

Empirical Evidence for AGN Feedback

Christy Tremonti

MPIA (Heidelberg) / U. Wisconsin-Madison

Aleks Diamond-Stanic (U. Arizona), John Moustakas (NYU)

Much observational and theoretical evidence supports a basic evolutionary scenario

Merger of
gas-rich disks



Ultra-Luminous
Infrared Galaxy



Quasar

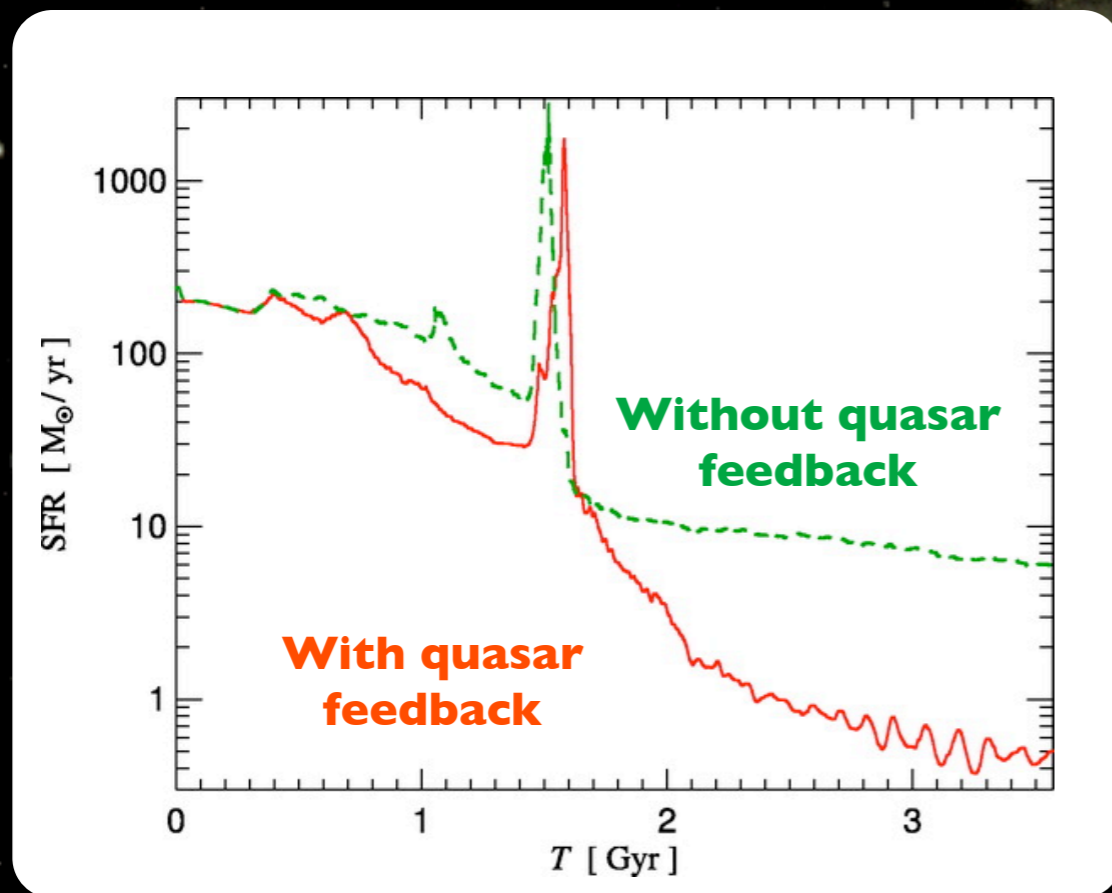


Massive
Early-type



Do quasars play an important role in this transition?

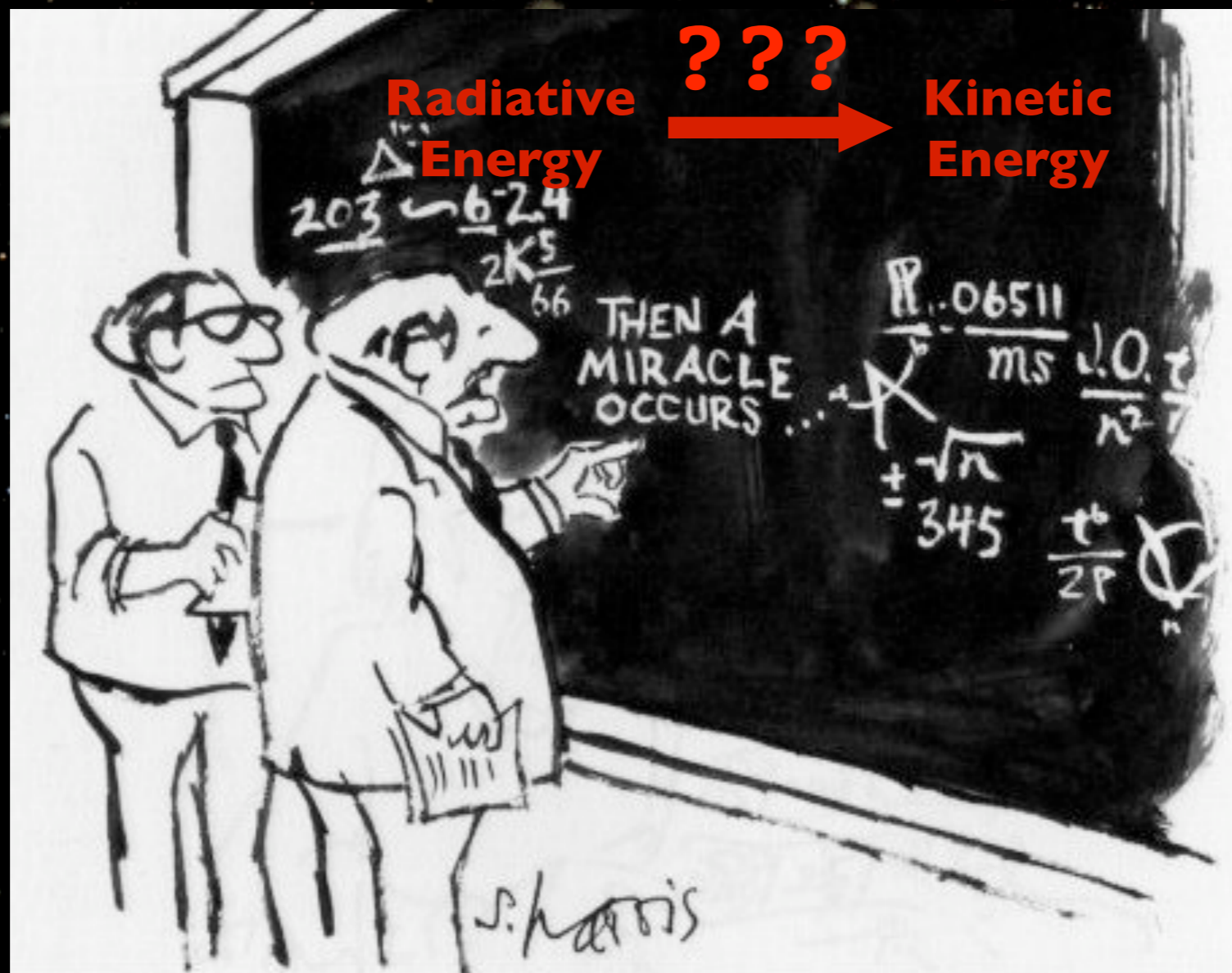
Theoretical models that include quasar-mode feedback shut-off star formation quickly following the merger



Springel et al. 2005



Why should we be skeptical?



"I think you should be more explicit here in step two."

from *What's so Funny about Science?* by Sidney Harris (1977)

We lack a good physical model... and we have limited observational evidence

We have abundant observational evidence for feedback from star formation

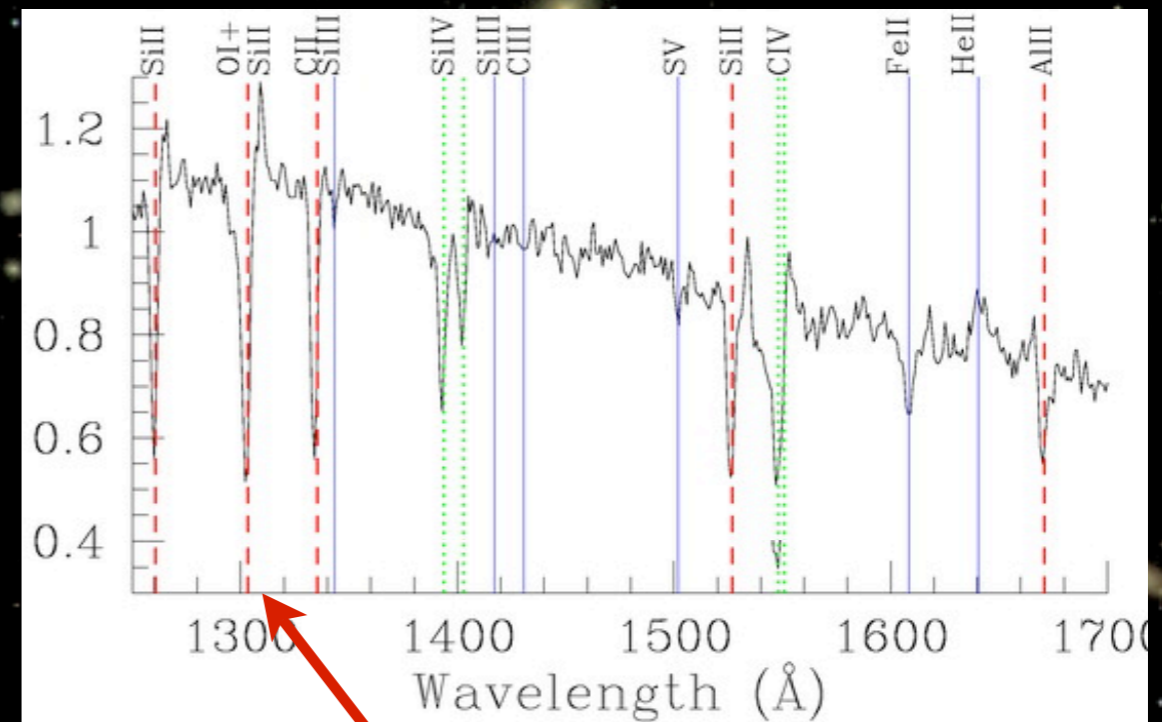
M82

Soft X-rays (Chandra)
R-band (HST)
PAHs (Spitzer)



Starburst driven outflows are usually detected via blueshifted absorption lines (from clumps of cold gas entrained in the hot wind)

Composite UV spectrum of local starbursts
Schwartz & Martin 2006



Blueshifted ISM lines

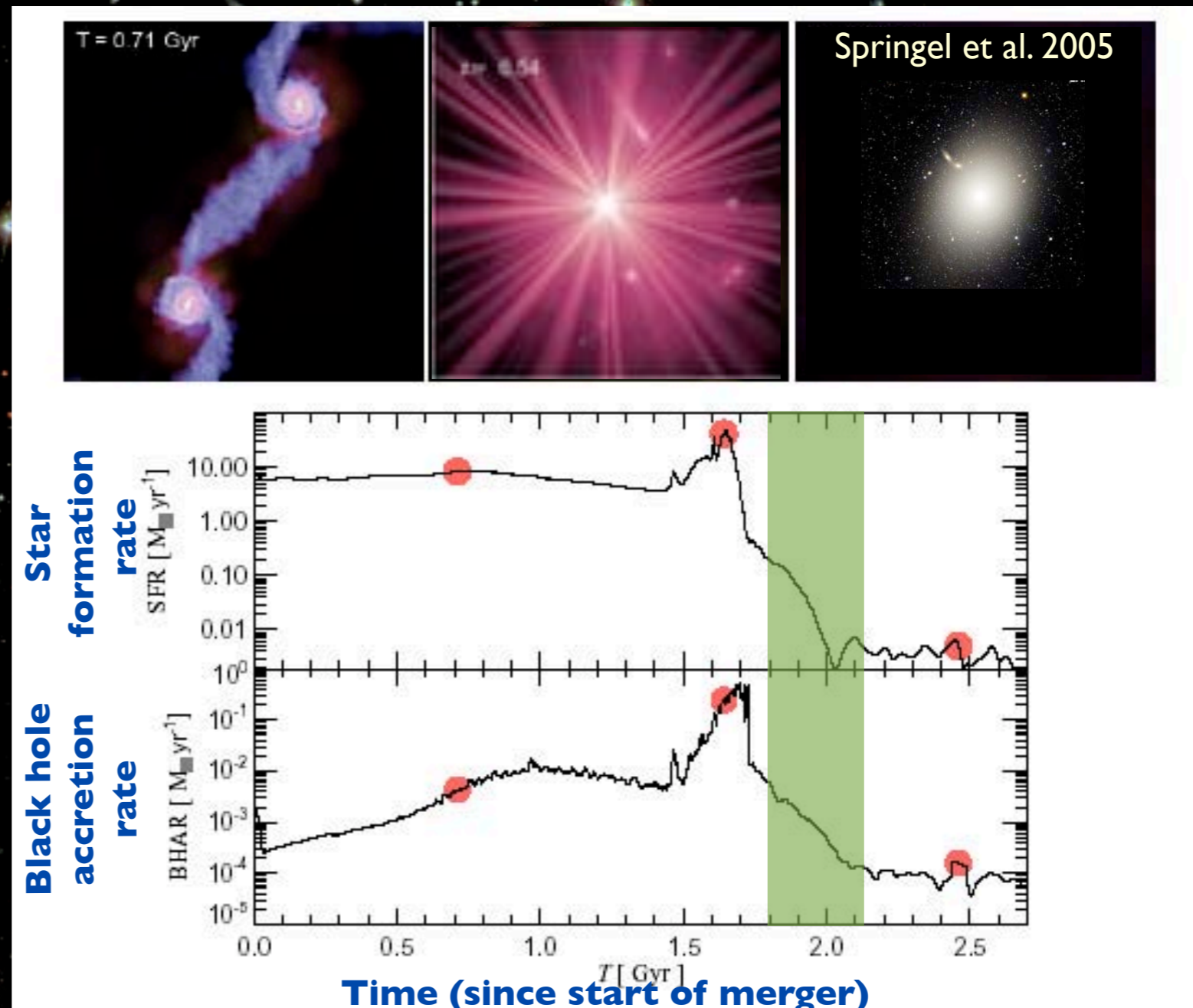


Can we detect quasar powered outflows the same way?

Feedback energy is maximum during the quasar phase, but this is not the best time to look for outflows

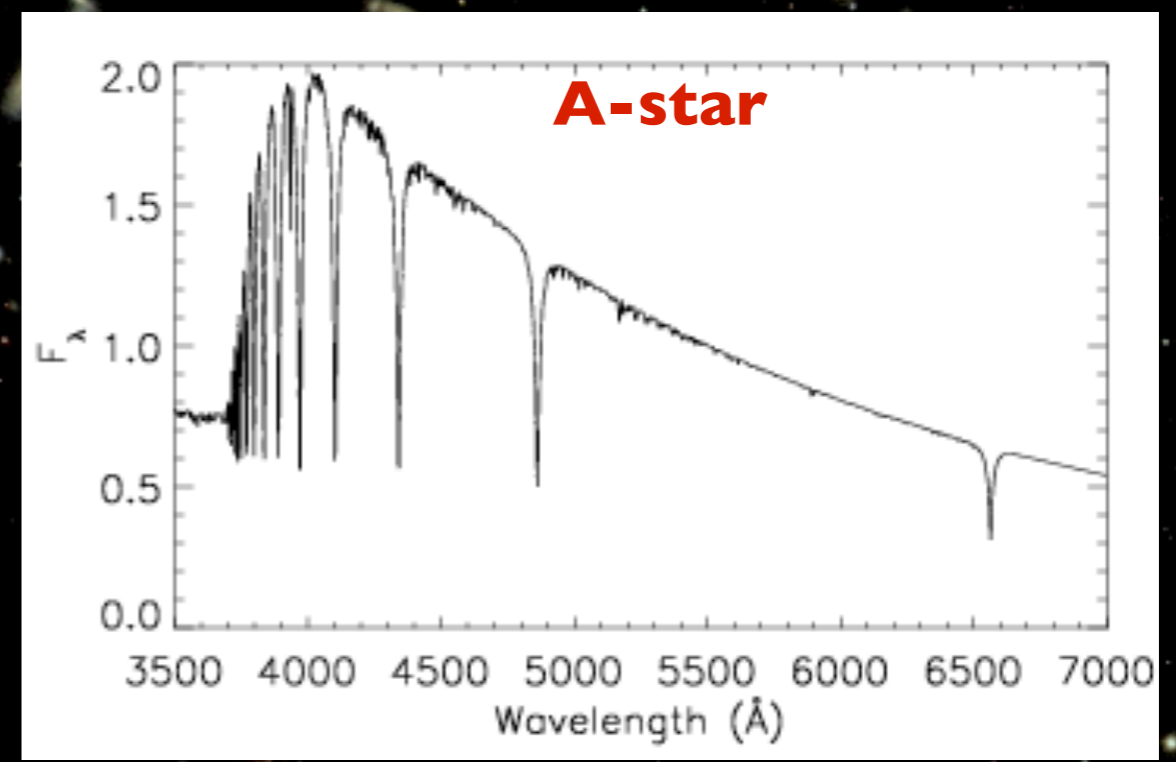
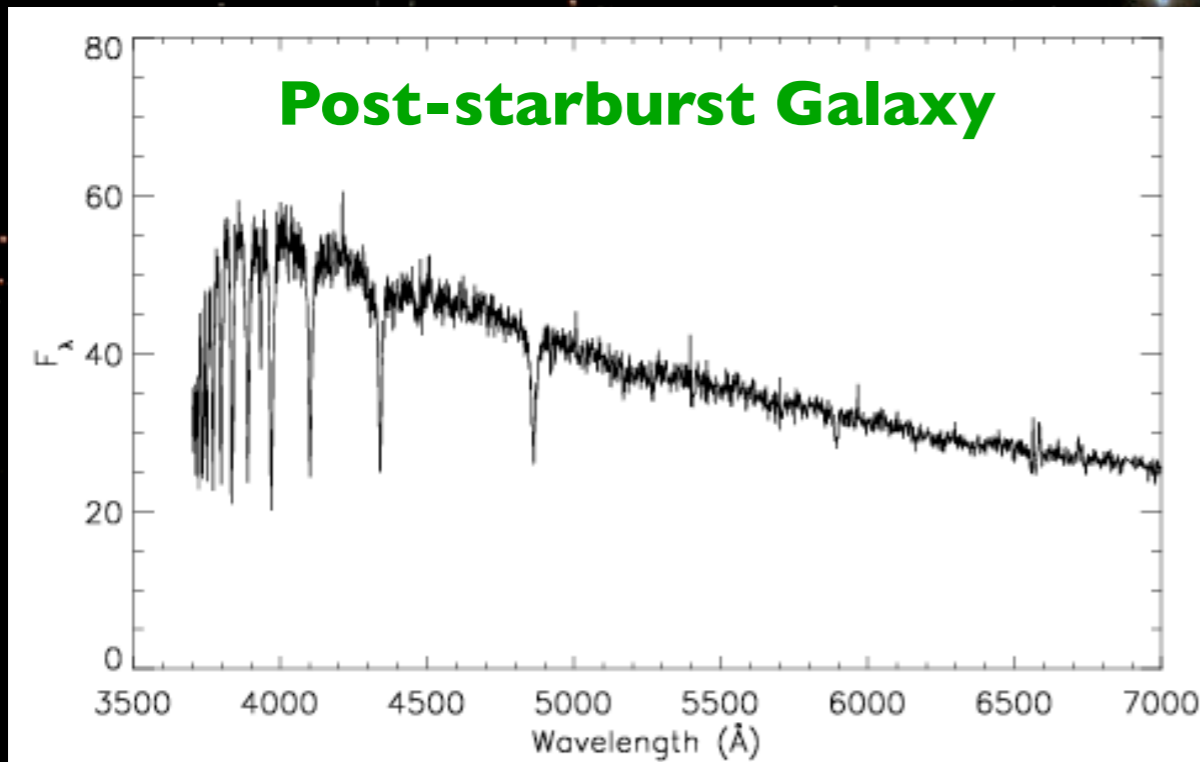
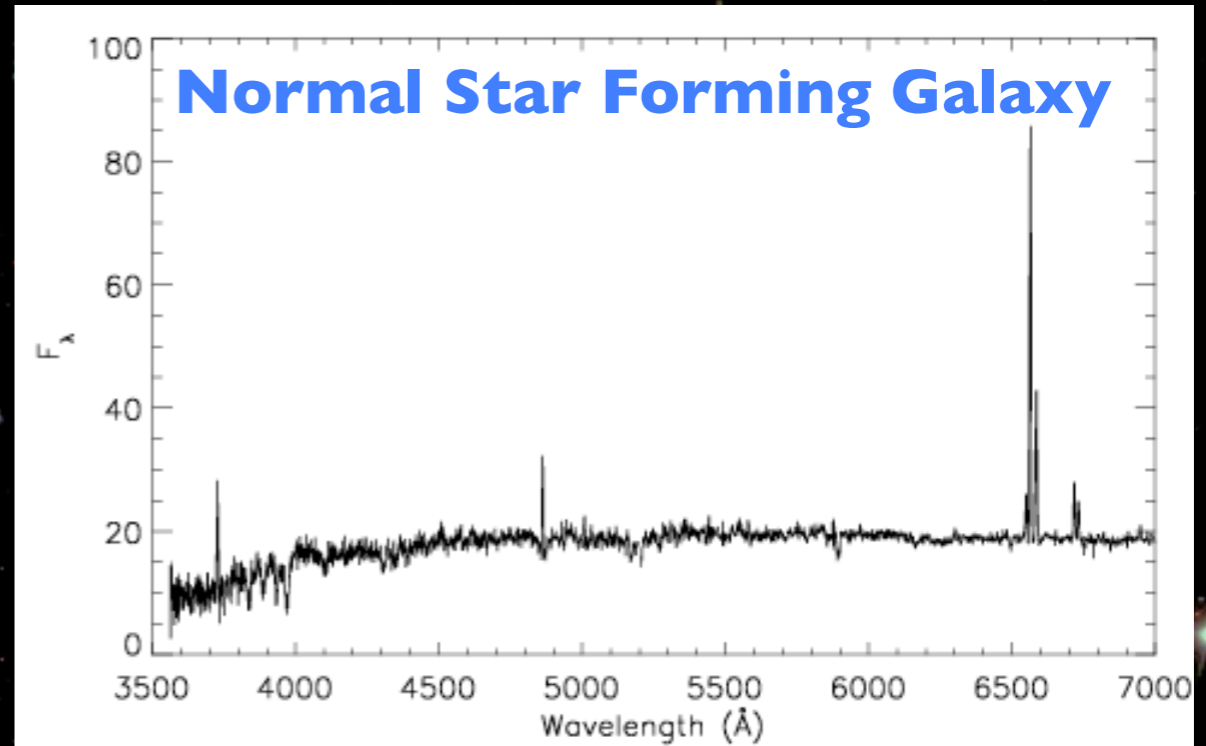
1) quasars ionize the cold gas used to trace outflows

2) absorption lines are harder to interpret: a small cloud near the quasar is indistinguishable from a galaxy-wide outflow



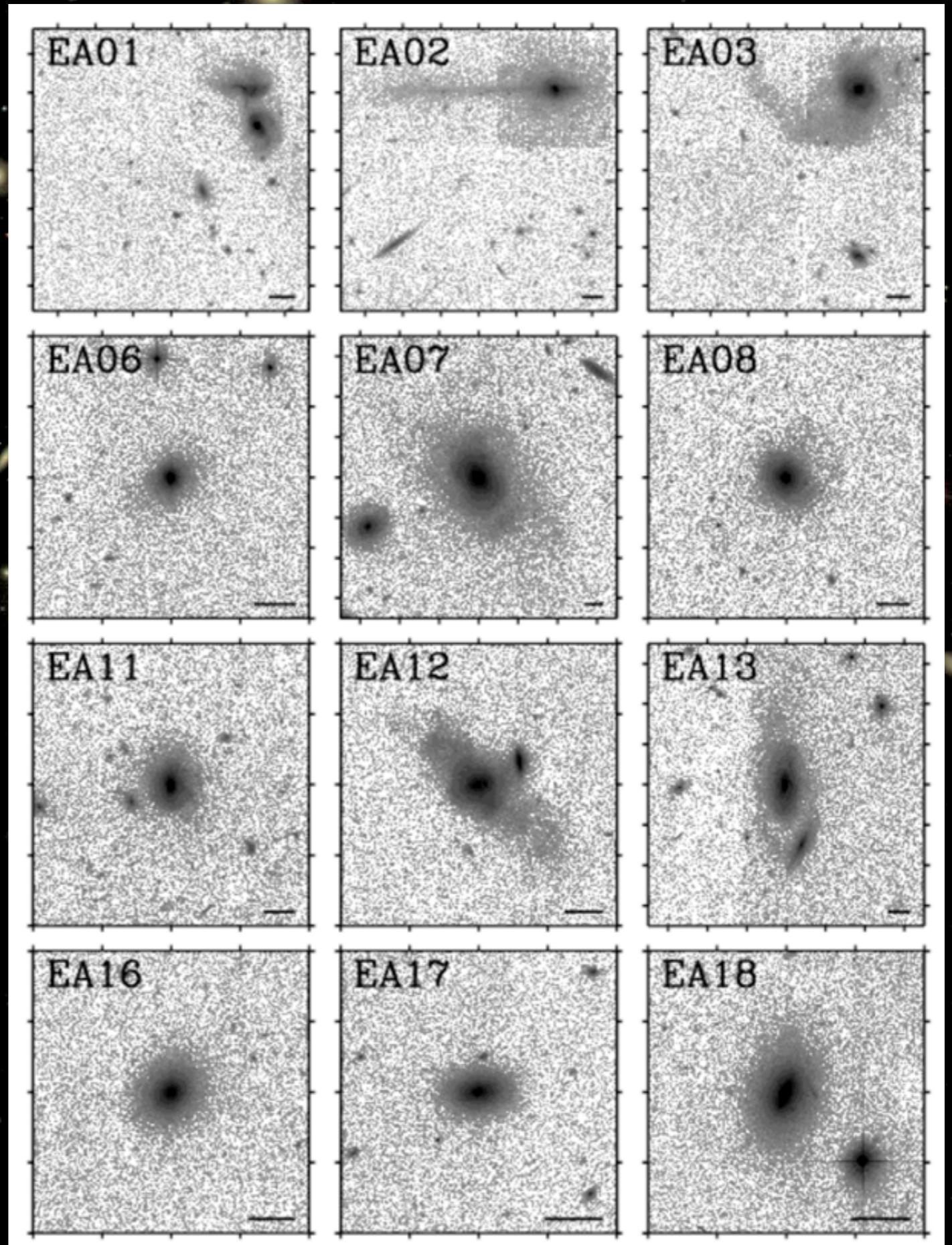
Post-starburst phase

Post-starburst galaxies can be identified by their distinctive spectra which lack strong nebular lines and resemble A-stars

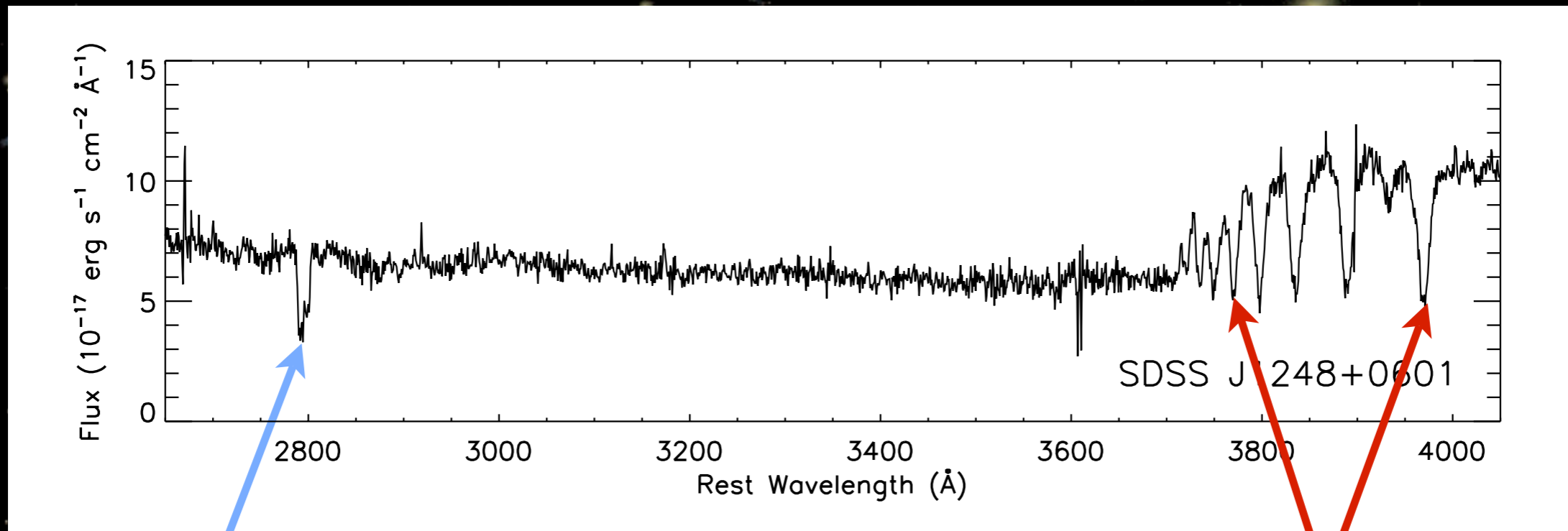


Local post-starbursts (E+A, K+A galaxies) are consistent with being late-stage mergers

- group environments
Zabludoff et al. 1996
- bulge-dominated morphologies
with faint tidal features
Yang 2004, 2007
- Low-luminosity AGN (LINERS)
Yan et al. 2005
- Rare locally ($\sim 1\%$ of galaxies)
but more common at high- z
Le Borgne et al. 2006



Our Goal: observe a sample of massive post-starbursts at $z=0.4 - 0.8$ to look for evidence of quasar-powered outflows



Interstellar Mg II

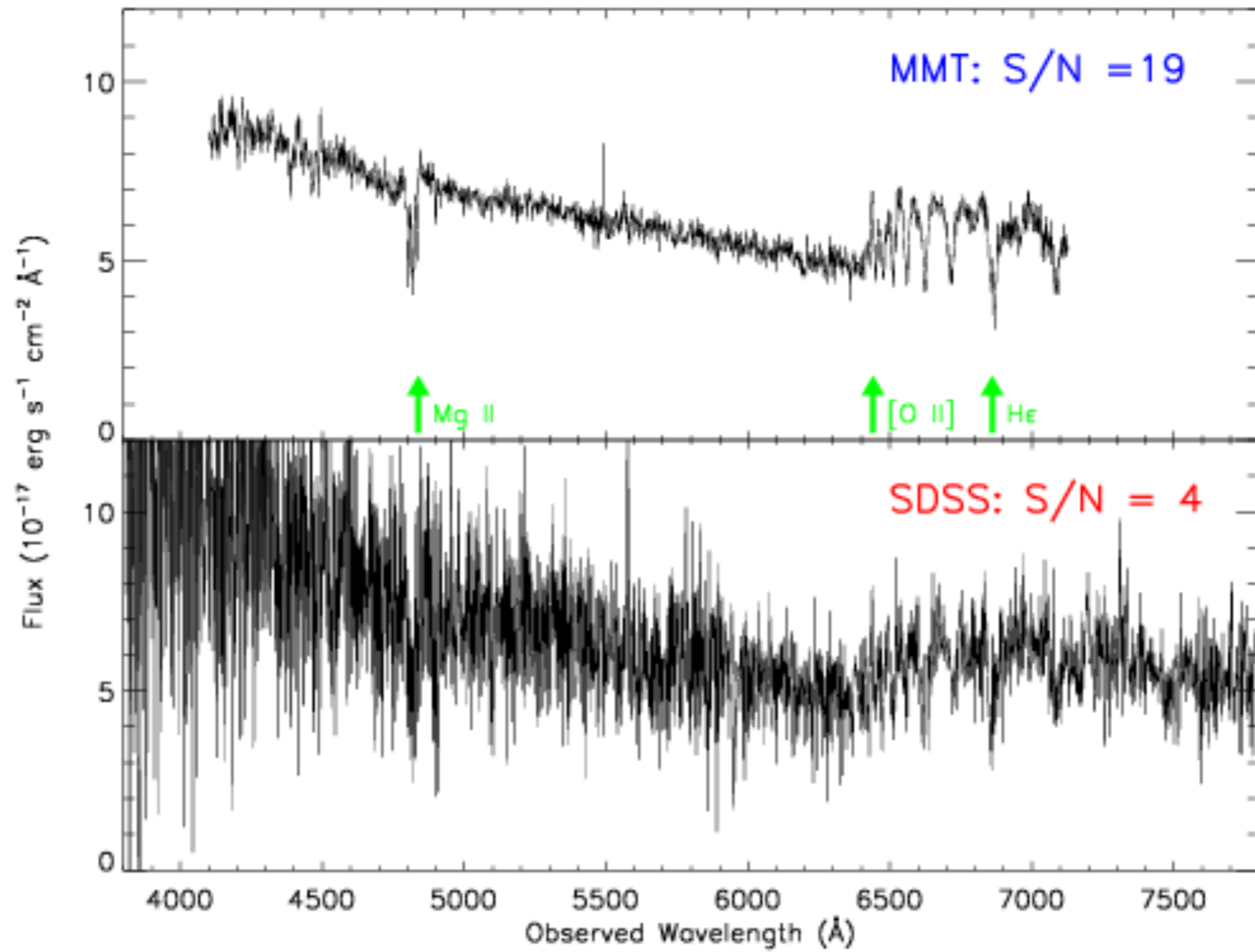
Stellar Balmer Lines



70 targets selected from SDSS

37 followed up with 6.5-m MMT

SDSS J2116-0634

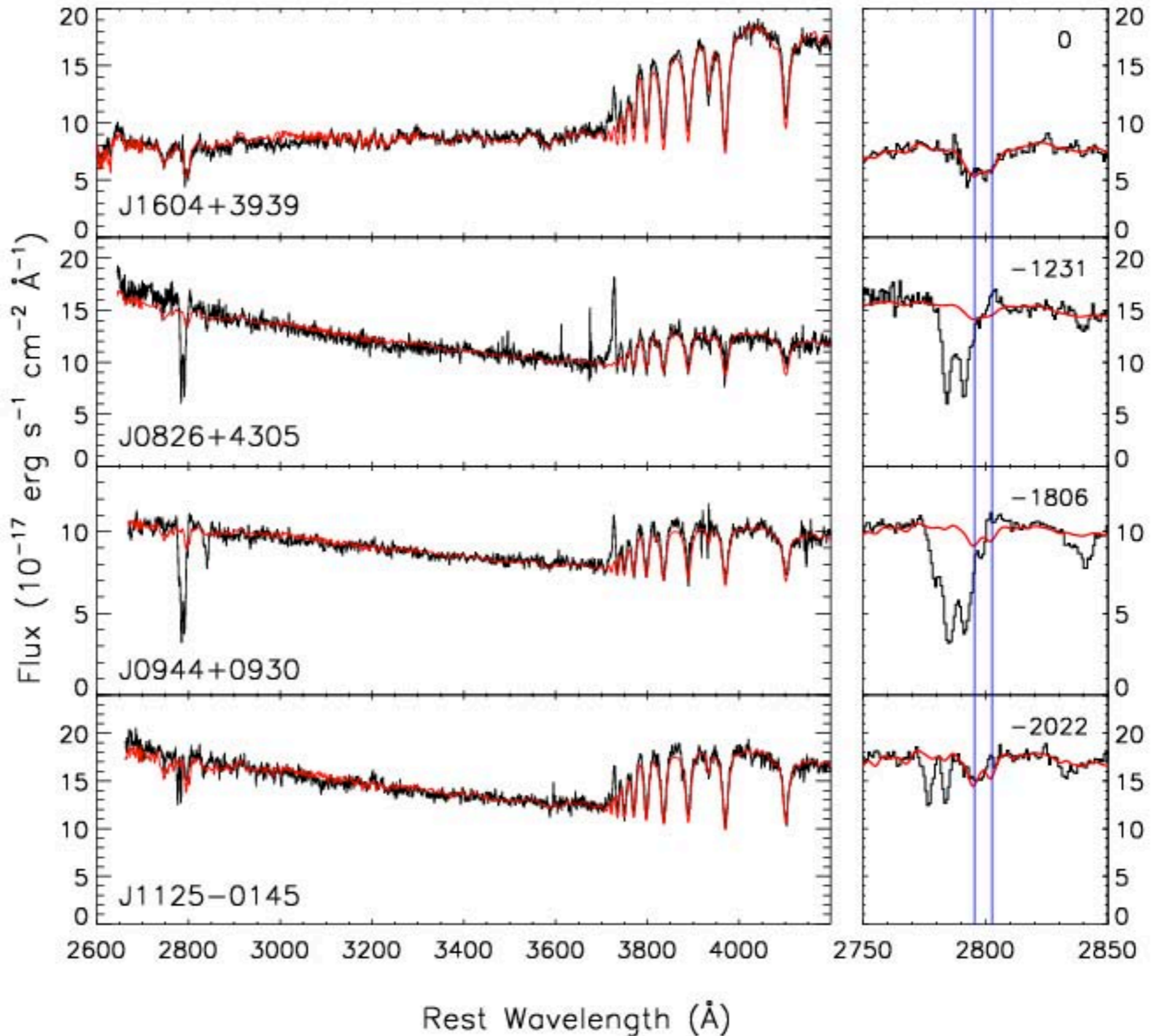


Mg II detected
in all 37
sources

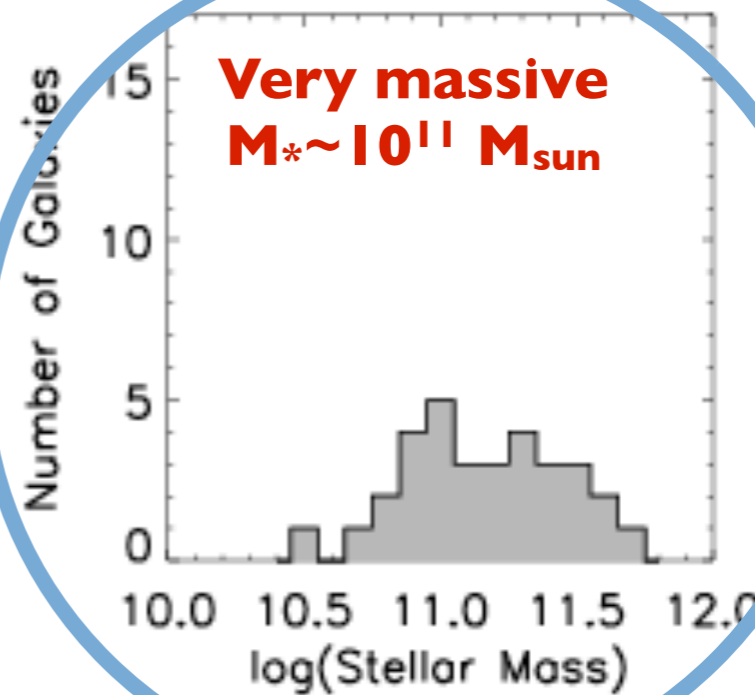
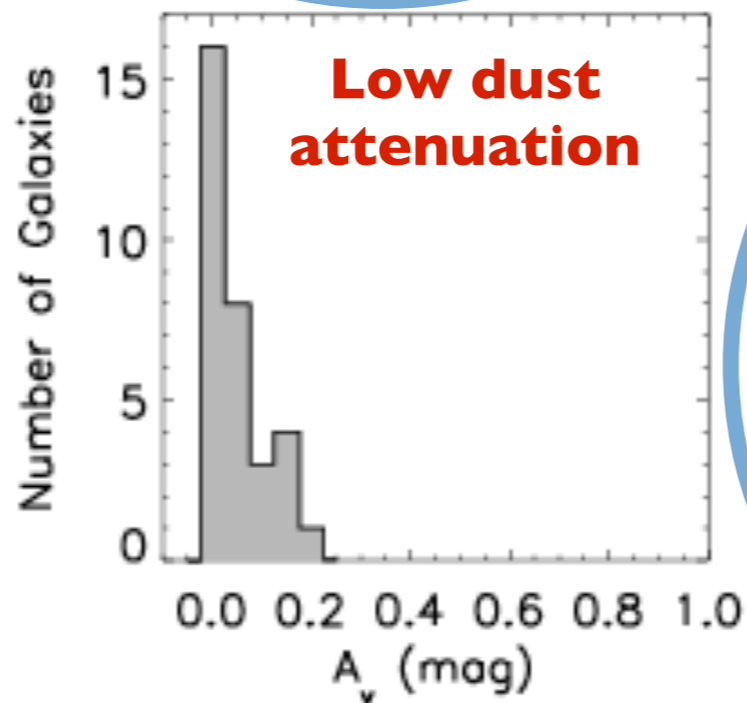
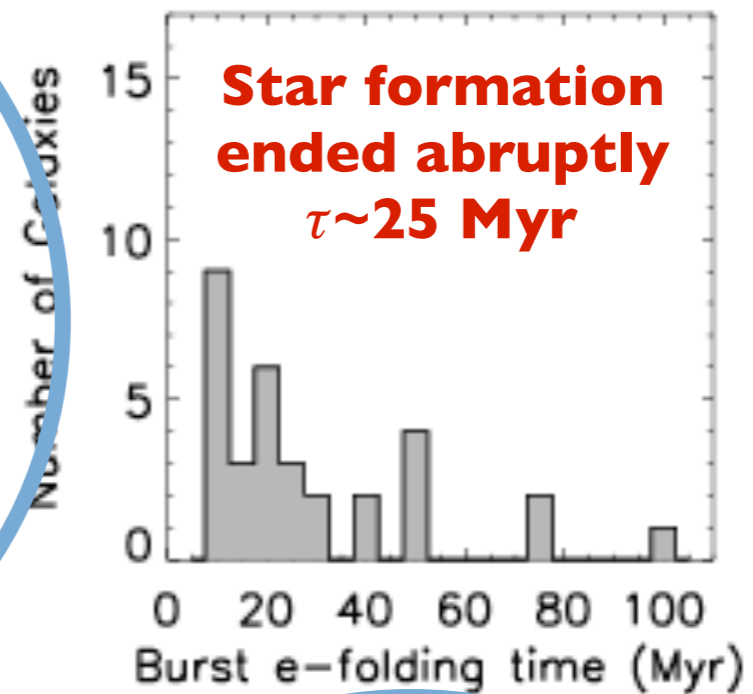
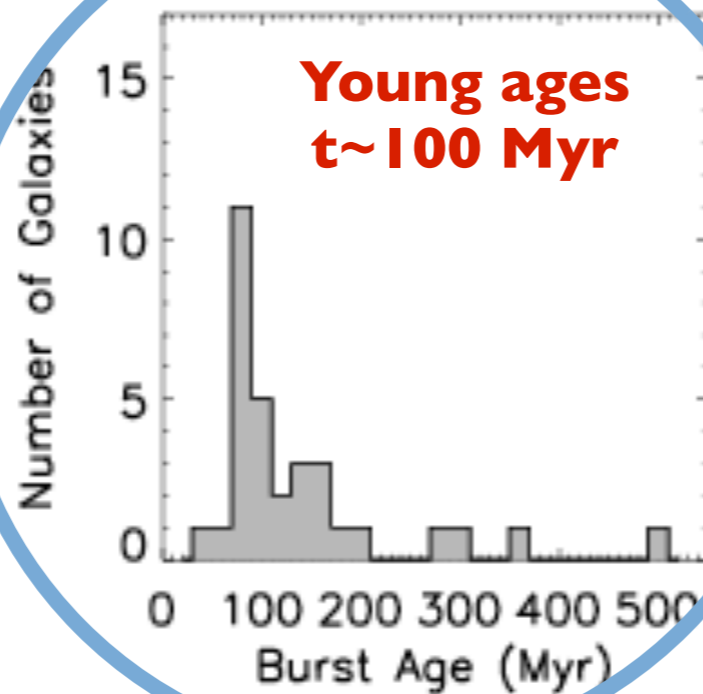
We fit **models**
to determine
the stellar
contribution

ISM absorption
in 2/3

Strongly
blueshifted!



We model the
galaxies stellar
populations



Fossil Galactic Winds

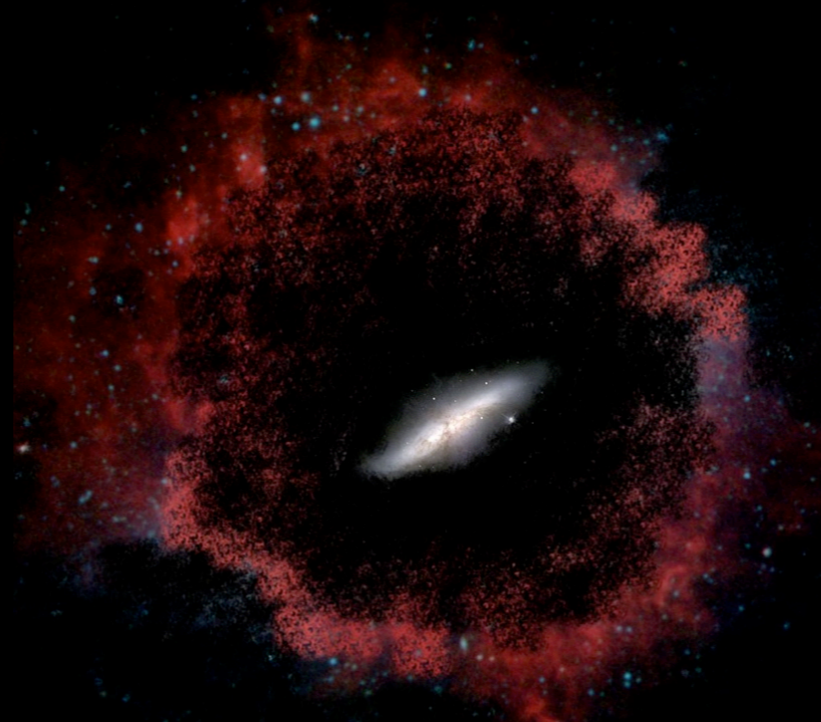
Interstellar Mg II in 24/37 galaxies

In all cases it is blueshifted from the stellar lines by 400 - 2000 km/s

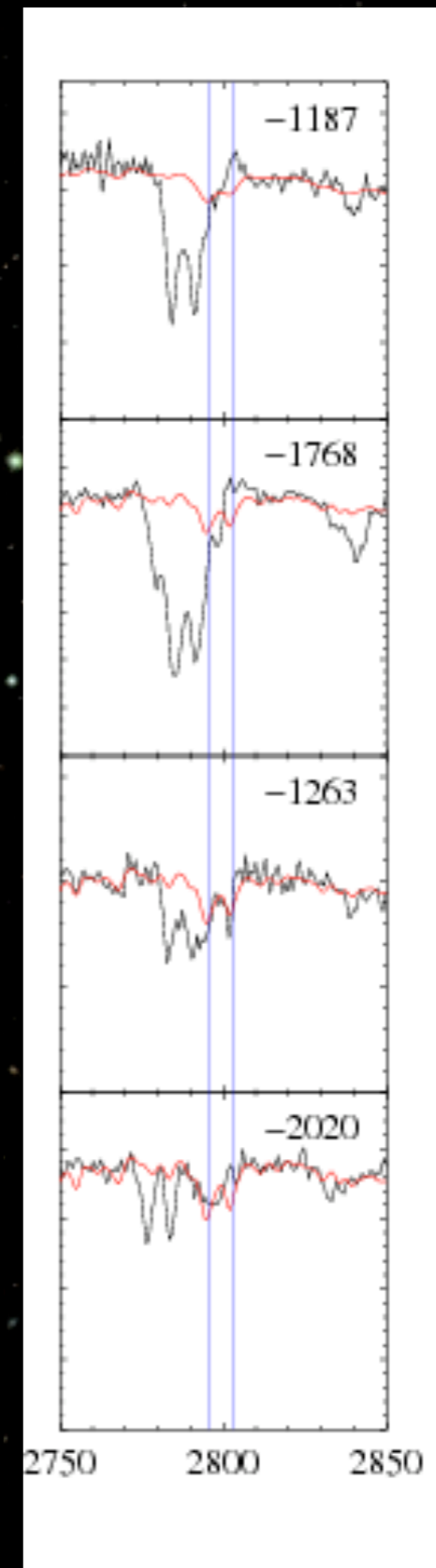
We assume that the outflows were launched at the peak of the galaxies' activity ~ 100 Myr ago



starburst, $t=0$ Myr



post-starburst, $t=100$ Myr



How far does the wind get?

$$d = v t$$

$$v = 1000 \text{ km/s}$$

$$t = 100 \text{ Myr}$$

$$d = 100 \text{ kpc}$$

winds will escape

How much mass is in the outflow?

$$N(\text{H}) \sim 10^{20} \text{ cm}^{-2}$$

$$M_{\text{wind}} = 10^{9.3} - 10^{11} M_{\text{sun}}$$

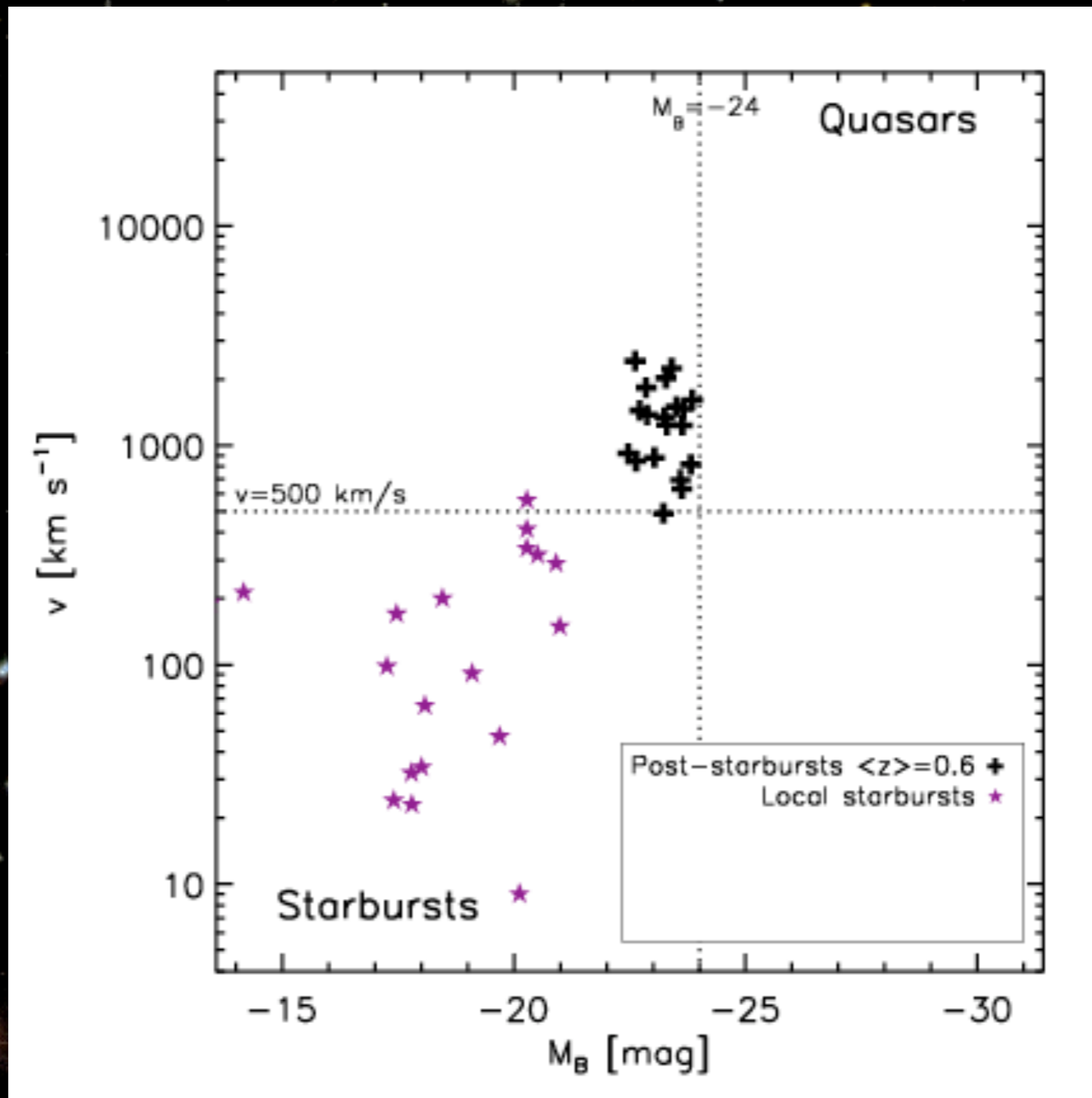
depends strongly on assumed geometry!

5 - 50% of the galaxy's baryons



Are these outflows really powered by quasars?

Outflow velocity

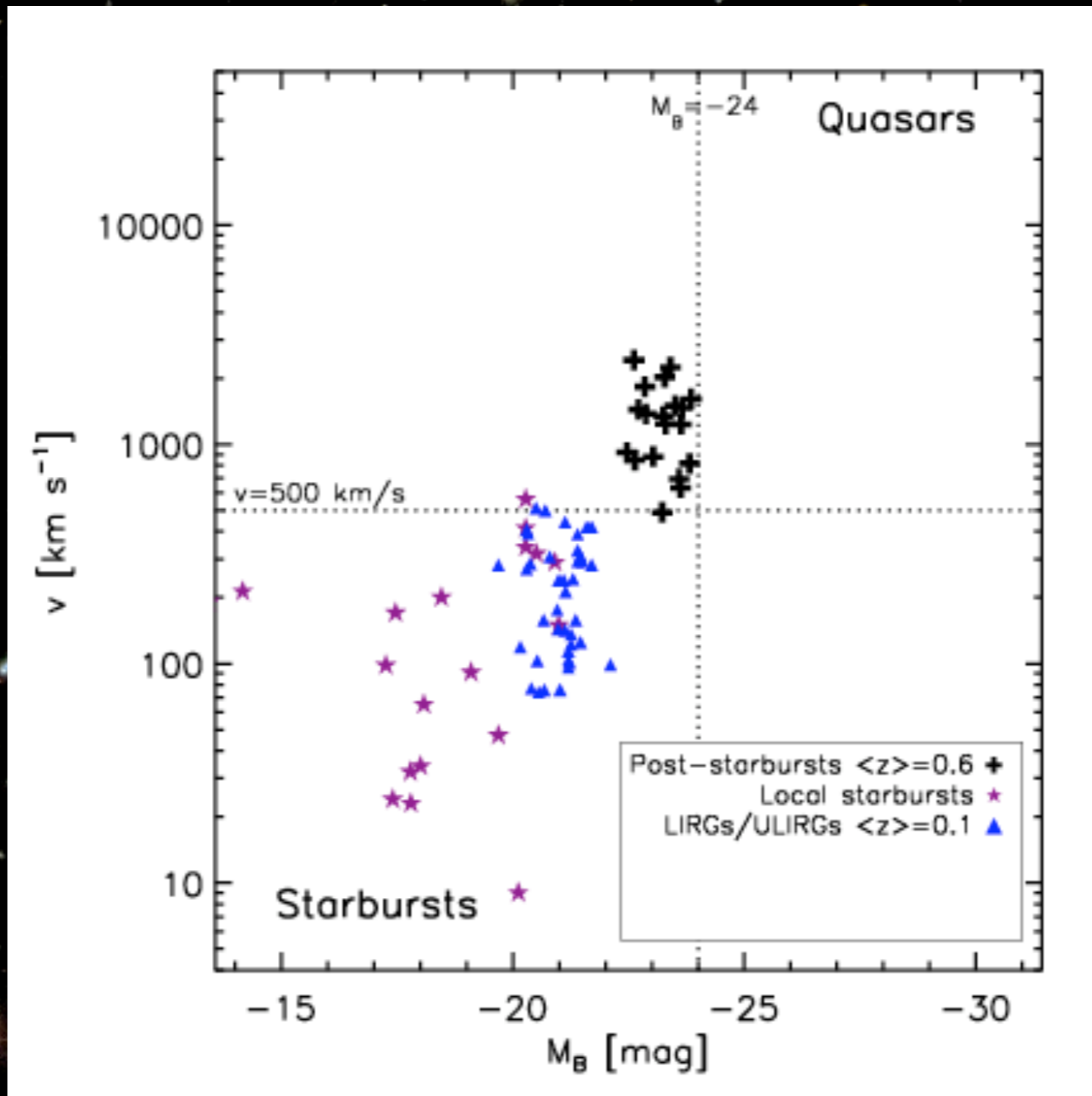


Local Starbursts
Schwartz &
Martin 2004, 2006

B-band Magnitude

Are these outflows really powered by quasars?

Outflow velocity

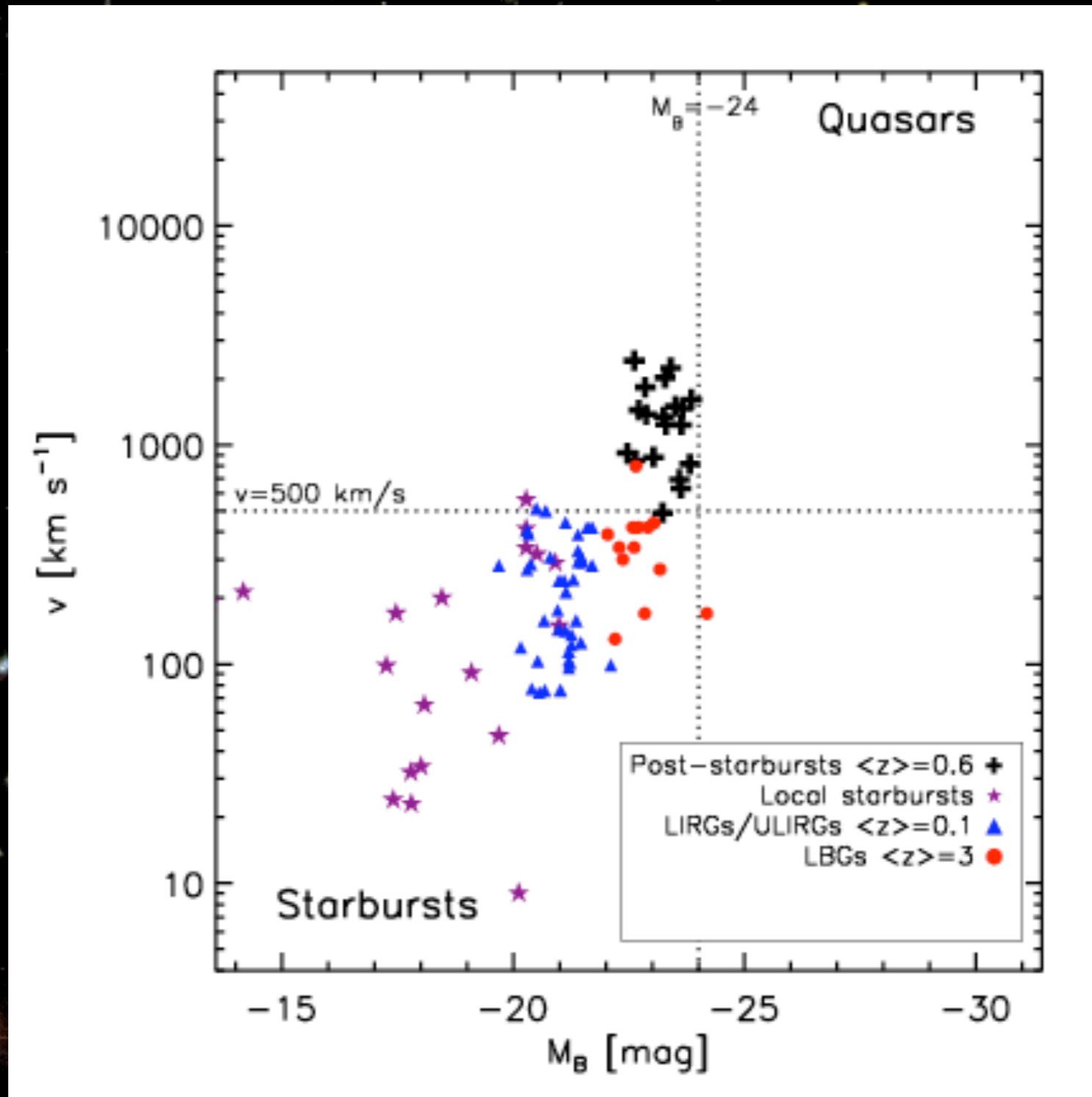


LIRGs/ULIRGs
Rupke et al. 2005
Martin 2005

B-band Magnitude

Are these outflows really powered by quasars?

Outflow velocity

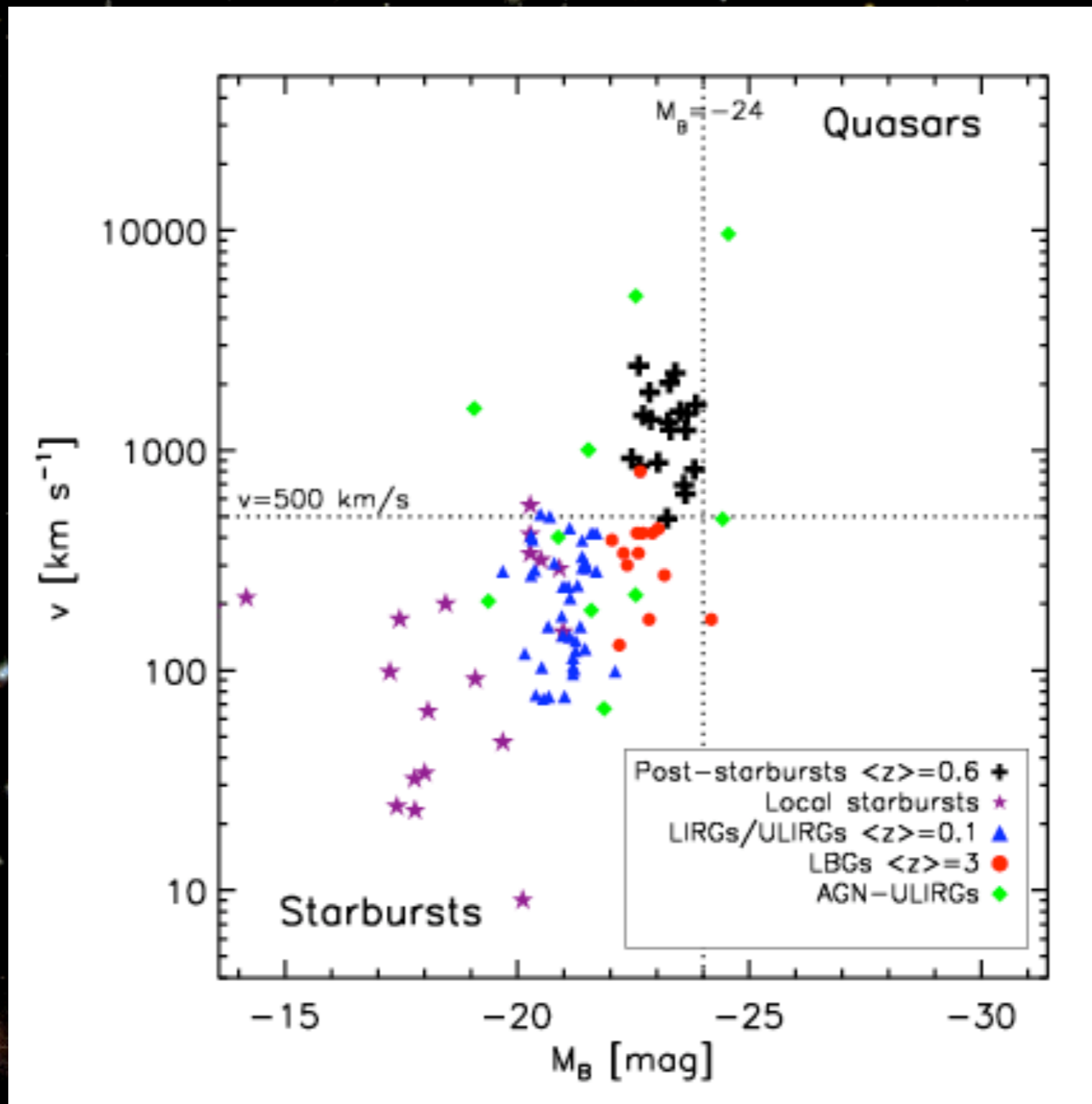


Lyman Break
Galaxies ($z=3$)
Pettini et al. 2001

B-band Magnitude

Are these outflows really powered by quasars?

Outflow velocity

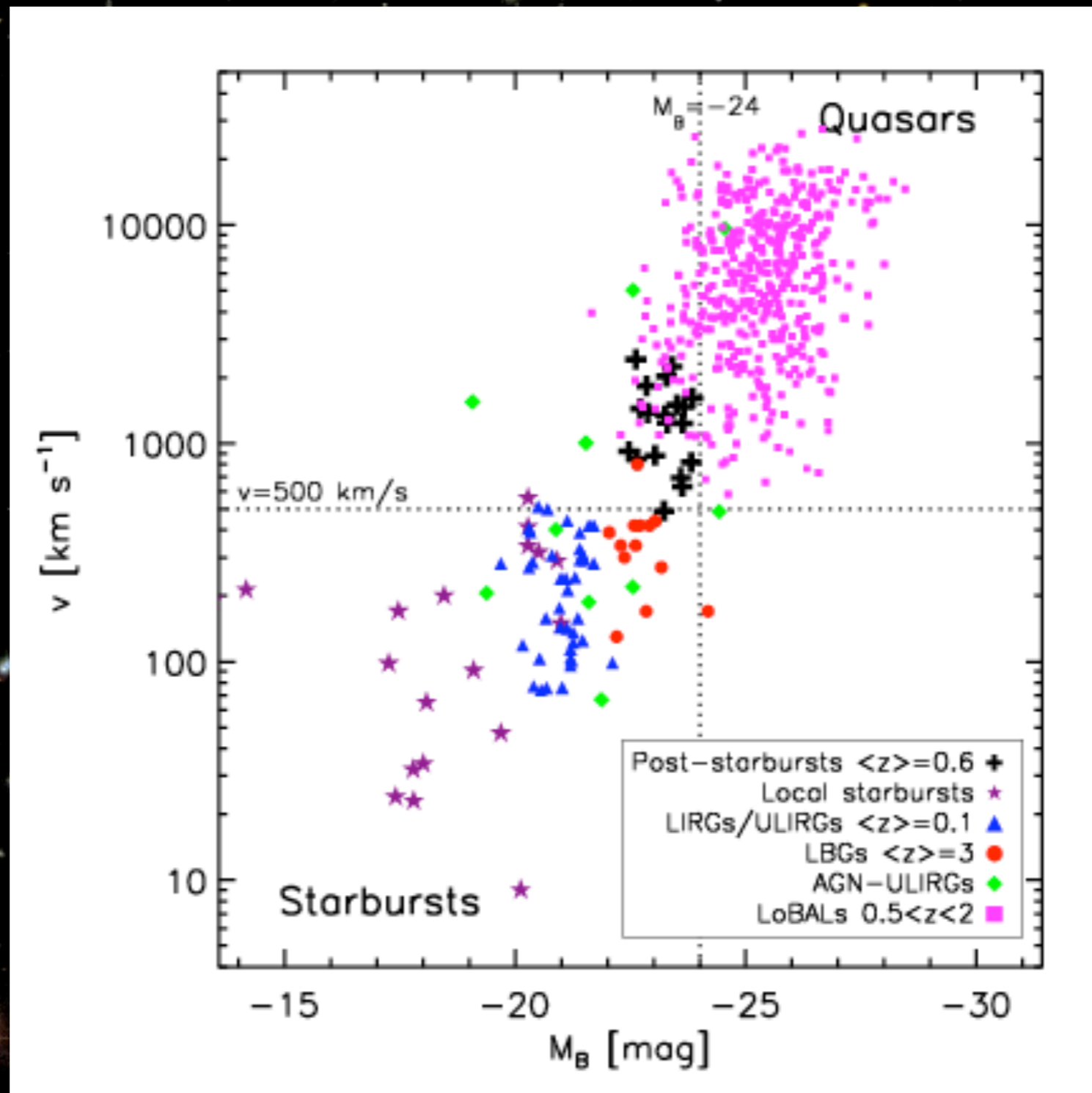


AGN-ULIRGs
Rupke et al. 2005

B-band Magnitude

Are these outflows powered by quasars? Probably

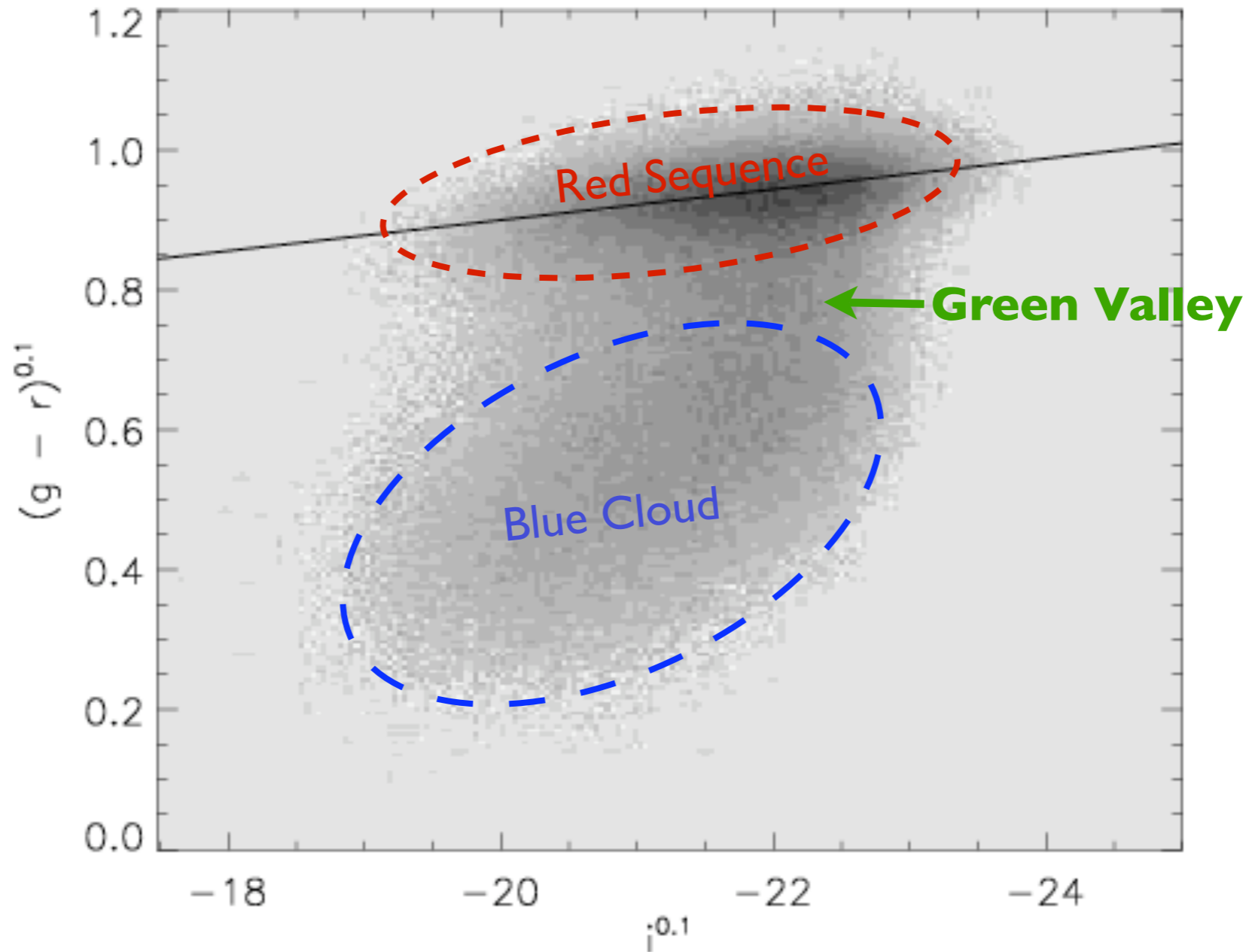
Outflow velocity



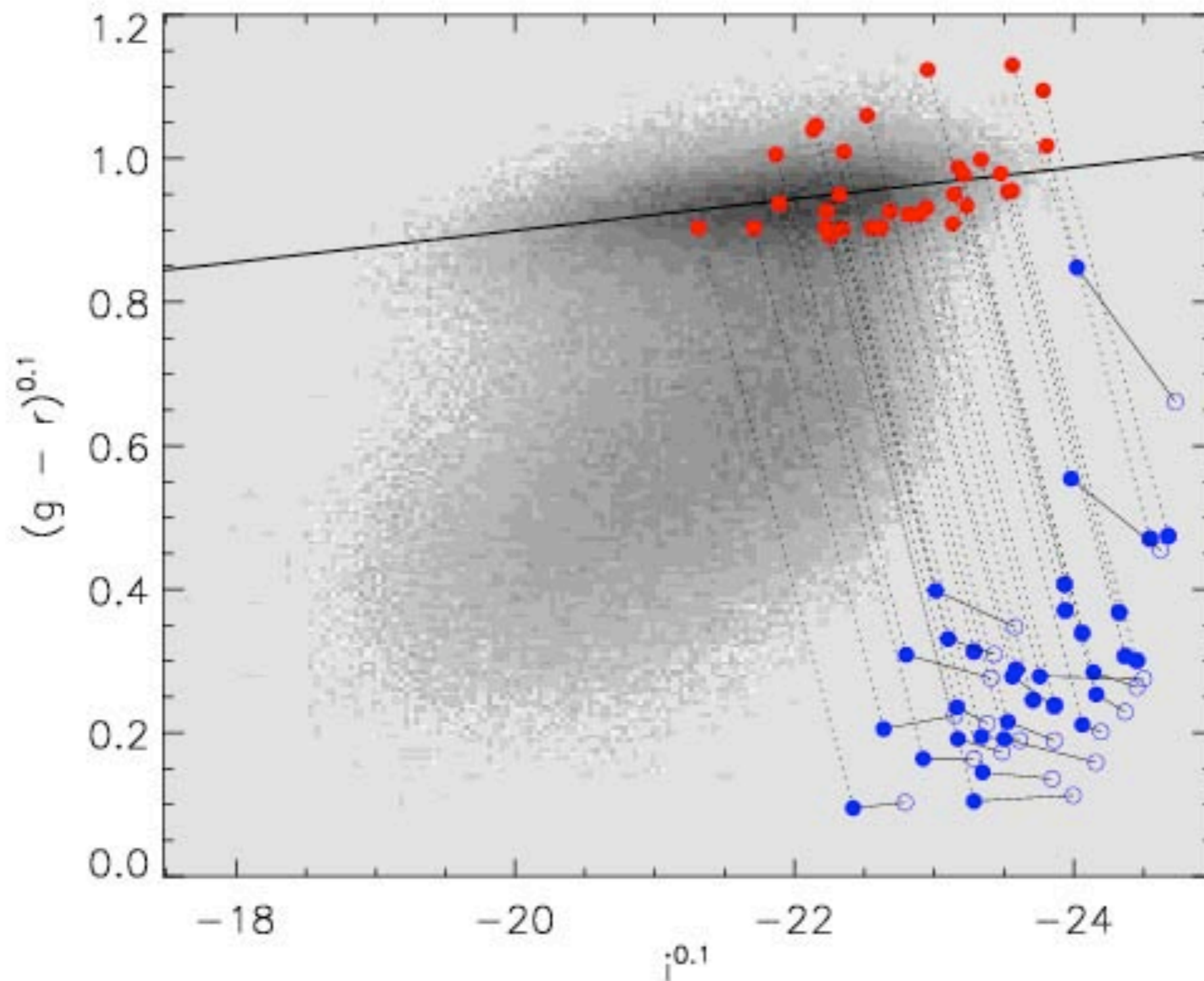
LOBAL QSOs
Trump et al. 2006

B-band Magnitude

What will these galaxies become?



Our post-starbursts will evolve passively into intermediate mass to massive early-types



**1 Gyr
later**

**Now
($z \sim 0.6$)**

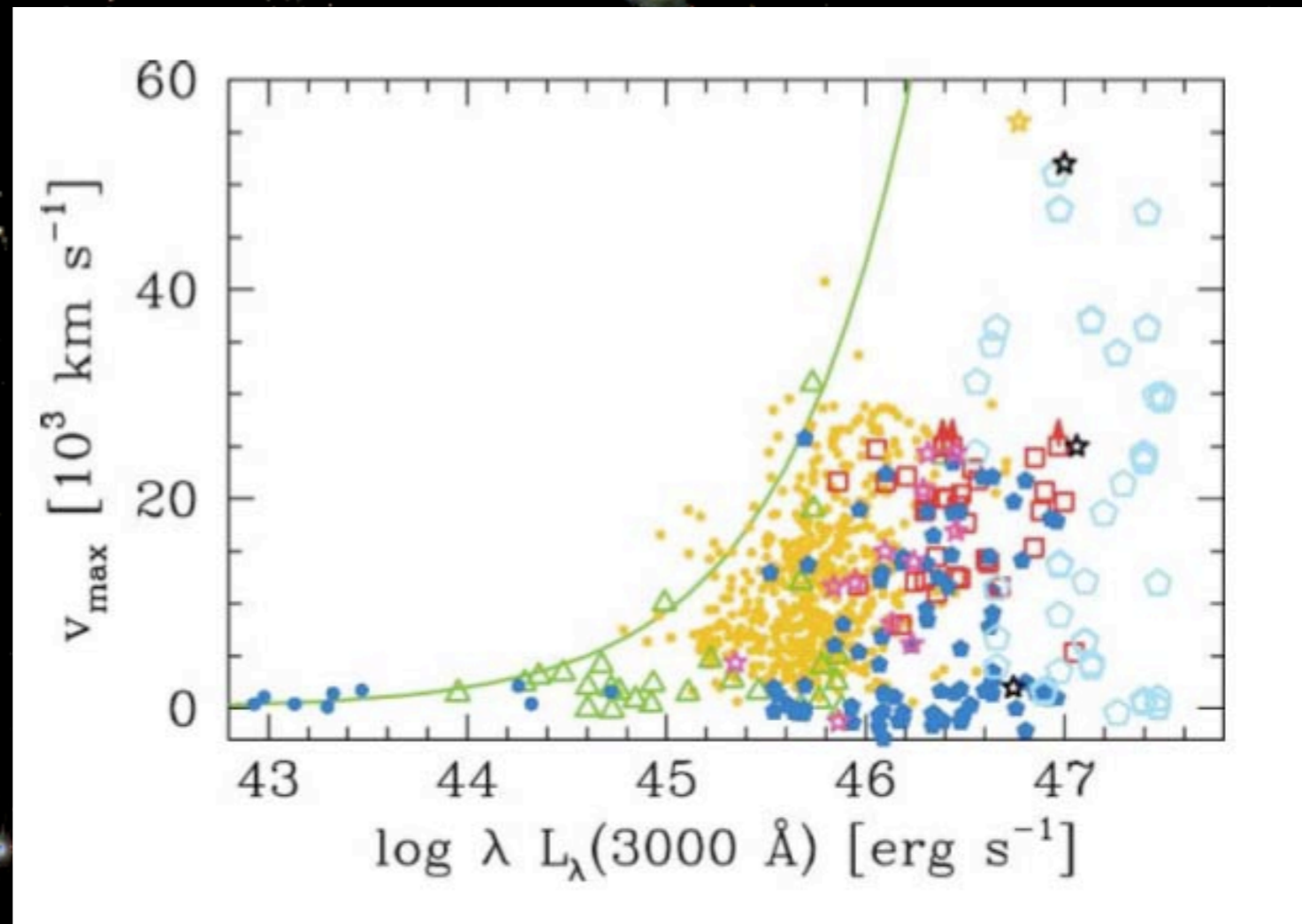
How does quasar feedback work?

I) Momentum-driven winds

Radiation pressure on dust grains?

Line-driving?

Some evidence from BAL Quasars



Ganguly & Brotherton 2008

2) Mechanical Energy Injection by Radio Jets

Emission line nebula in
 $z \sim 2$ radio galaxies

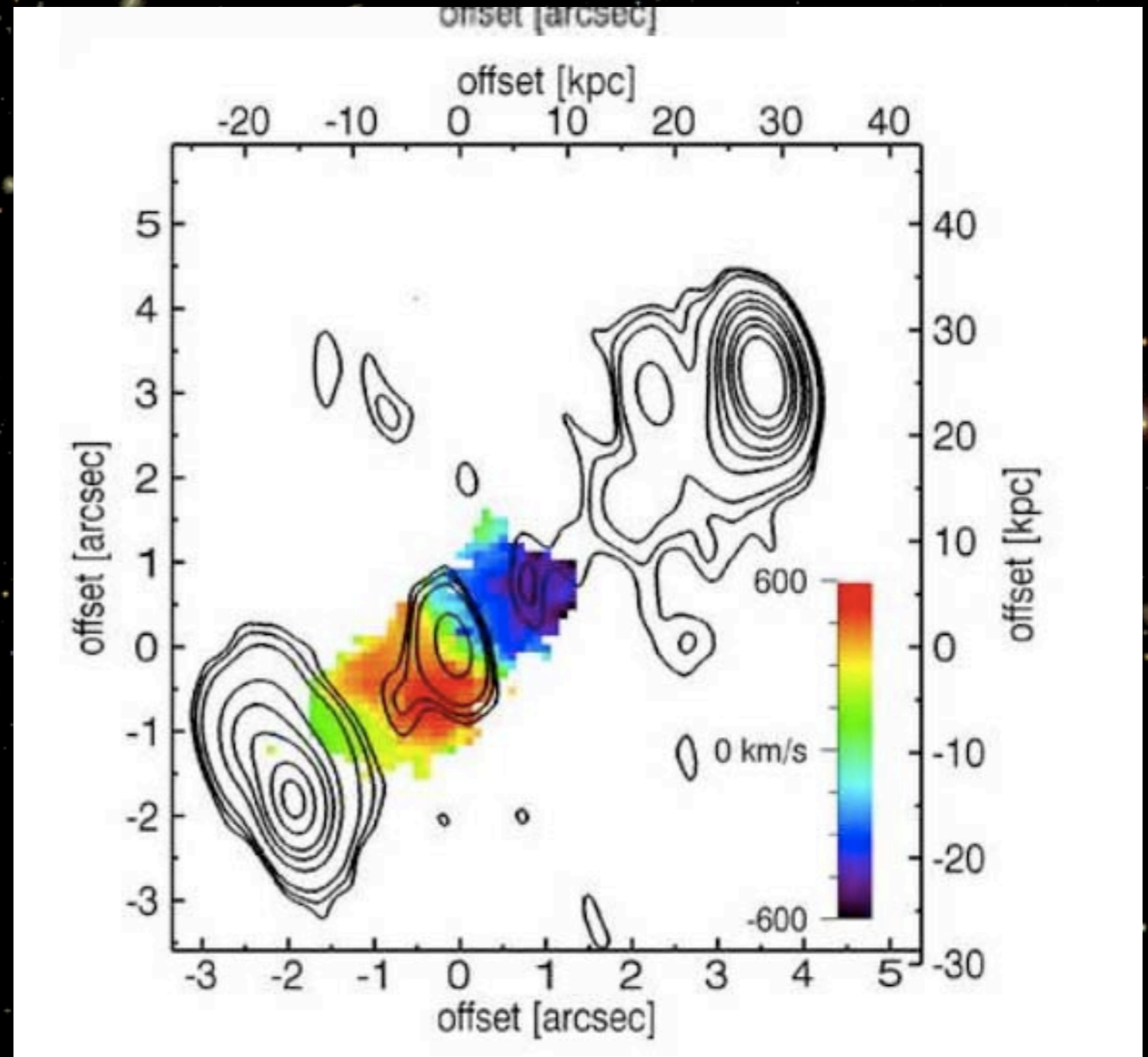
SINFONI maps of [OIII]

sizes 10 - 20 kpc

~ 500 km/s outflows

~ 1000 km/s line widths -
powerful shocks

$10^{10} M_{\text{sun}}$ of gas

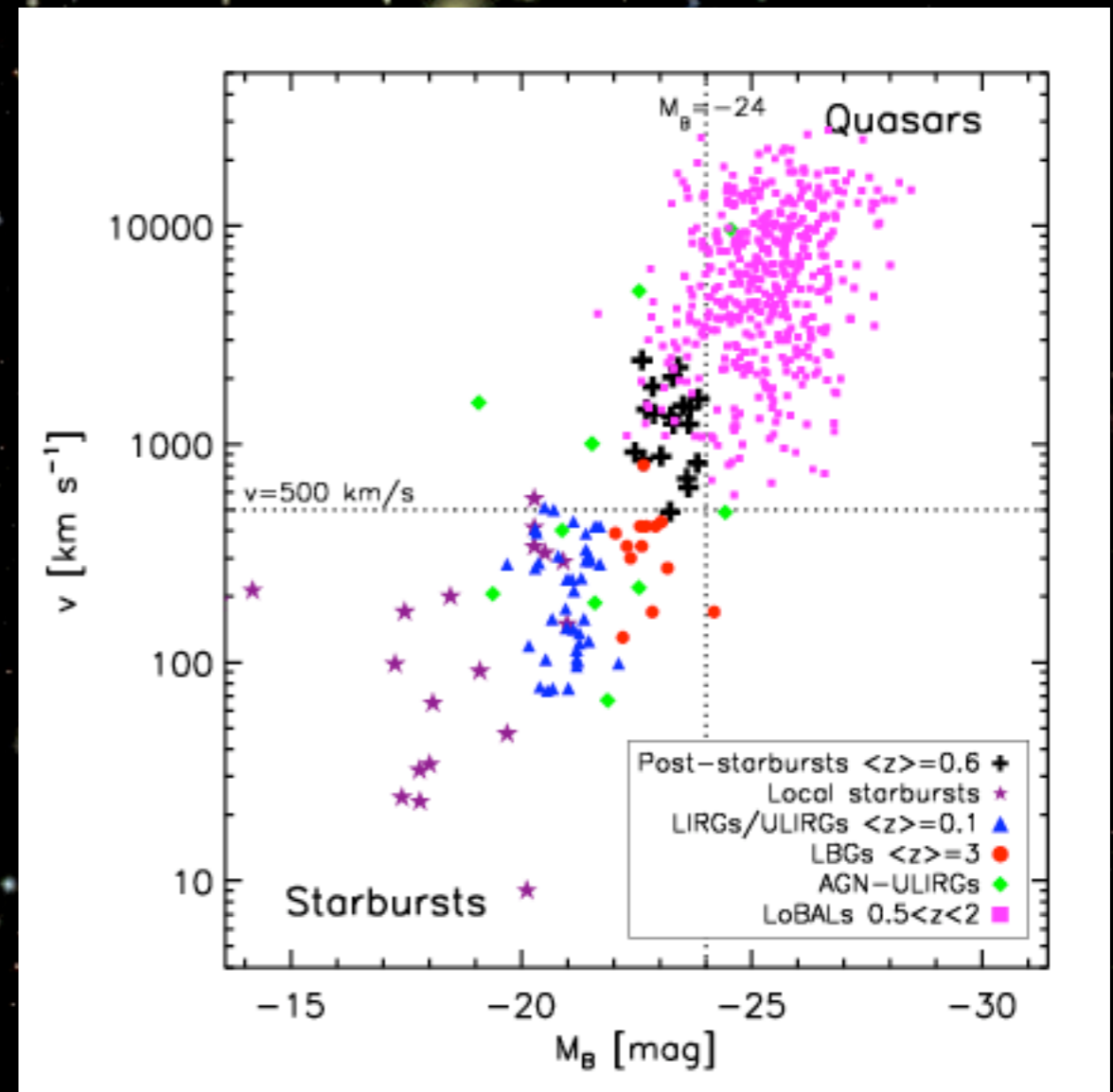


Nesvadba et al. arXiv:0809.5171v2.

Summary

We looked for signs of fossil galactic winds in 37 post-starburst (post-quasar) galaxies

- We detected blueshifted Mg II in 2/3 of the sample
- Outflow velocities are $v=400 - 2000$ km/s
- The velocities exceed those of starbursts and show some overlap with BAL quasars
- The outflows entrain 5 - 50% of the galaxies' baryons



Quasars are likely to have played a role in expelling the cool gas and quenching star formation