

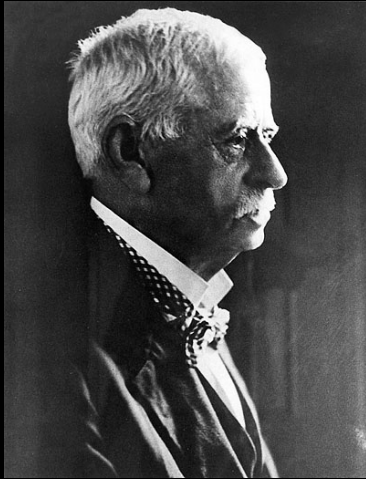
McDONALD OBSERVATORY
THE UNIVERSITY OF TEXAS AT AUSTIN

*Update on the McDonald Observatory Search for
Planets Around White Dwarf Stars:
Dying Stars, Living Planets*

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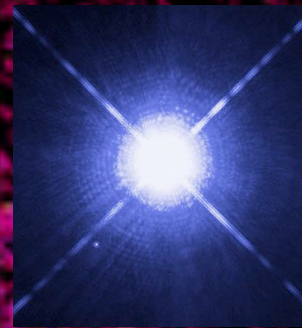




“One of the first tasks to be undertaken by the staff of the McDonald Observatory will be to investigate further the mysteries of the white dwarfs.”

-- May 5th, 1939

White Dwarf Stars: Eddington's *“Impossible”* Star



What Are White Dwarf Stars?

- *Endpoint* of evolution for most stars
- *Homogeneous* in mass and surface composition
- *Uncomplicated* in structure and composition; evolution is just cooling

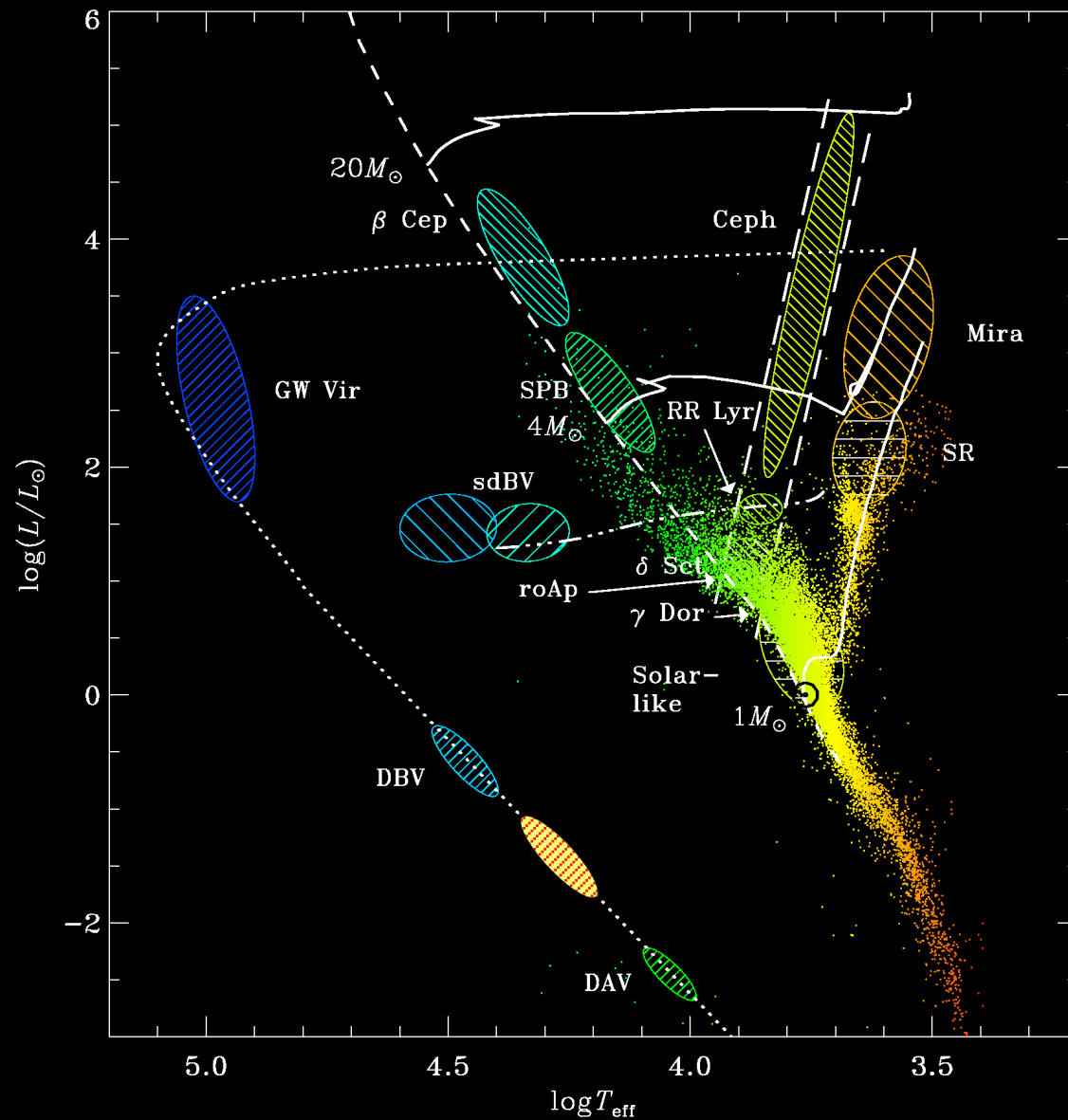
They Shed Their Complexity!



... and Why Should I Care?

- **Representative** (and personal)
 - 98% of all stars, including our sun, will become one
 - Archeological history of star formation in our galaxy
- **Extreme physics:** plasmon neutrinos, internal crystallization, and dark matter, calibrate inertial confinement (laser fusion)
- *Envelopes in same EOS domain as planetary interiors*
- *Debris disks and accreting rocks/asteroid material – flame testing extrasolar rocks ...*
- *A way to find Solar Systems dynamically like our own*

The first new class of pulsating white dwarf found in the last 25 years!



All four have been discovered at McDonald Observatory!

Surface
Brightness
Variations
100-1,000 s



$l = 2$ modes

Nonradial Gravity
Modes: g-modes



$l = 3$ modes

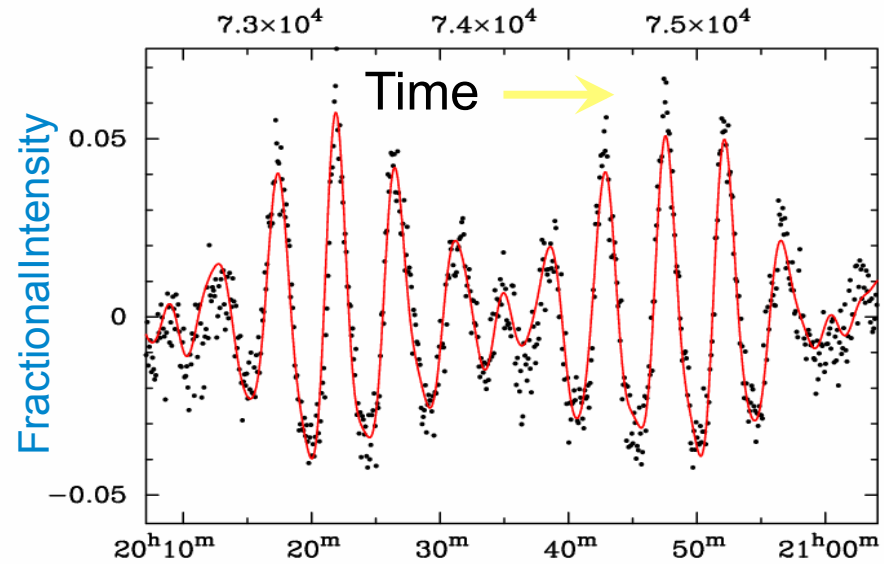
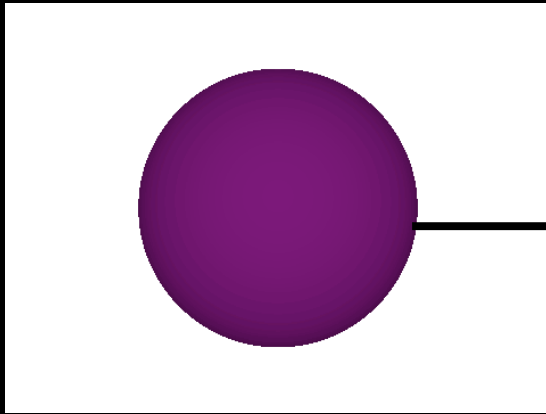
Quantum numbers

1,m,n

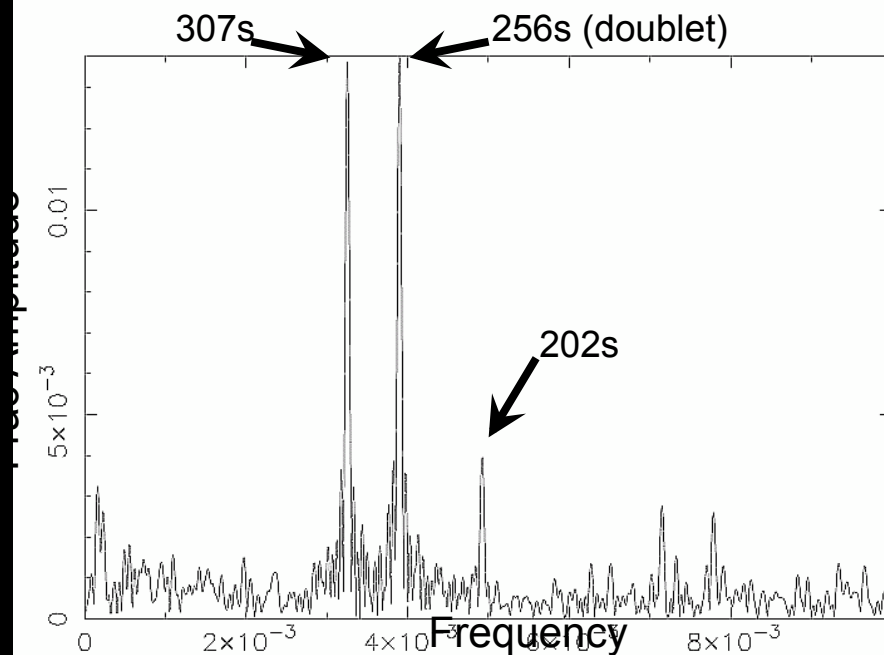
← Most commonly
observed
modes are $l=1$



$l = 1$ modes



fergal 10-Oct-2005 12:02



fergal 25-Apr-2005 17:30

- All (?) DAs become variable at $\approx 12,000\text{K}$
- Non-radial multi-periodic *g*-mode pulsations
- Periods between 100-1000s, amplitudes $\sim 1\%$

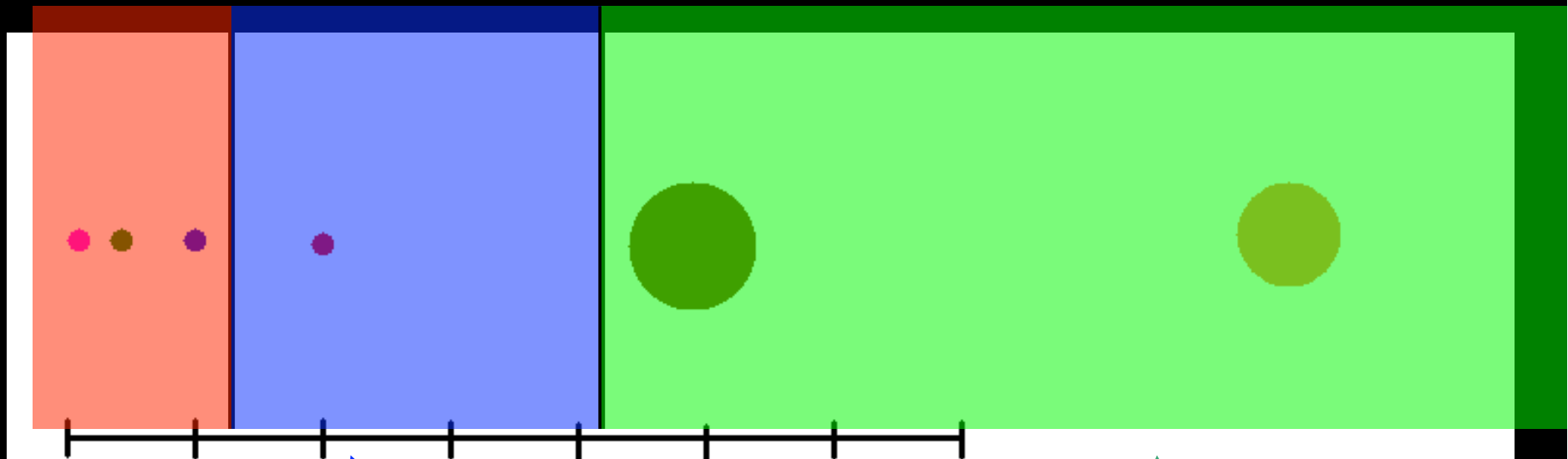
Extreme Stability Pulsators: **DBV** and **HDAV Stars**

- Stellar evolution becomes a spectator sport!
- Extreme Physics: Plasmons and axions
- The most stable optical clocks—only a few millisecond pulsars are more stable (or are they ...?)

Motivation: Why Look for Planets Around White Dwarf Stars?

- Clocks in orbit => information on companions. Watch for change in light travel time due to reflex orbital motion.
- Planet search samples a unique range of planet masses and separations, similar to our own Solar System
- $\sim 10^{-3}$ fainter than progenitor, makes follow up by direct observation possible (Mullally et al. 2007)
- They come with age estimates ...



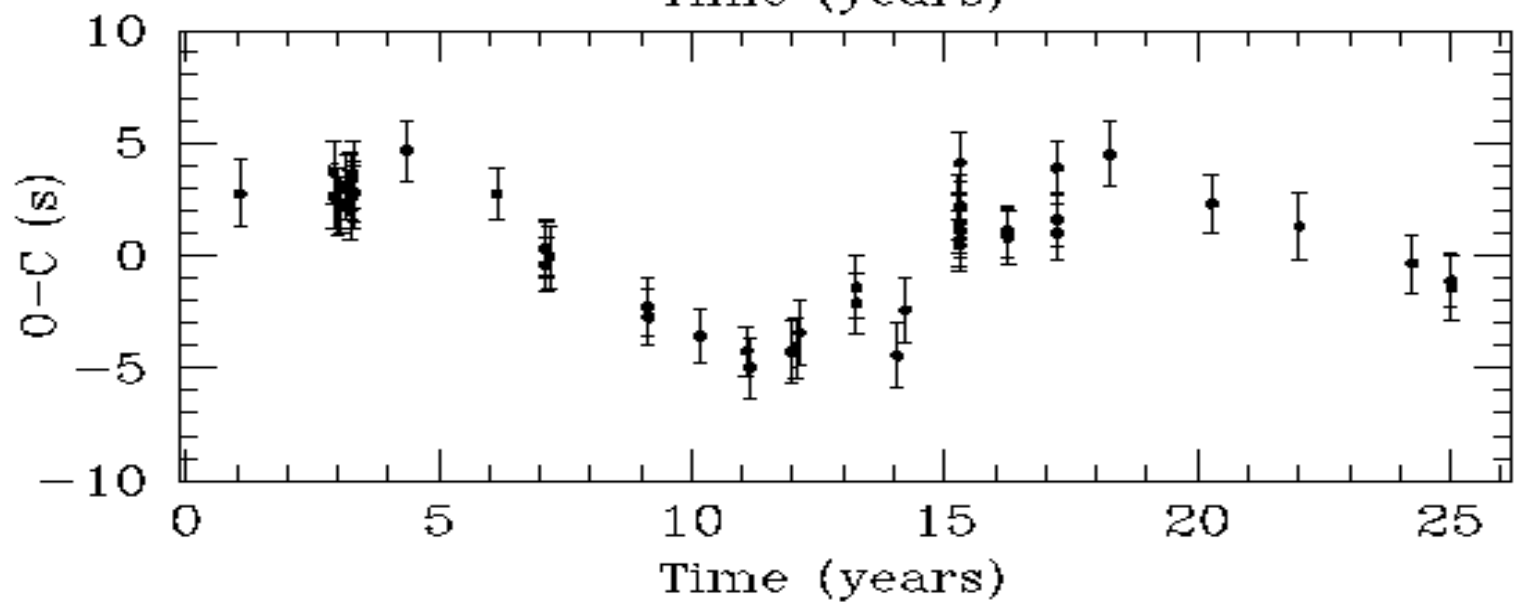
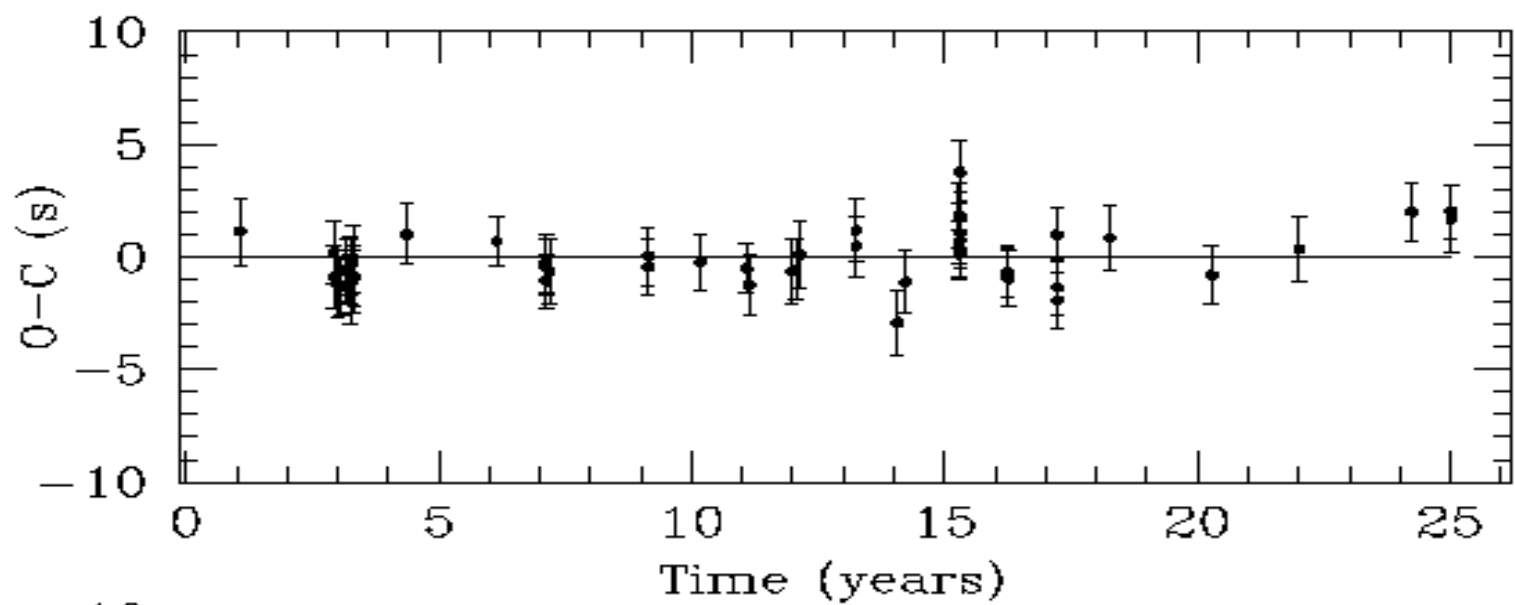


Destroyed In Red
Giant Envelope

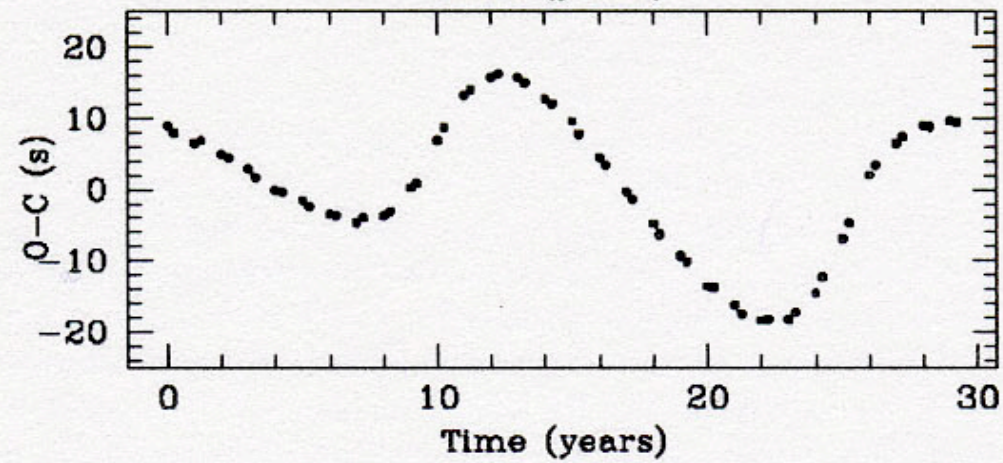
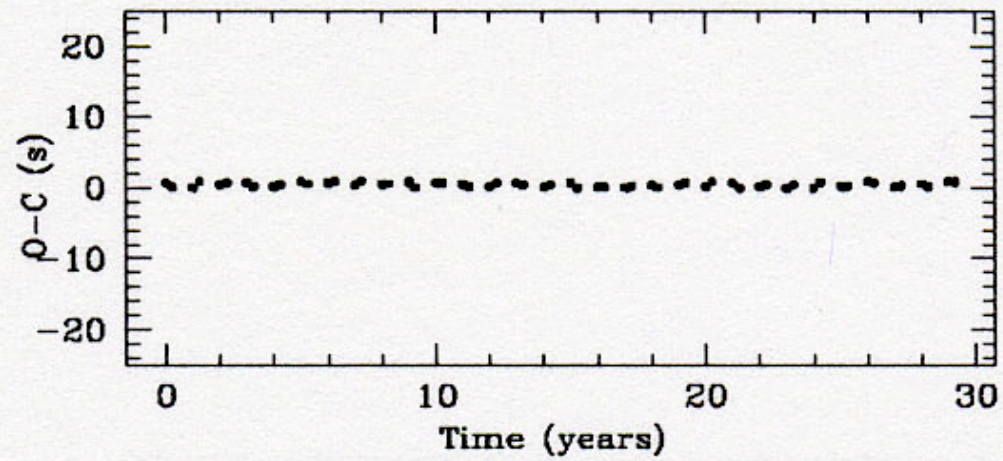
Spiral in to become
close companion (CV
progenitor?)

Move out adiabatically
with mass loss

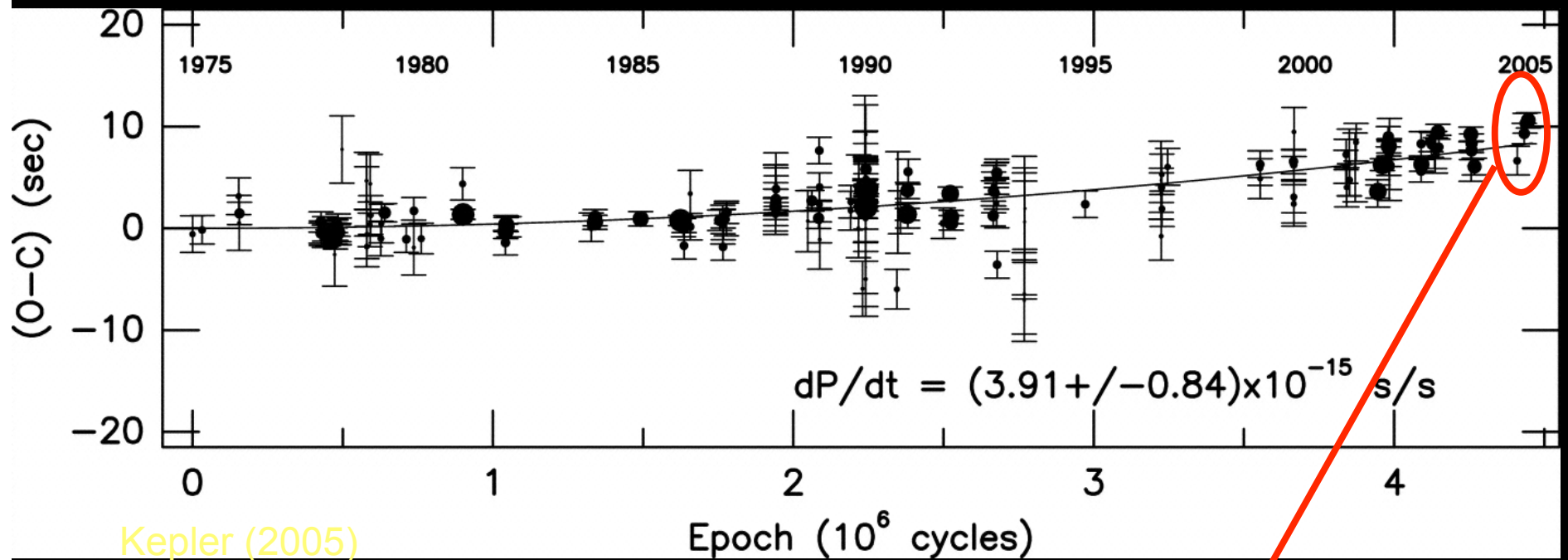




Multiple Planets



DAV pulsations are very stable

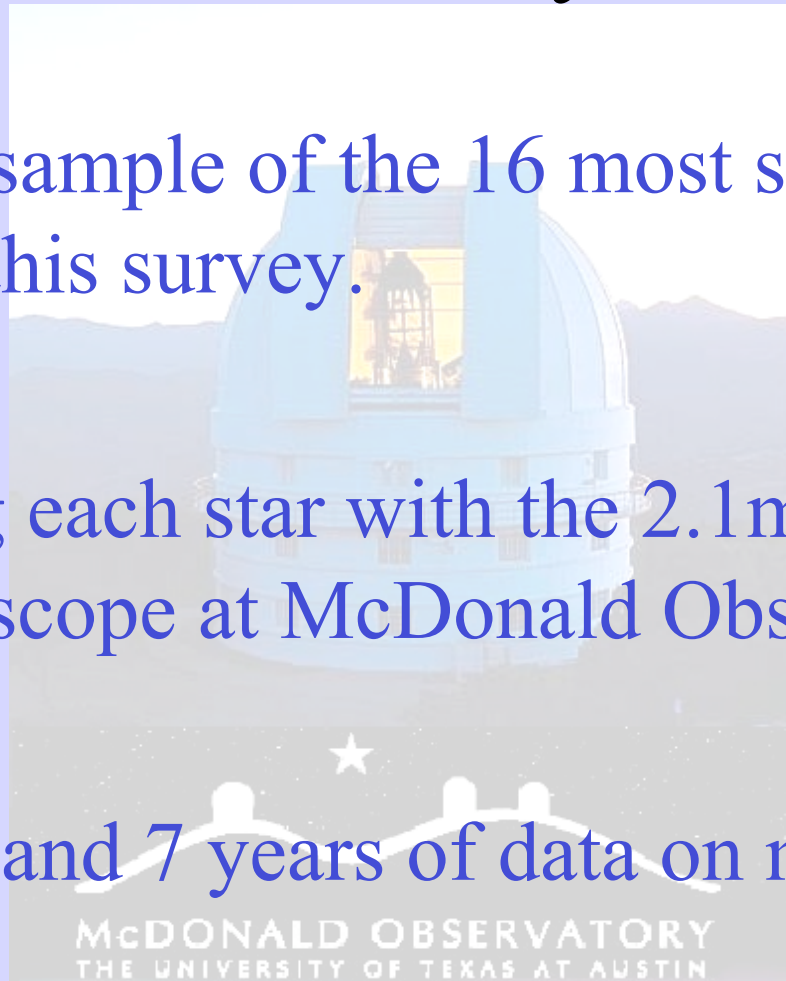


Kepler (2005)

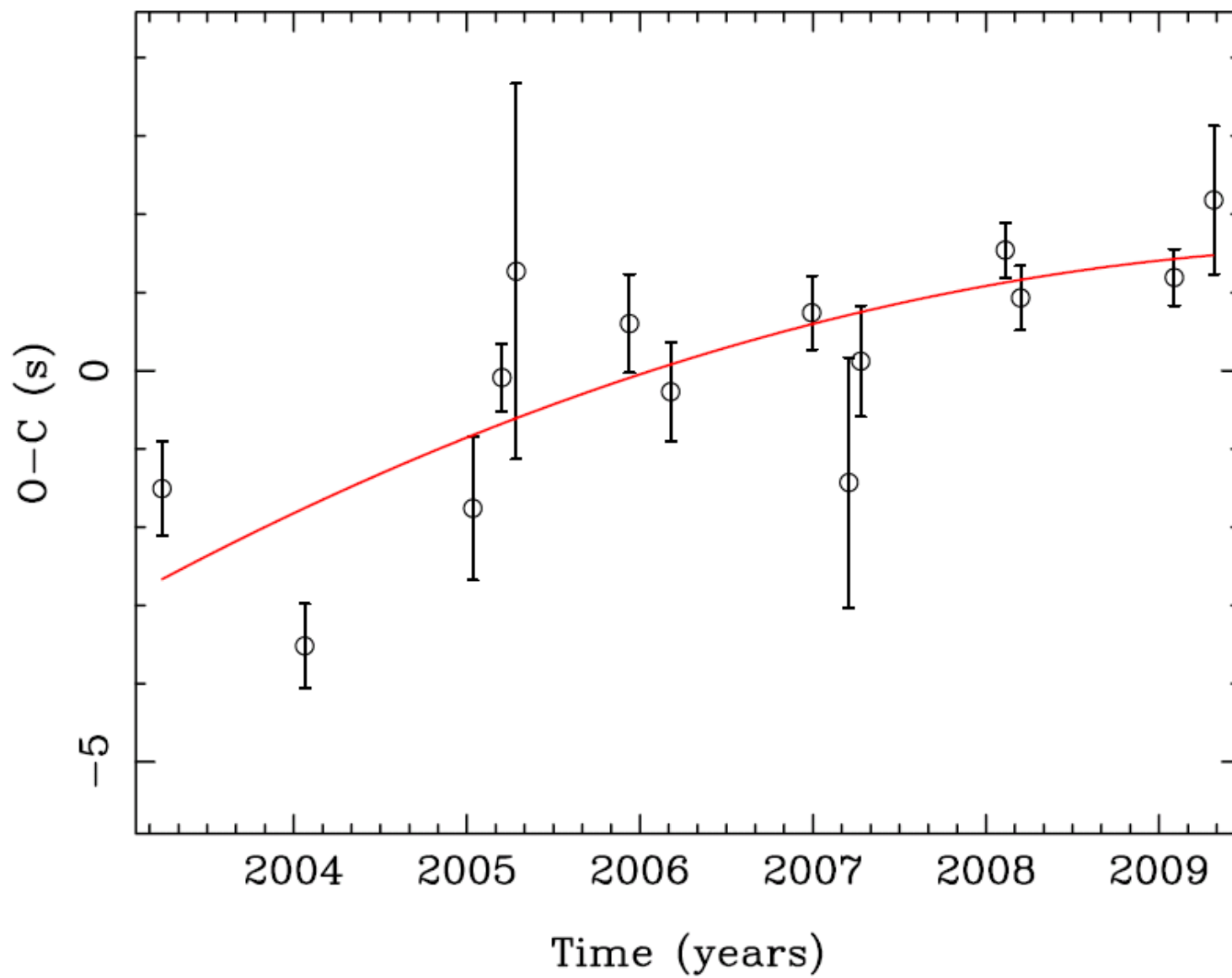
Argos Data

Our Survey

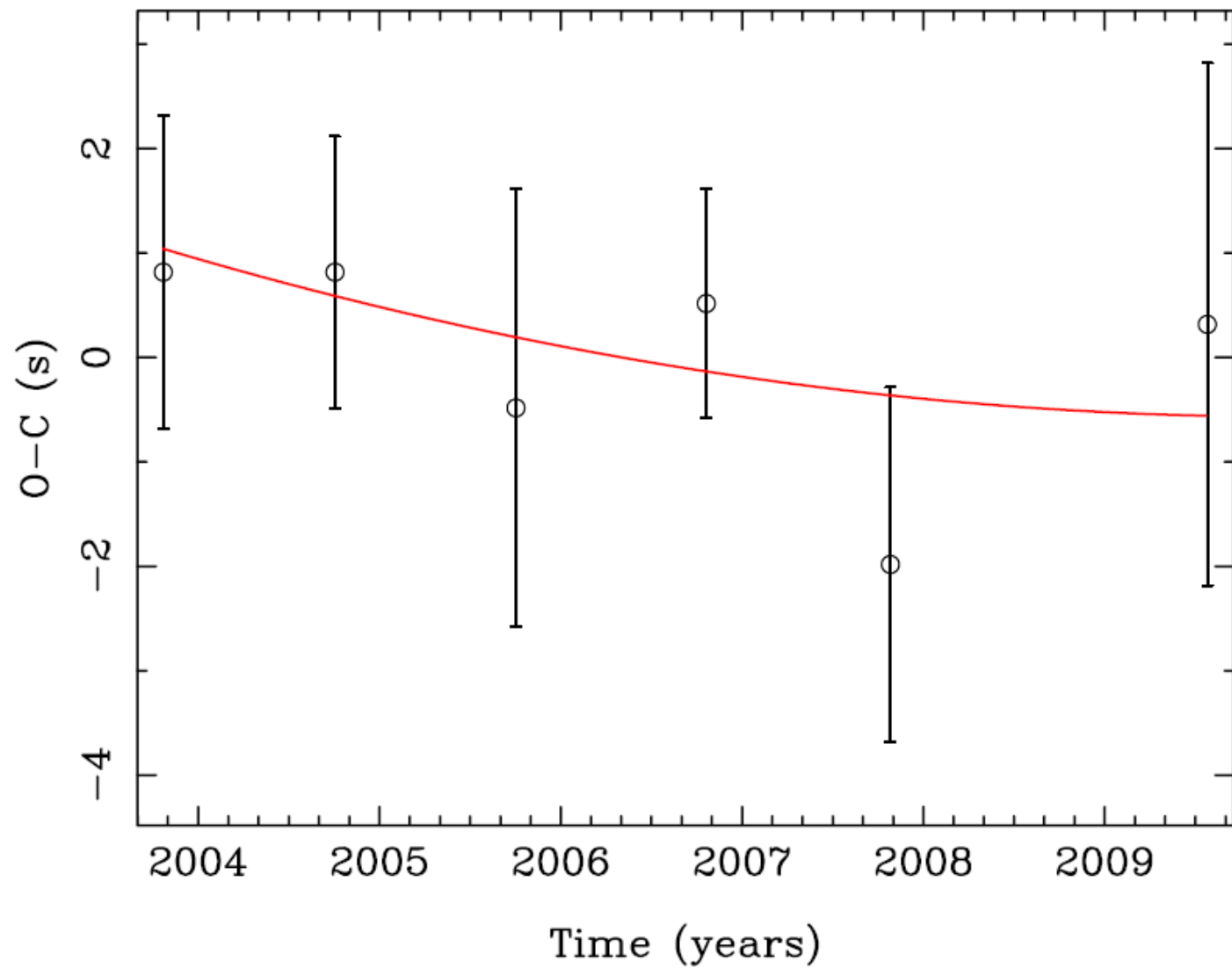
- Selected a sample of the 16 most suitable DAVs for this survey.
- Monitoring each star with the 2.1m Otto Struve telescope at McDonald Observatory
- Between 5 and 7 years of data on most stars.



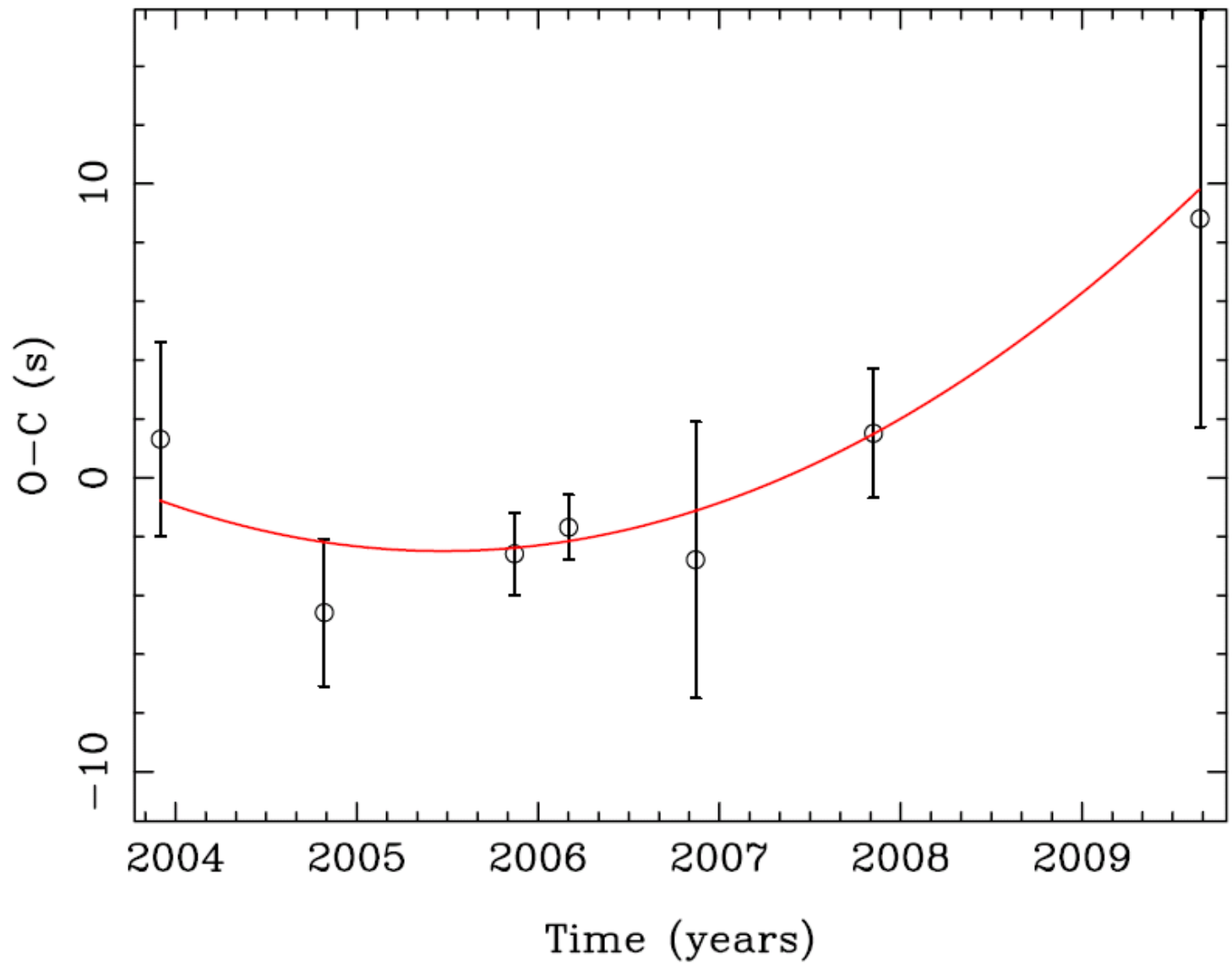
G117-B15A_215s_period



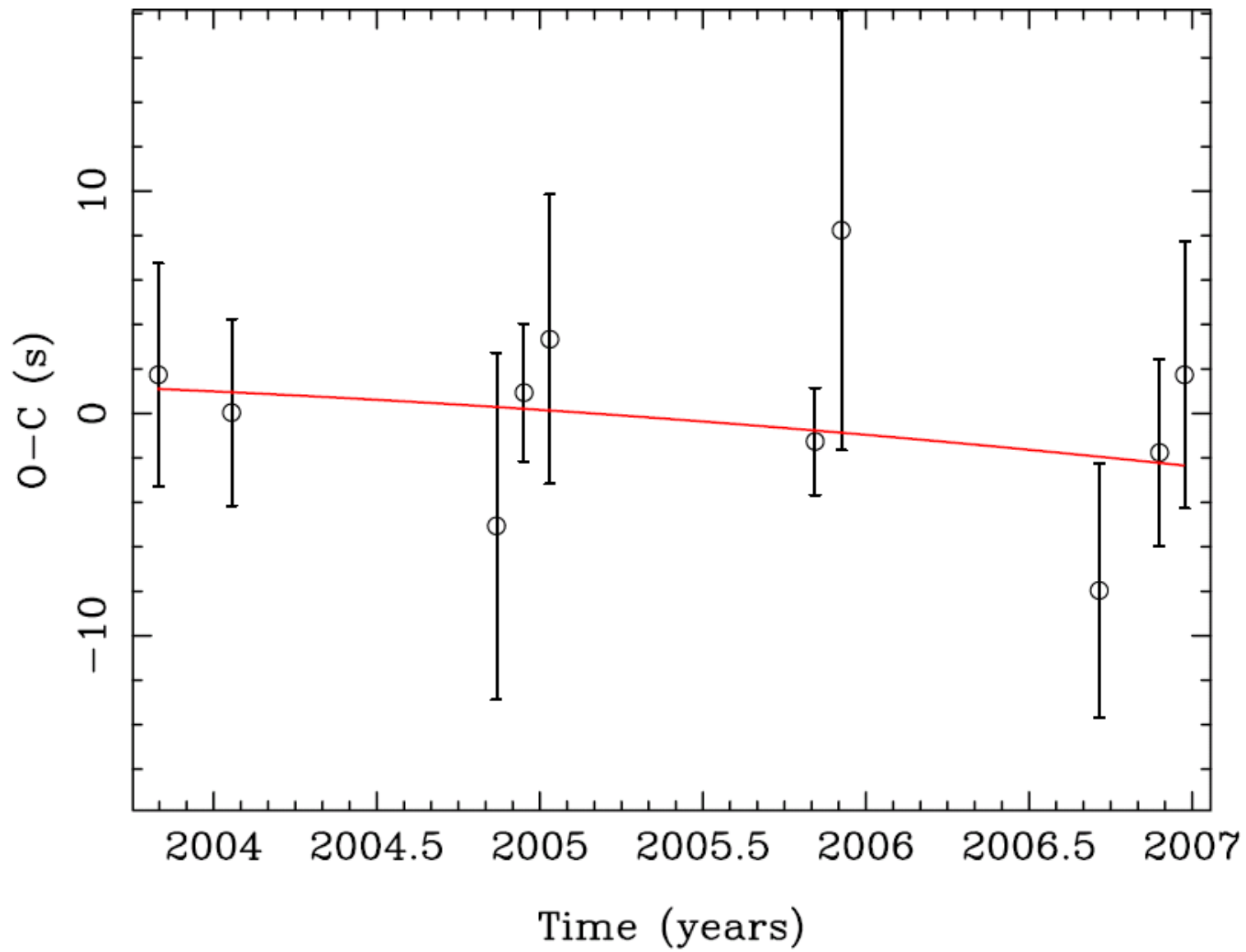
GD244_203s_period



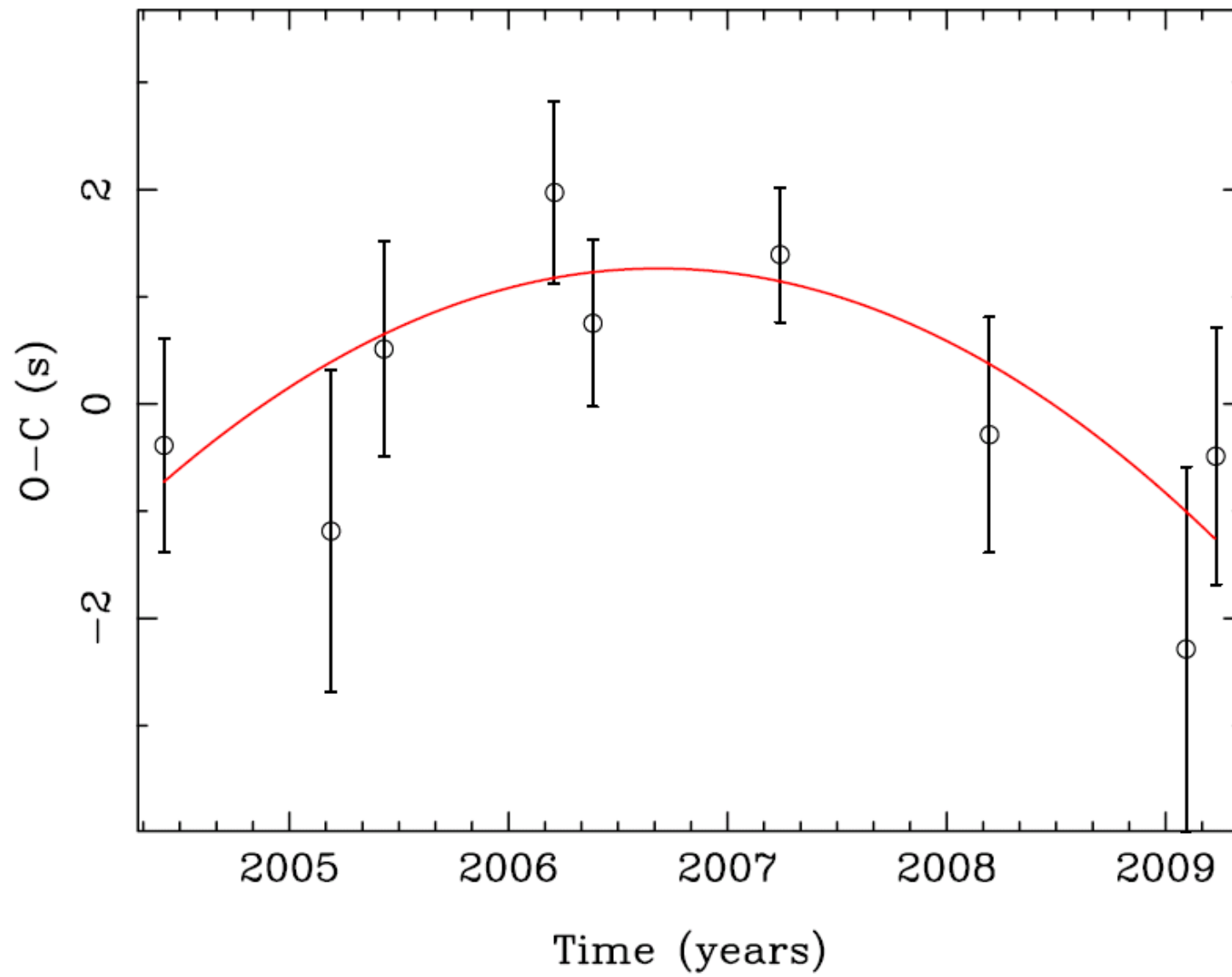
WD0018+0031_258s_period



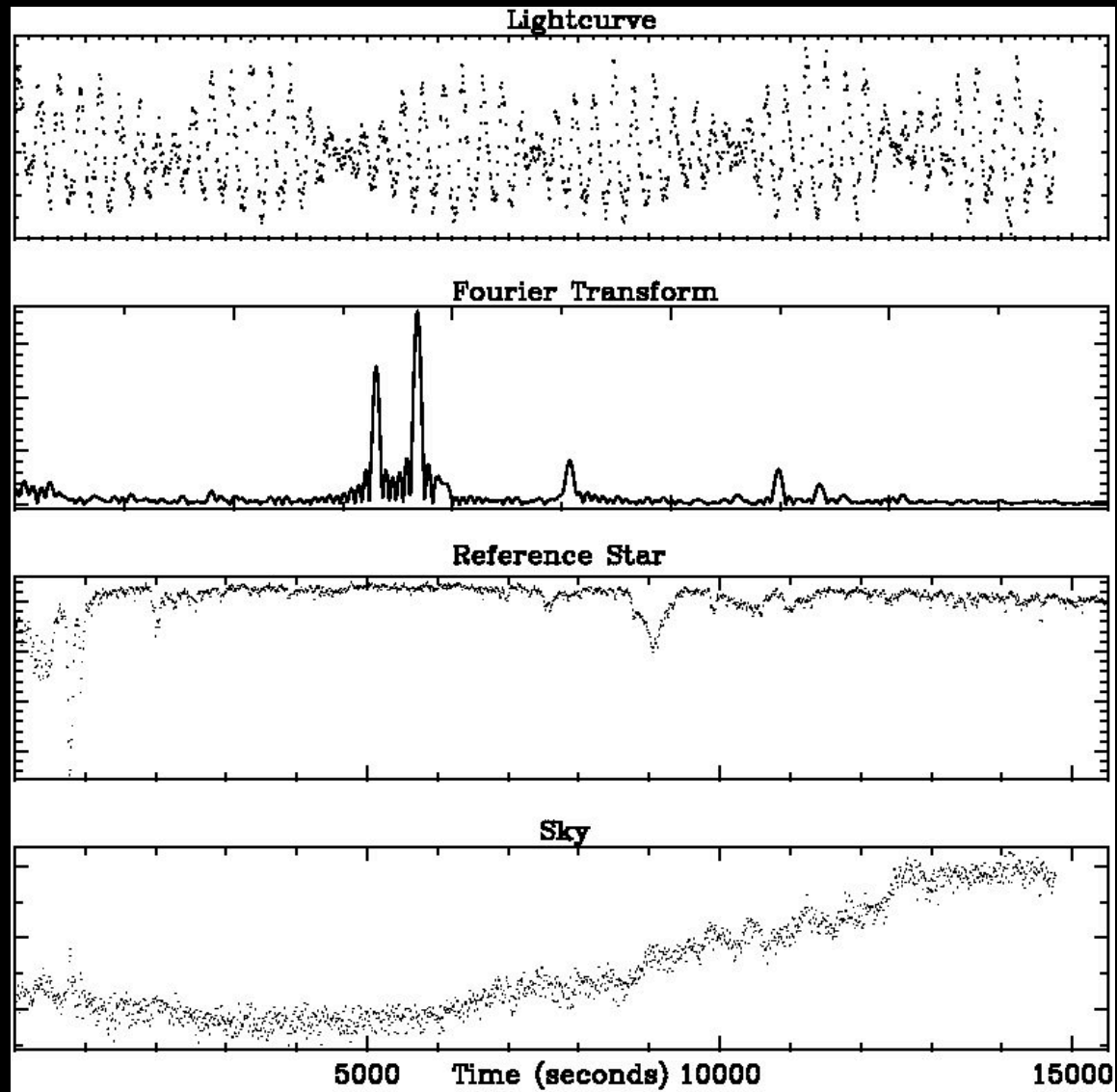
WD0214-0823_262s_period

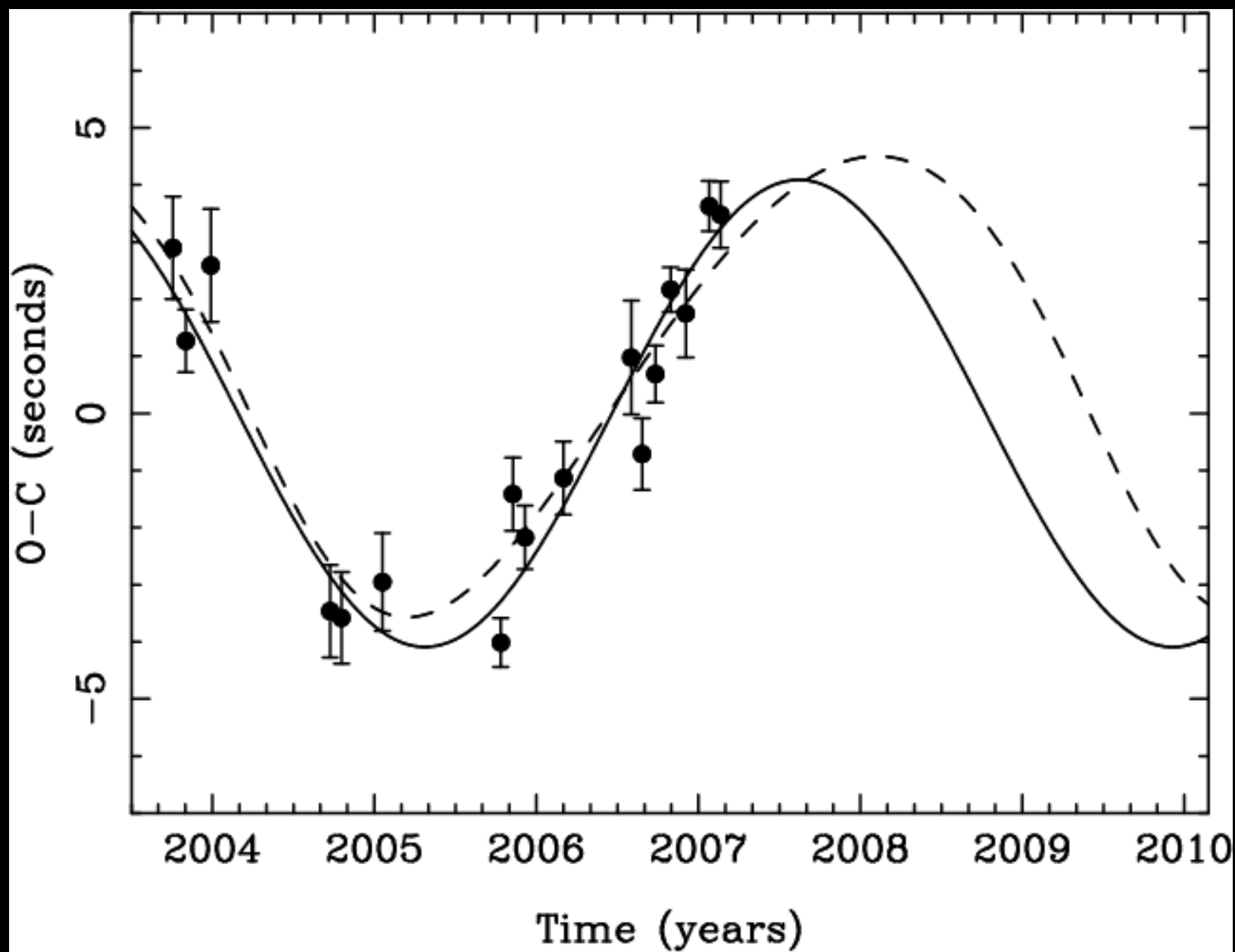


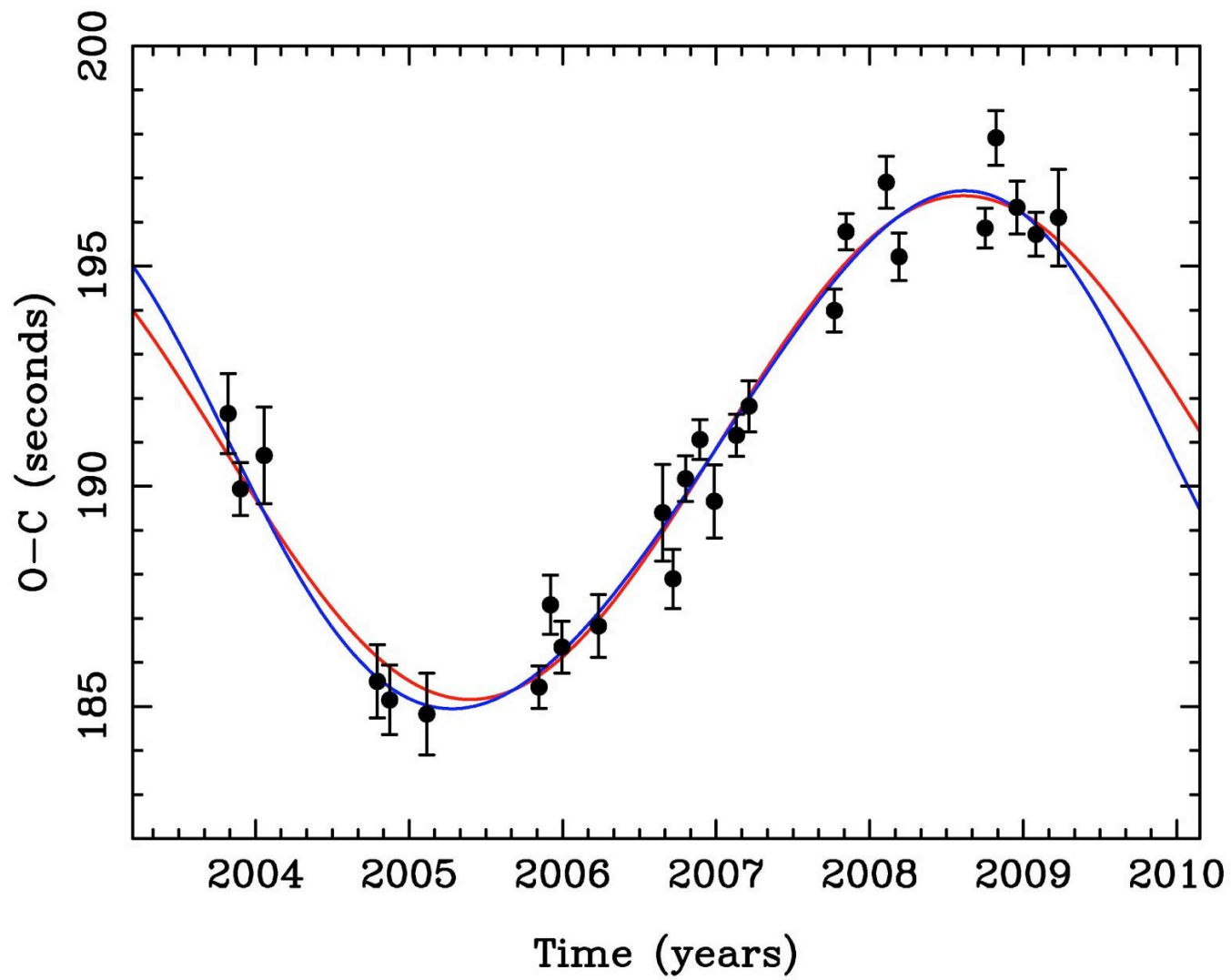
WD1354+0108_198s_period



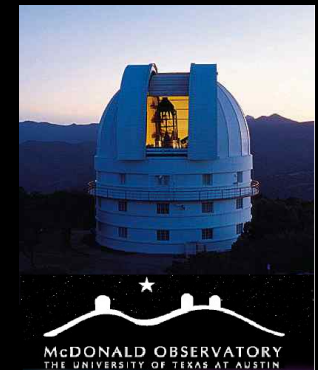
GD66 Lightcurve



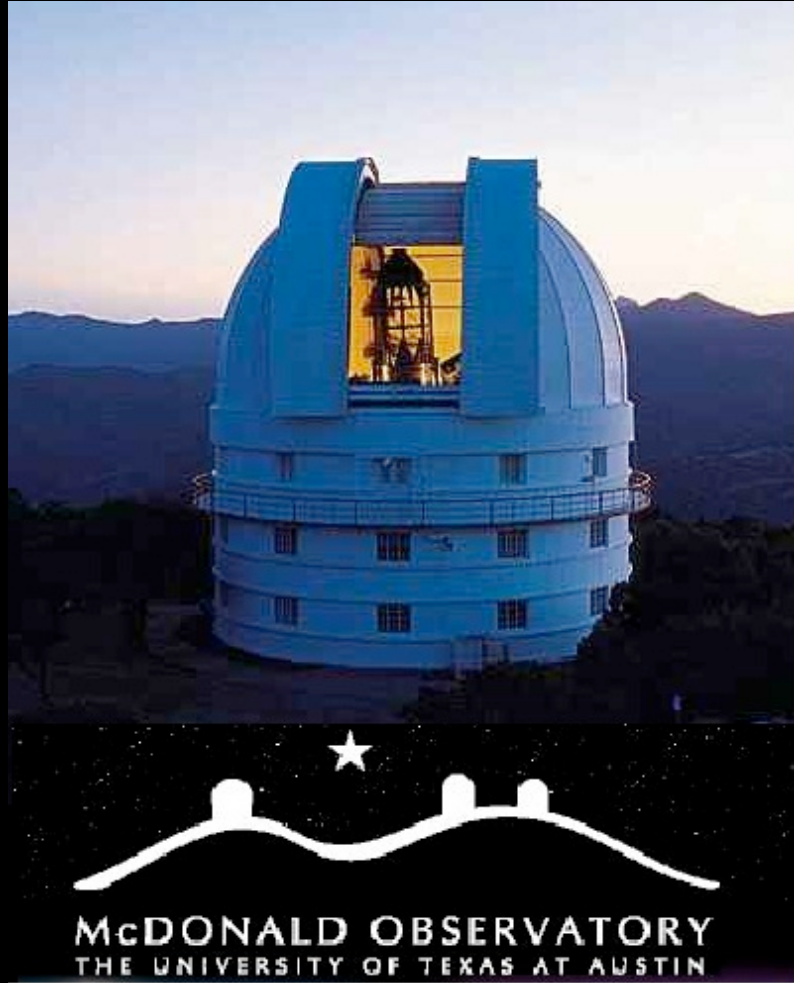




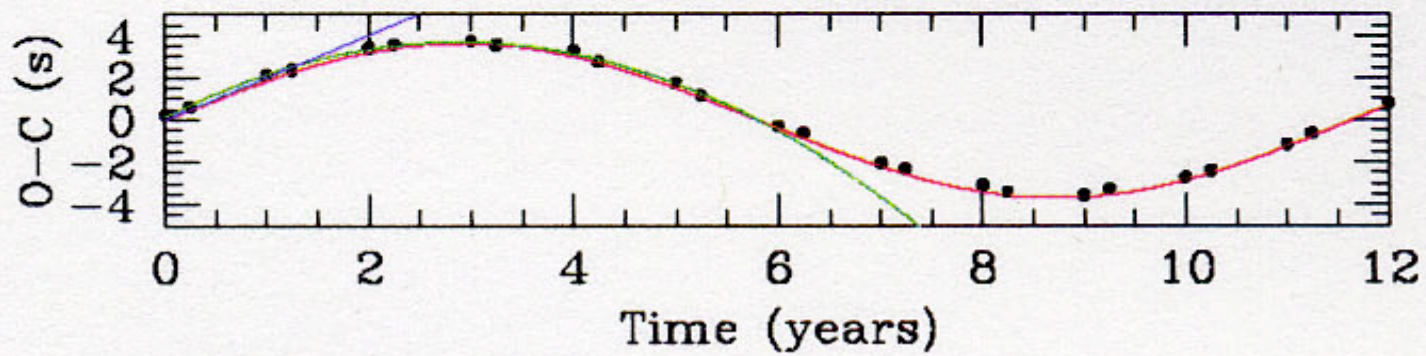
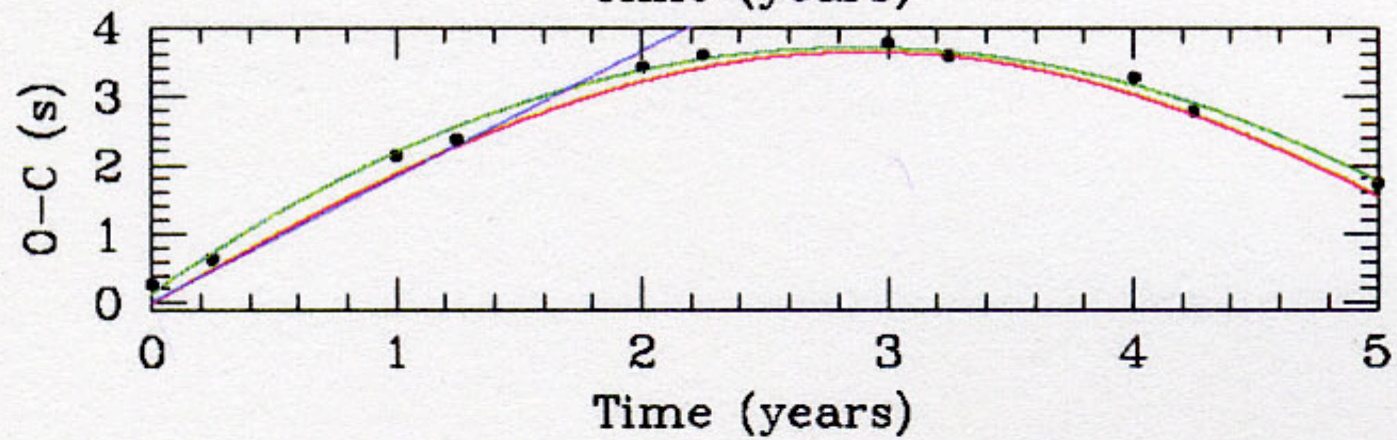
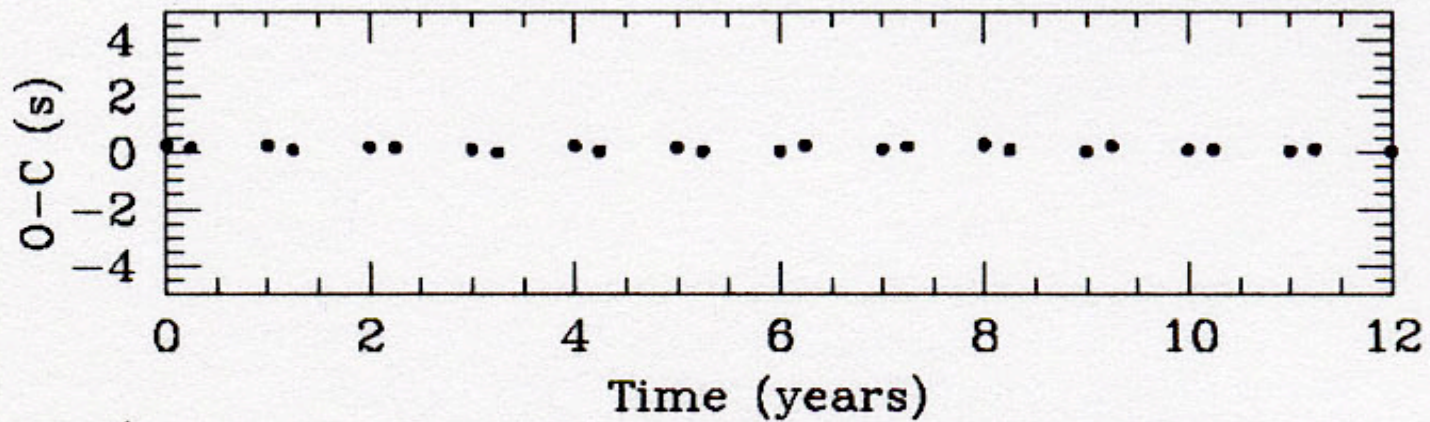
Current Results of Planet Search

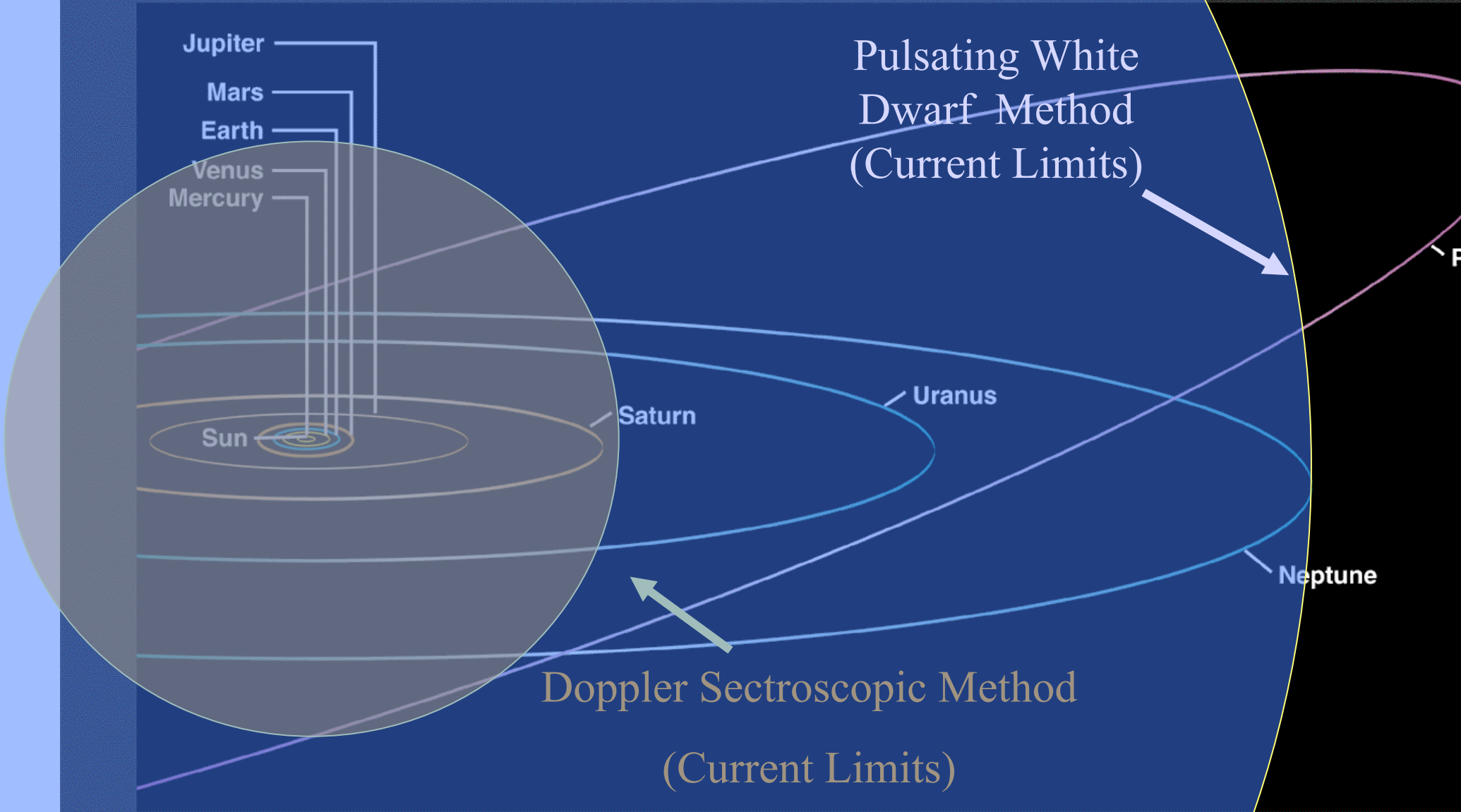


- We have evidence for *the first planet in orbit around a white dwarf star!*
- The planet mass is about $2 M_J$. The period is about 7 years.
- We have found a planet in a system that is *dynamically similar to our own!*
- Evidence for 1-2 more in sample of 16 stars:
3/16!



Early warnings

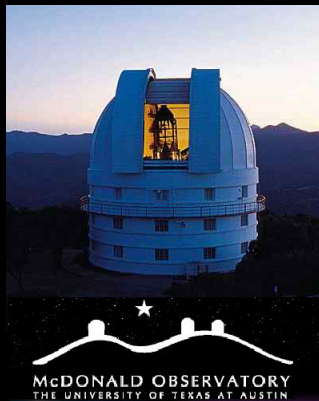




Other Explanations

Artifact of Data Reduction

Intrinsic Behavior of star



Instrument Clock

Fitting Procedure

Conversion between timing systems

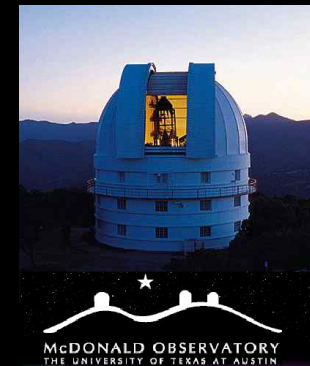
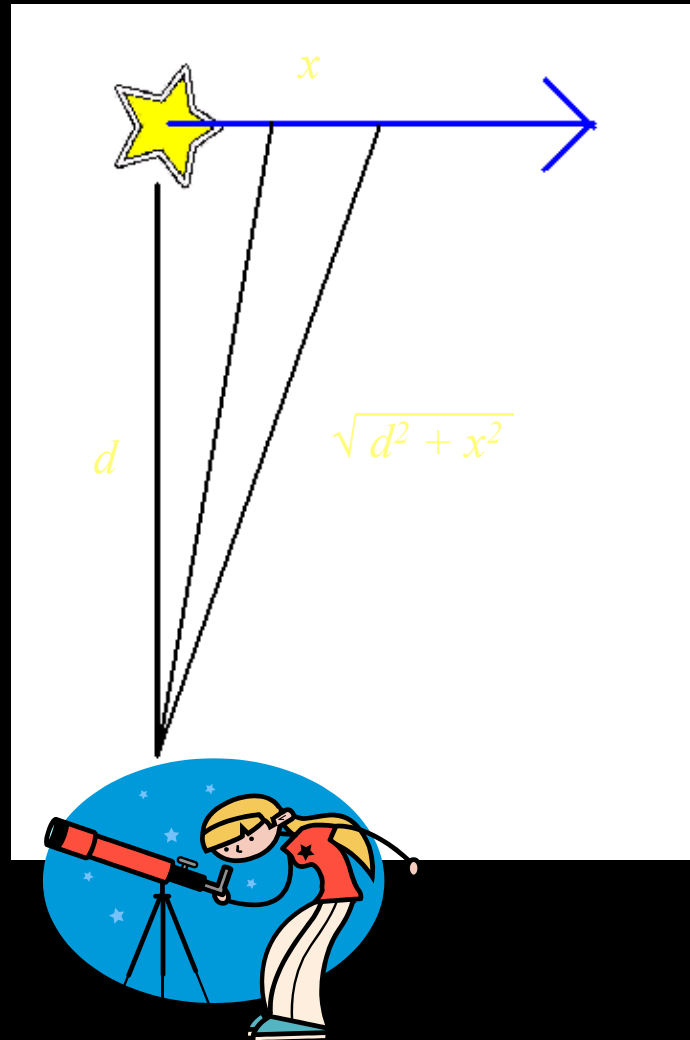
Stellar Cooling

Known DAV weirdness

Proper Motion

Core cry

Proper motion creates a \dot{P} term



Other Effects – Proper Motion



- $\mu = 133 \text{ mas/yr}$

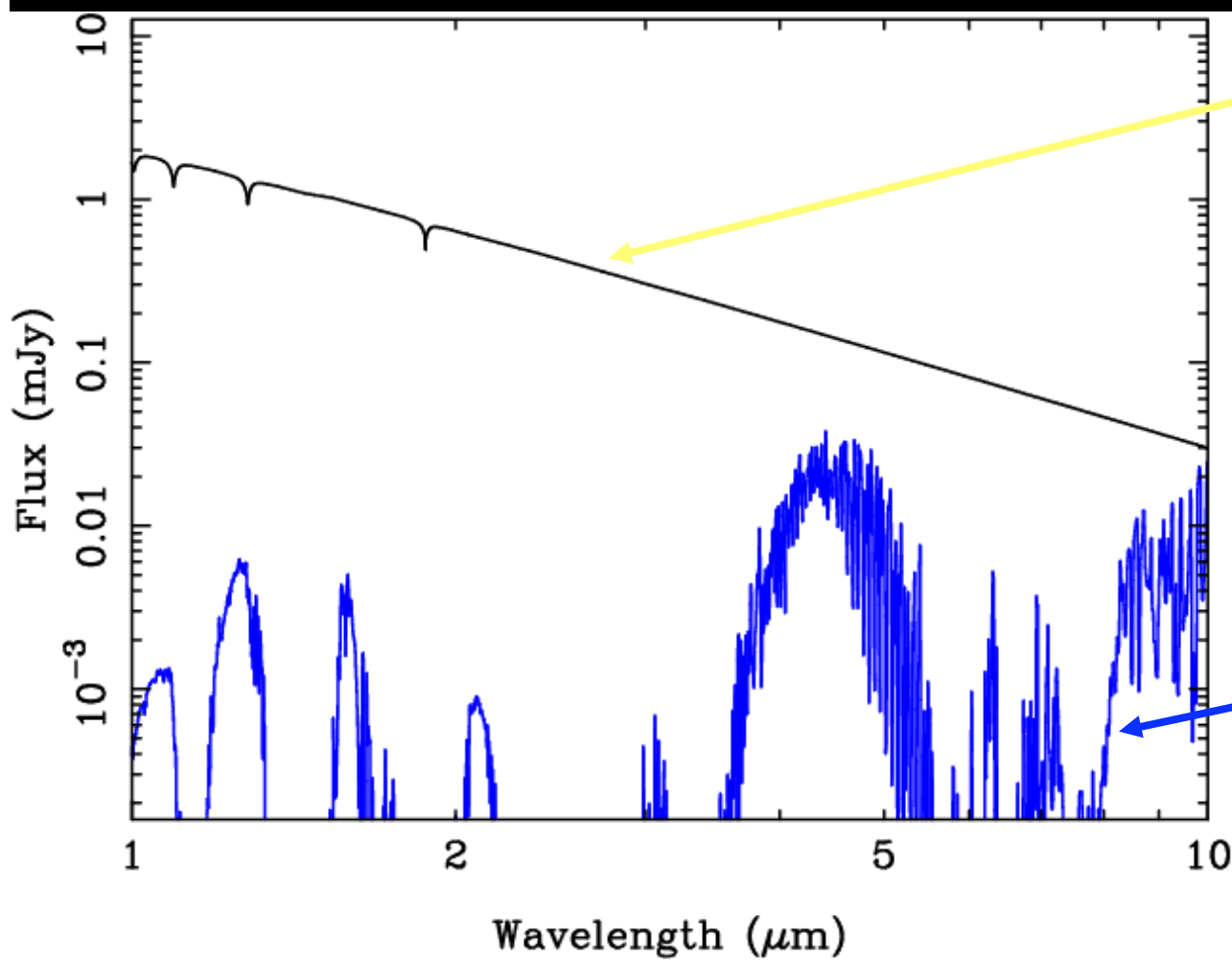
- $\dot{P}_{\text{pm}} = 2.430 \times 10^{-18} P \mu^2 d,$

Pajdosz (1995)

- $P = 302 \text{ s}$ $d \approx 50 \text{ pc}$

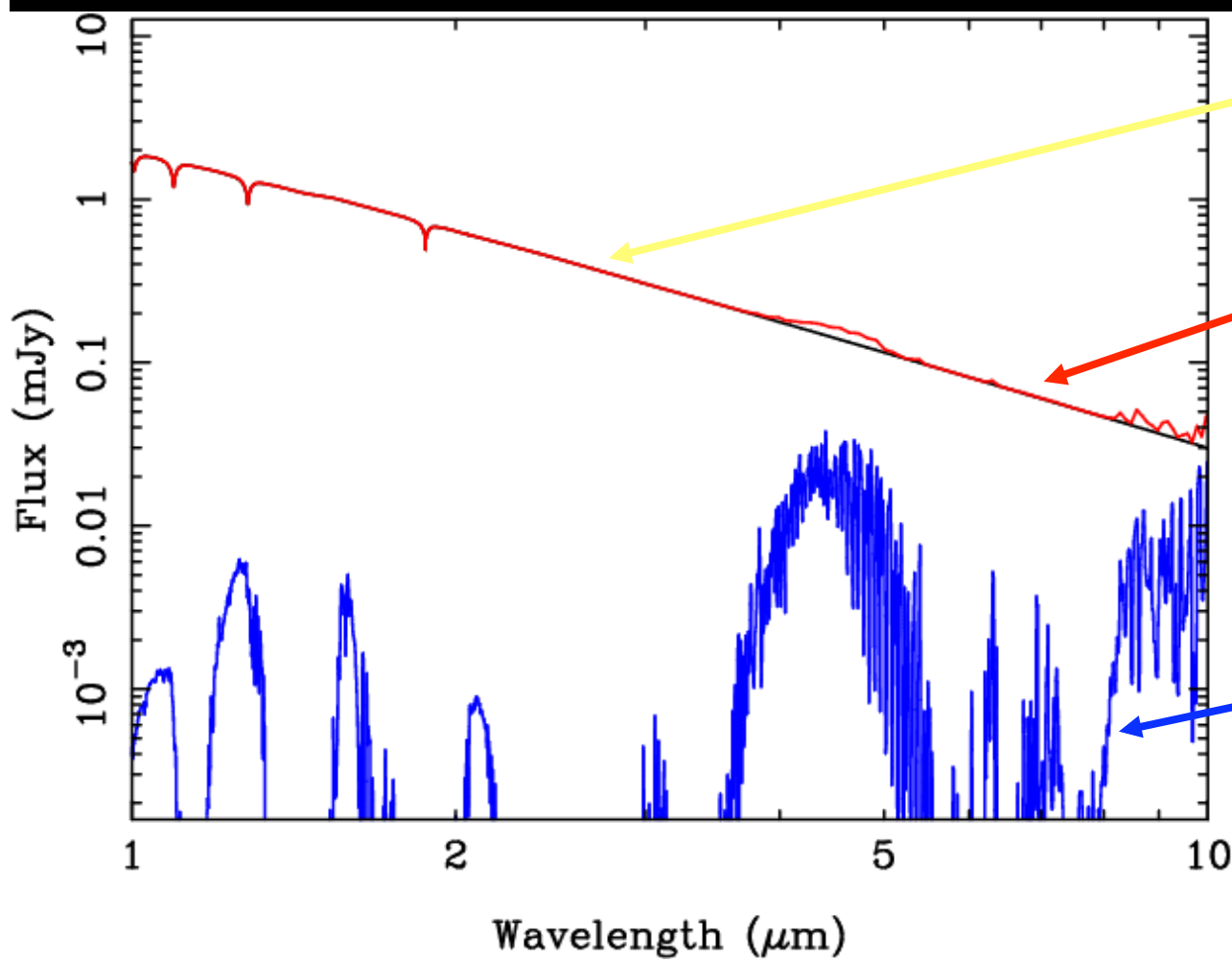
- $\dot{P}_{\text{pm}} = 6.5 \times 10^{-16}$

- $\dot{P}_{\text{obs}} \sim 10^{-12}$



12kK DA White Dwarf model spectrum

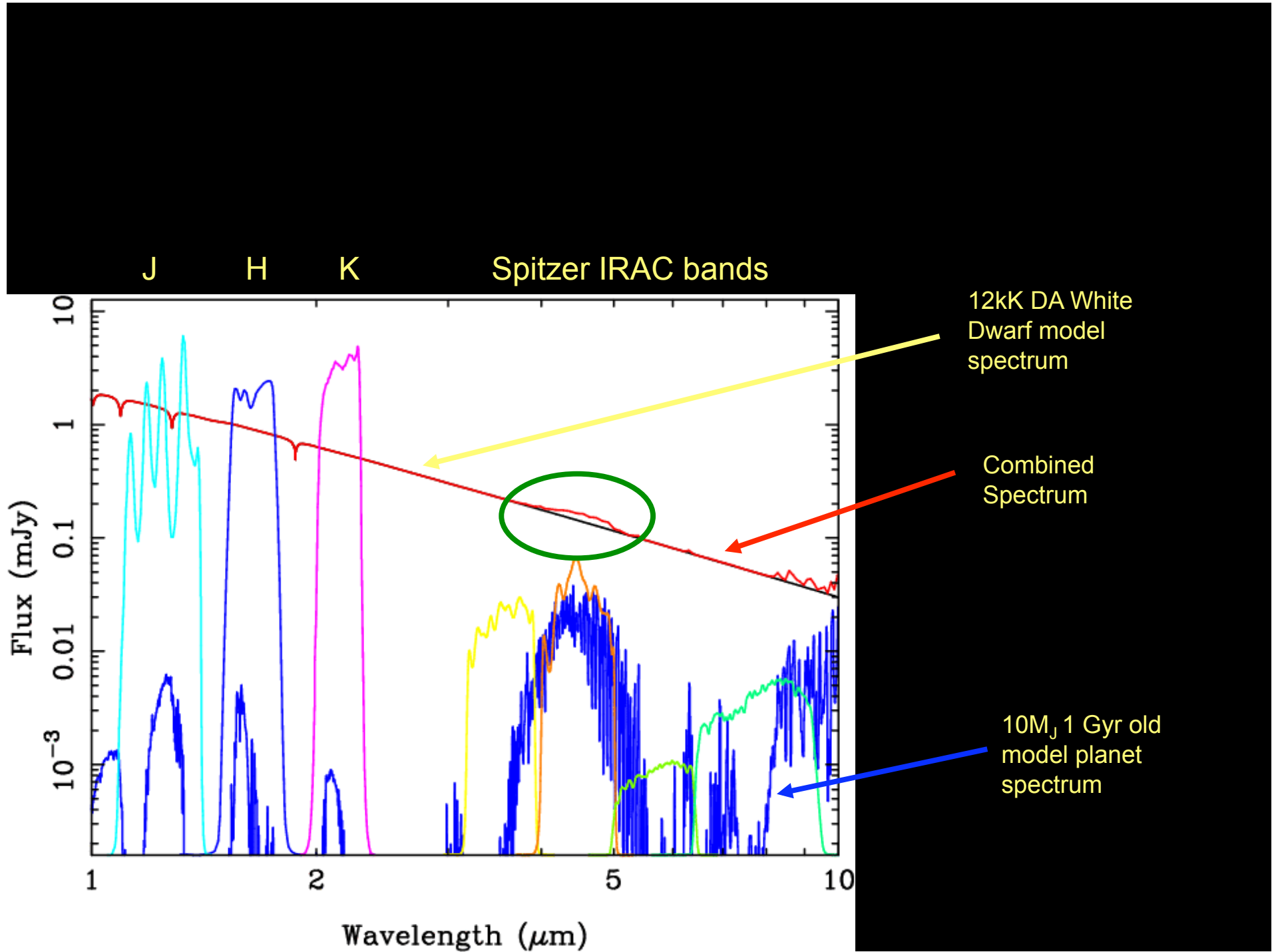
$10M_j$ 1 Gyr old model planet spectrum

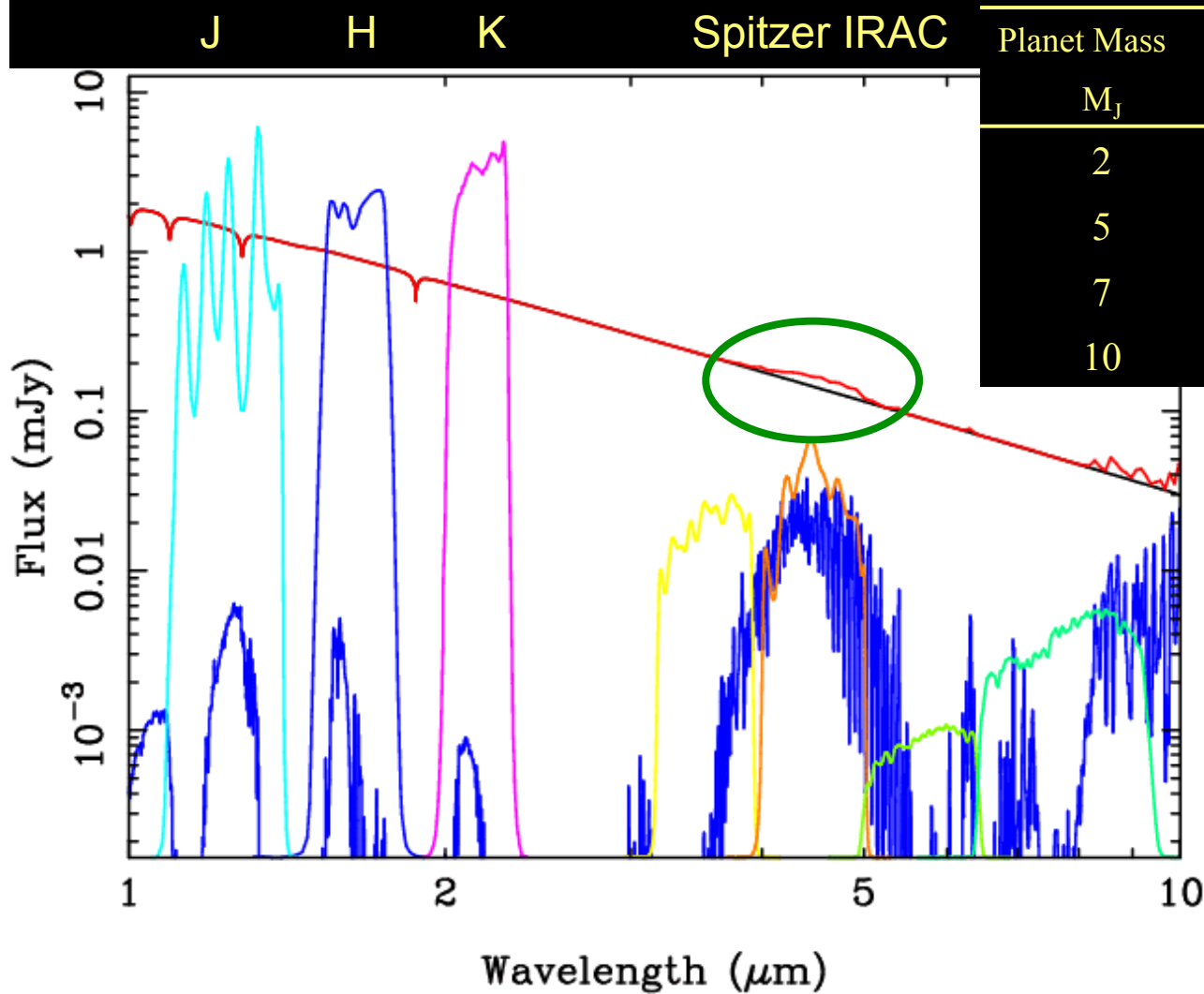


12kK DA White Dwarf model spectrum

Combined Spectrum

$10 M_J$ 1 Gyr old model planet spectrum





Planet Mass M_J	Contrast at $4.5\mu\text{m}$	Observing Time for 1σ detection
2	1.003	6.8hrs
5	1.043	111s
7	1.078	33s
10	1.127	13s

$$\frac{\text{Flux}(\text{Star} + \text{Planet})}{\text{Flux}(\text{Star})}$$

