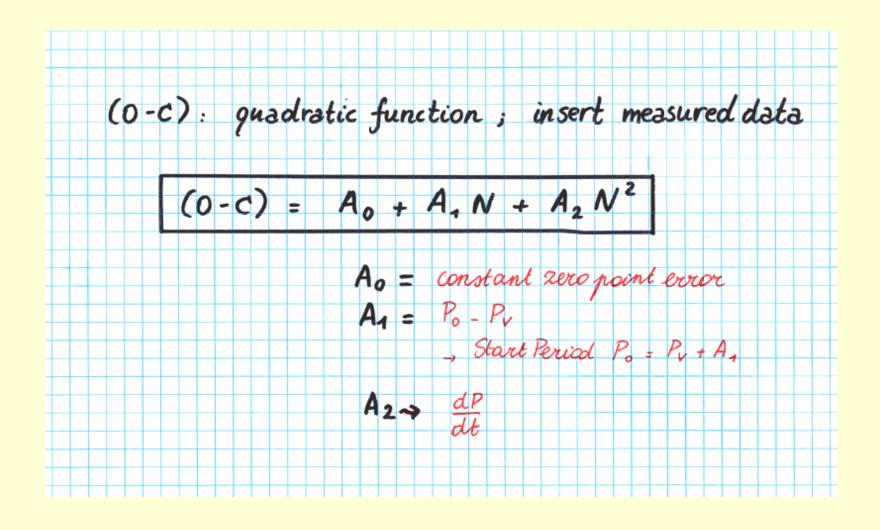
$$= \frac{NP_0}{1 - \frac{1}{2} \frac{dP}{dt} N} - NP_V$$

$$1 - \frac{1}{2} \frac{dP}{dt} N$$
Series expansion ,  $\frac{1}{2} \frac{dP}{dt} N \ll 1$ 

$$(O-C) = NP_0 + \frac{1}{2} N^2 P_0 \frac{dP}{dt} - NP_V$$

$$(O-C) = N(P_0 - P_V) + N^2 (\frac{1}{2} P_0 \frac{dP}{dt})$$

$$L_V = \frac{1}{2} \frac{dP}{dt} + \frac{1}{2} \frac{dP}{dt} = \frac{1}{2} \frac{dP}{dt} + \frac{1}{2} \frac{dP}{dt} = \frac{1}{2} \frac{dP}{dt} = \frac{1}{2} \frac{dP}{dt} + \frac{1}{2} \frac{dP}{dt} = \frac{1}{2} \frac{dP}{dt} =$$



$$(0-c) = NP_0 + \frac{1}{2}N^2P_0 \frac{dP}{dt} - NP_v$$

$$(0-c) = N(P_0 - P_v) + N^2(\frac{1}{2}P_0 \frac{dP}{dt})$$

$$L \quad quadratic function$$

If the adopted period 
$$P_{v}$$
 is correct:  $P_{v} = P_{o}$ , then
$$(0-C) = \frac{1}{2} N^{2} P_{o} \frac{dP}{dt}$$

$$= \frac{1}{2} t^{2} \left(\frac{1}{P_{o}} \frac{dP}{dt}\right)$$

### O-C and period changes

Some remarks concerning Period Changes:

1. (0-c) in d (days) = ... + 
$$\frac{1}{2}$$
 N<sup>2</sup>P<sub>0</sub>  $\frac{dP}{dt}$ 

$$\stackrel{\cong}{=} \dots + \frac{1}{2} \left( \frac{1}{P} \frac{dP}{dt} \right) t^{2}$$

$$\stackrel{d}{=} \frac{d^{2}}{d^{2}}$$

NB:  $\frac{1}{P} \frac{dP}{dt}$  often expressed in  $yr^{-1}$  (thus 365.25 x larger!)

2. In publications we often find:

HJD (max) = 
$$T_0 + P_0 \cdot N + \alpha N^2$$
 ( $T_0$  and  $P_0$  in  $d$ ),  
so  $\frac{1}{2P} \frac{dP}{dt} \cdot t^2 = \alpha N^2$  with  $t/p = N$ 

$$\frac{1}{P}\frac{dP}{dt} = \frac{2\alpha}{P^2} \quad (again in d^{-1}, \quad or \quad \alpha = \frac{1}{2} \frac{P}{dt} \quad (in d)$$

$$not \quad yr^{-1}$$

### O-C and period changes

Remarks (continued):

3.) Often HUD (max) = 
$$T_0 + P.N + \frac{B}{2}N^2$$
  
then  $G = 2\alpha = P\frac{dP}{dt}$ 

$$\frac{d(lnP)}{dt} = \frac{1}{P} \frac{dP}{dt}$$

Example: BL Cam

$$HJD(max) = 2443125,8042 + 0,0391.E + 6,153.10^{-13}.E^2$$

$$\left(\frac{1}{P}\frac{dP}{dt}\right) = \frac{2\alpha}{P^2} = 8.05 \cdot 10^{-10} d^{-1} \text{ or } 2.94 \cdot 10^{-7} \text{ yr}^{-1}$$

Example: 
$$\frac{dP}{dt} = 10^{-11}$$
, time span: 10 years

 $(0-C)$  in days =  $\frac{1}{2}$   $(10 \times 365, 25)^2 \times 10^{-11}$  / Po

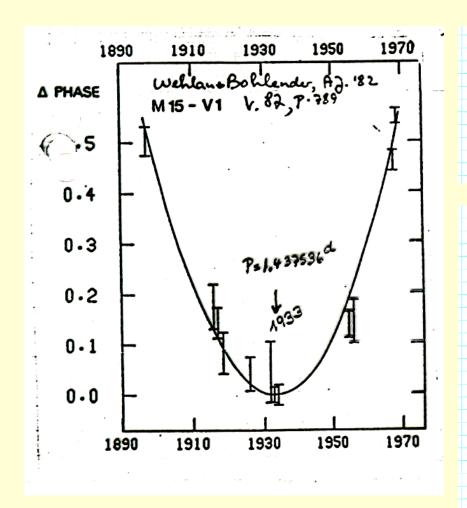
= 6,67 × 10<sup>-5</sup> / Po

Cepheid:  $P = 10$  d  $\Rightarrow$   $(0-C) = 6,67 \times 10^{-6}$  d = 0,58 s

Scuti Star:  $P = 0,1$  d  $\Rightarrow$   $(0-C) = 6,67 \times 10^{-4}$  d = 57,6 s

Pulsar:  $P = 0,03$  s = 3,472 × 10<sup>-4</sup> d

 $\Rightarrow$   $(0-C) = 0,0192$  d : 27,6 min  $?$ 



Example: Period Changes in cluster BL Her Stars

BL Her Stars: Pop. II Cepheids

20 stars known with P between 1.13-7.90 d

(Wehlau L Bohlender, AJ 1982)

1932-1970: 
$$\triangle \phi \sim 0.55 \Rightarrow (0-C) = 0.8 d$$

# Gycles: ca. 9600

Po = 1.437536 d

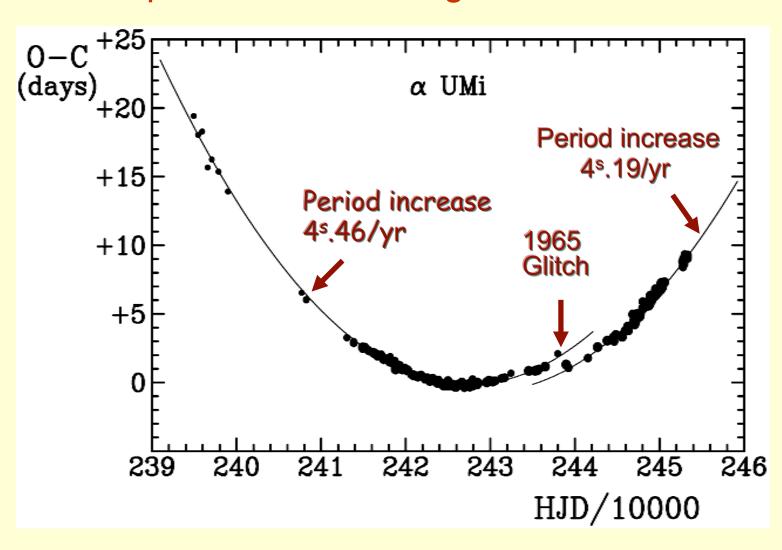
 $(0-C) = \frac{1}{2} N^2 P_0 \frac{dP}{dt}$ 

0.8 =  $\frac{1}{2} (9600)^2 1.437536 \frac{dP}{dt}$ 
 $\Rightarrow \frac{dP}{dt} = 1,2 \times 10^{-8} d/d$ 

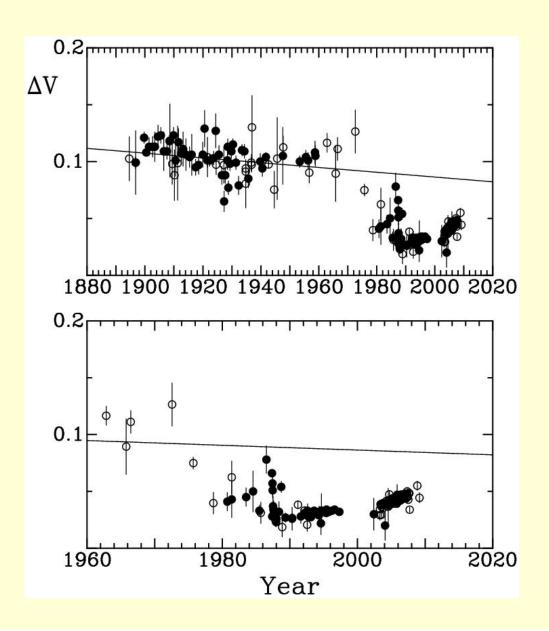
or  $4,4 \times 10^{-6} d/yr$ 

Note: We started at the bottom of the curve!

# Polaris - Period always increasing except for an unusual "glitch" around 1965



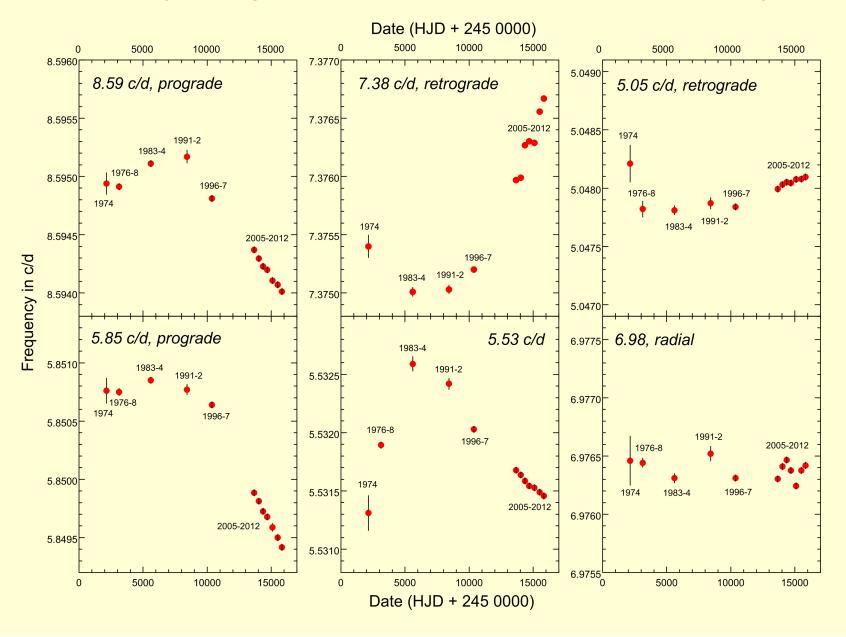
### Polaris – amplitudes near "glitch" around 1965

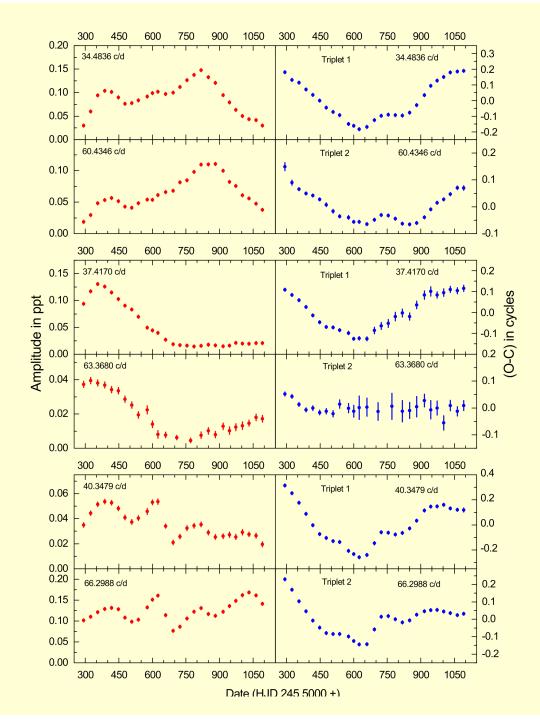


Amplitudes from V<sub>R</sub>

– filled circles
Pe, pg, visual photometry
- open circles
(See Turner 2009)

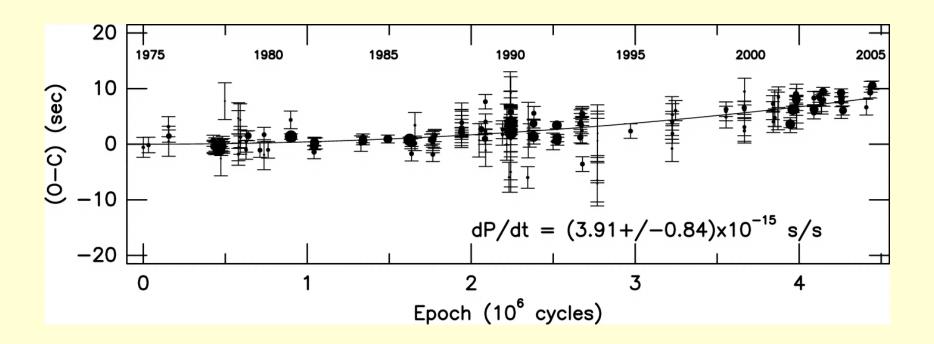
#### Frequency changes of the Delta Scuti star 4 CVn over 40 years





### O-C changes of the Delta Scuti star KIC 8054146

G117-B15A: the most stable optical clock known (Kepler et al. 2005, ApJ 634, 1311)

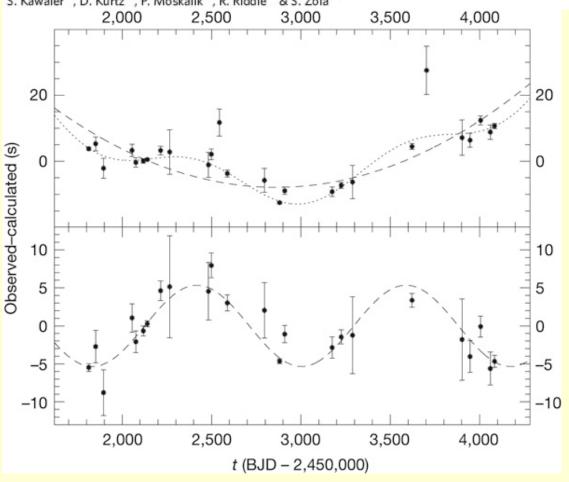


Note: period change due to proper motion (Pajdosz 1995, A&A 295, L17)

### A giant planet orbiting the 'extreme horizontal branch' star V 391 Pegasi

Nature 449, 189 (2007)

R. Silvotti<sup>1</sup>, S. Schuh<sup>2</sup>, R. Janulis<sup>3</sup>, J.-E. Solheim<sup>4</sup>, S. Bernabei<sup>5</sup>, R. Østensen<sup>6</sup>, T. D. Oswalt<sup>7</sup>, I. Bruni<sup>5</sup>, R. Gualandi<sup>5</sup>, A. Bonanno<sup>8</sup>, G. Vauclair<sup>9</sup>, M. Reed<sup>10</sup>, C.-W. Chen<sup>11</sup>, E. Leibowitz<sup>12</sup>, M. Paparo<sup>13</sup>, A. Baran<sup>14</sup>, S. Charpinet<sup>9</sup>, N. Dolez<sup>9</sup>, S. Kawaler<sup>15</sup>, D. Kurtz<sup>16</sup>, P. Moskalik<sup>17</sup>, R. Riddle<sup>18</sup> & S. Zola<sup>14,19</sup>



The pulsating subdwarf V391 Peg

Planet: 3 Jupiter masses, 1.7 AU, 3.2 yr period