Comets: What do they tell us about the early Solar System

Dr. Anita Cochran The University of Texas McDonald Observatory

Photo: Ed Szczepanski

Why Study Comets?

- Comets are the least altered objects left over from the formation of the Solar System and were the building blocks of planets in the outer part of the Solar System.
- Comets we see today in the inner Solar System come from one of two reservoirs:
 - The Kuiper belt a disk-like region located from Neptune outwards and mostly primordial
 - The Oort cloud a spherical halo of comets out to about 10⁵ AU; these objects were formed in the region of the giant planets



Composition

- The spectrum of a comet is mostly molecular gas emission, with a continuum formed from sunlight reflecting off the dust.
- Most of the spectral lines are formed via resonance fluorescence.
- Comets have high <u>excitation</u> temperature but not high physical temperatures.
- The presence of lines and their ratios <u>do not</u> indicate the density of the material. They are an indicator of the parent species



Relative Intensity

Are All Comets Chemically Similar?

- The vast majority of comets have similar composition
- There is a sub-group (generally Jupiter family comets) which have <u>different</u> spectral characteristics.



Measuring Isotope Ratios

- Isotope ratios yield information on the fractionation history of the gas.
- This is a function of temperature, density and pressure.
- We need to know the isotope ratios as a <u>function of</u> <u>formational distance</u>.



The B-X band of CN is a $\Sigma - \Sigma$ transition so it has no Q-branch. The strongest lines are of the (0,0) band. There is also evidence of the Pand R-branch lines of the (1,1) band within the (0,0) Rbranch lines, with a bandhead at 3871Å.

The isotopes are determined with synthetic spectra to model the isotopomers



Note the isotope shifts and different relative intensities.

Examples



¹²C/¹³C: Howell 90±10 ; de Vico 90±10

¹⁴N/¹⁵N: Howell 140±20; de Vico 140±15



Halley Type comets are denoted with **green circles**; Oort cloud comets are denoted with **blue squares**; Jupiter family comets are denoted with **red triangles**.

Work of A. Cochran (Texas), E. Jehin (ESO), J. Manfroid, D. Hutsemakers, C. Arpigny (Liege), J.-M. Zucconi (Besancon), J. Stuwe (Leiden) and R. Schulz (ESA) – various publications

Common Values for the Ratios

- C and N isotope ratios measured in 13 comets
 - 2 Halley Type comets, 2 Jupiter Family comets and 9 Oort Cloud comets
- All of these comets have the same ¹²C/¹³C ratio
 - ¹²C/¹³C= 91±21
 - The same as the Solar System and terrestrial value
- All of these comets also have the same ¹⁴N/¹⁵N ratio.
 - ¹⁴N/¹⁵N= 141±29
 - Solar System value <u>assumed</u> to be the terrestrial value of 272 (Anders and Grevasse 1989)
 - Our team was the first to measure this ratio in the optical

Nitrogen Isotopes Elsewhere



LMC = Large Magellanic Cloud HC¹⁵N [4]; ISM=Interstellar Medium HCN [6]; Solar wind from SOHO [14] and upper layers of lunar soils [16]; Jupiter from Galileo Probe [19] and Cassini/CIRS [1]; Comets CN, Mars[18], Titan Huygens GCMS [17], Sub-mm HCN [8], Venus [9]

Implications

- Owen et al. (2001) argue that the value of ¹⁴N/¹⁵N for Jupiter represents the solar value of this ratio since other isotope ratios (Carbon, Xenon) measured in Jupiter's atmosphere are the same as the Sun
 - The Local Interstellar Medium value (LISM, the ISM point at 8.2Kpc) agrees with the Jupiter value
 - The value measured for the solar wind from SOHO disagrees with the Jupiter value by a factor of two
- If the Solar value is ~400 but we measure ~140 from CN
 - Must be an additional parent for CN
 - That parent must have a ¹⁴N/¹⁵N much lower than the normal solar nebula value
 - Might point to large scale inhomogeneity in the solar nebula





An Opportunity to Determine the Role of Primordial Cometary Material in the Formation of Planets

