

The Hobby-Eberly Telescope Dark Energy Experiment (HETDEX)

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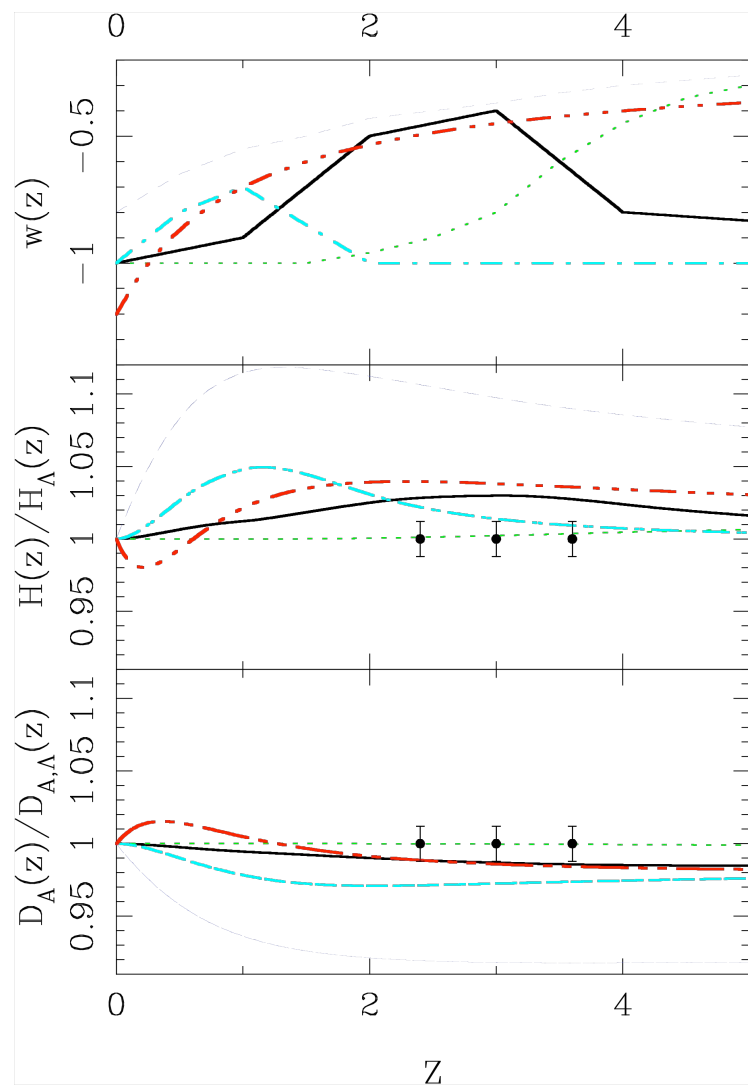
Goals for HETDEX

- HETDEX measures redshifts for about 1 million objects from $2 < z < 4$
- Baryonic oscillations determine H and D_A to 1% and 1.4% in 3 bins
- Constraints on constant w to about one percent
- Tightest constraints on evolving w at $z=0.4$ (to a few percent)

The issues are:

- measurements of H , D_A , and redshift range
- and then going from H , D_A to w

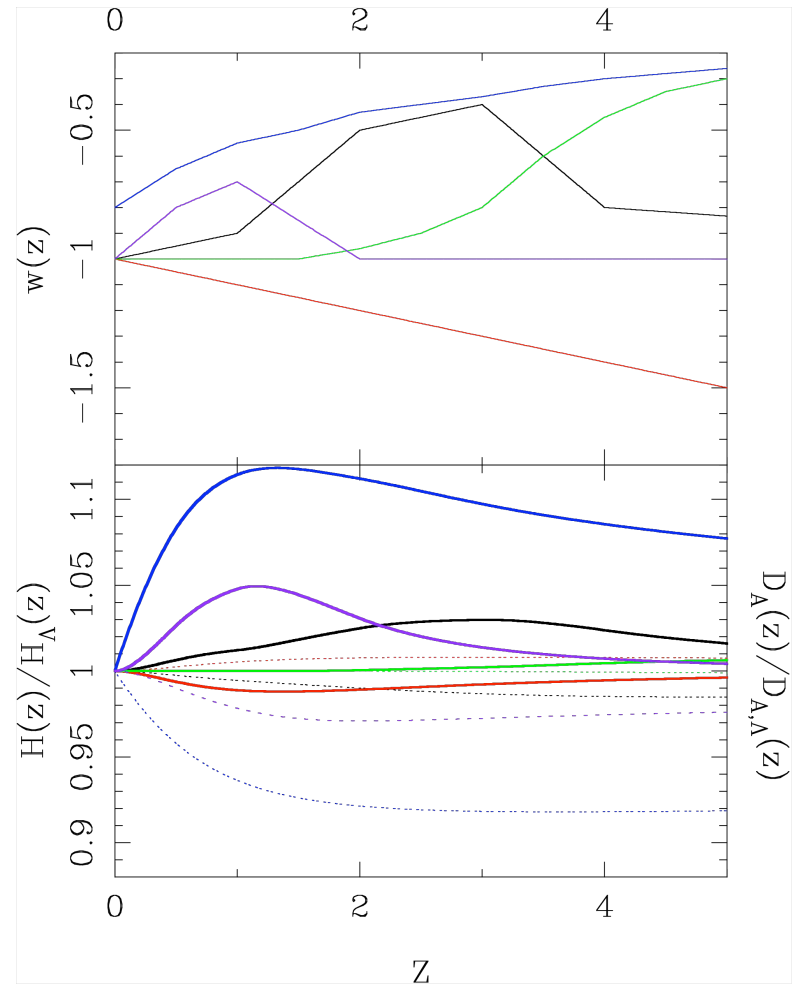
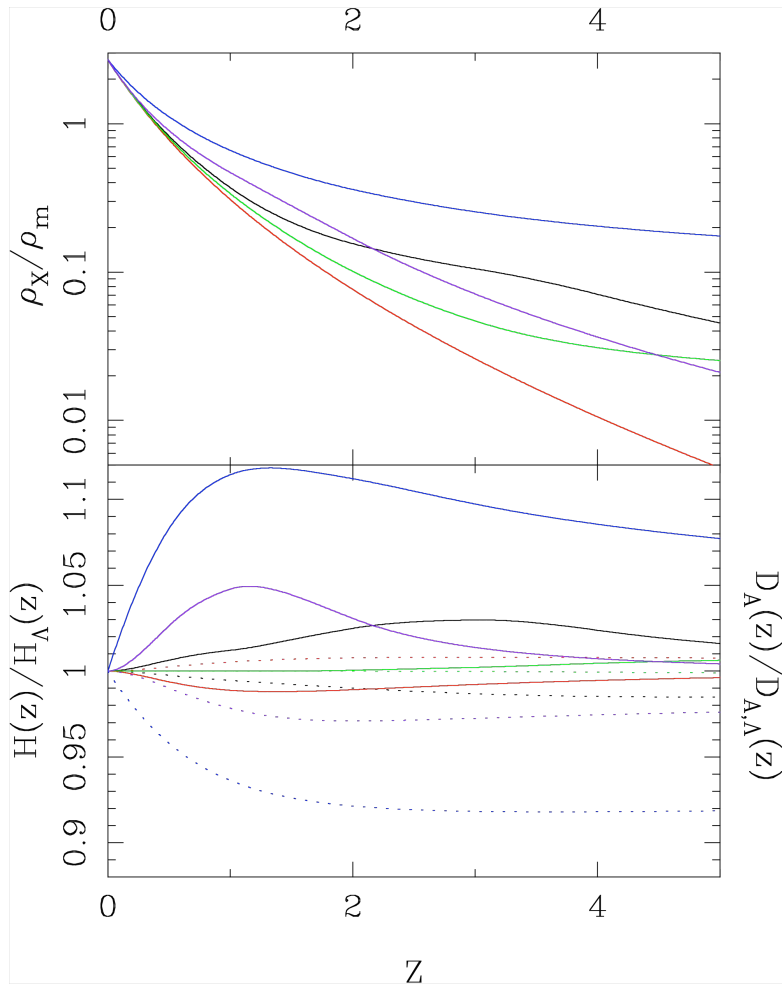
H(z), Da(z), and w(z)

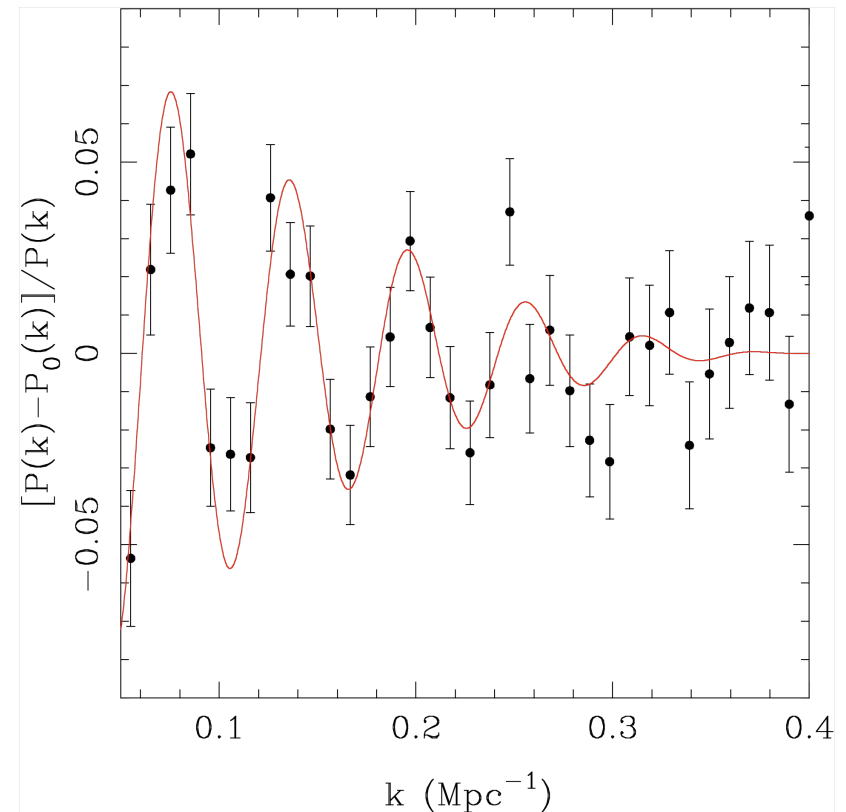
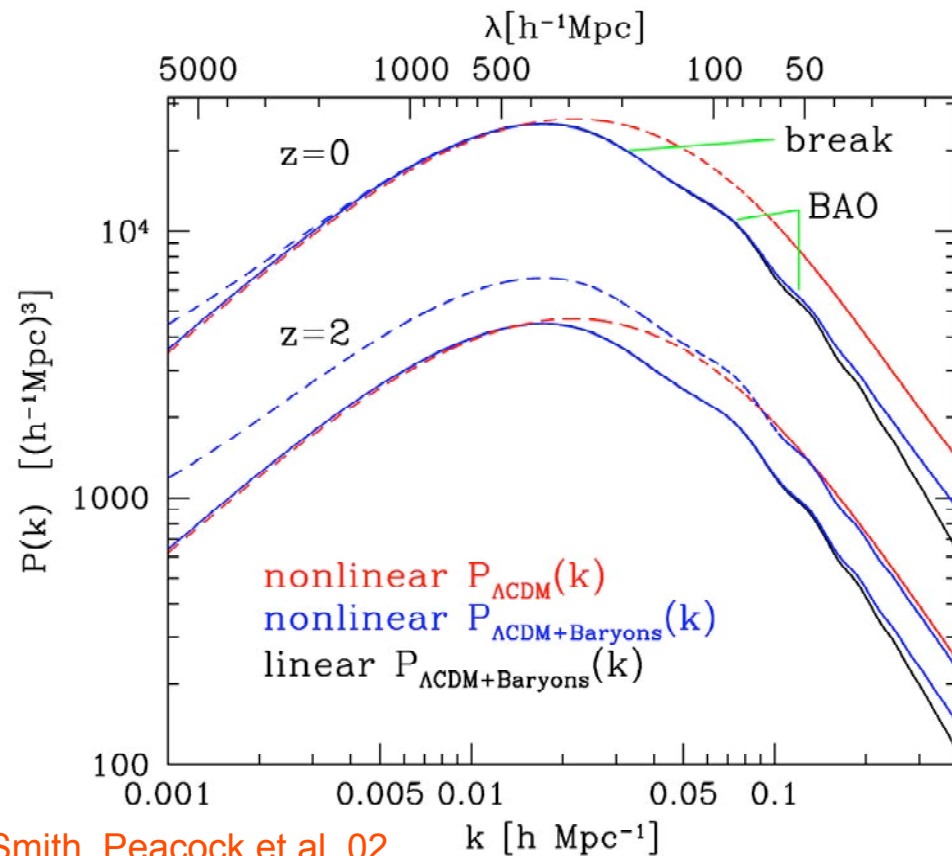


$$H(z) = h \sqrt{\Omega_m (1+z)^3 + \Omega_x \exp\left[3 \int_0^z \frac{1+w(z)}{1+z} dz\right]}$$

The integral dependence of H on w allows low- z constraints from high- z observations

The need for accuracy vs. z

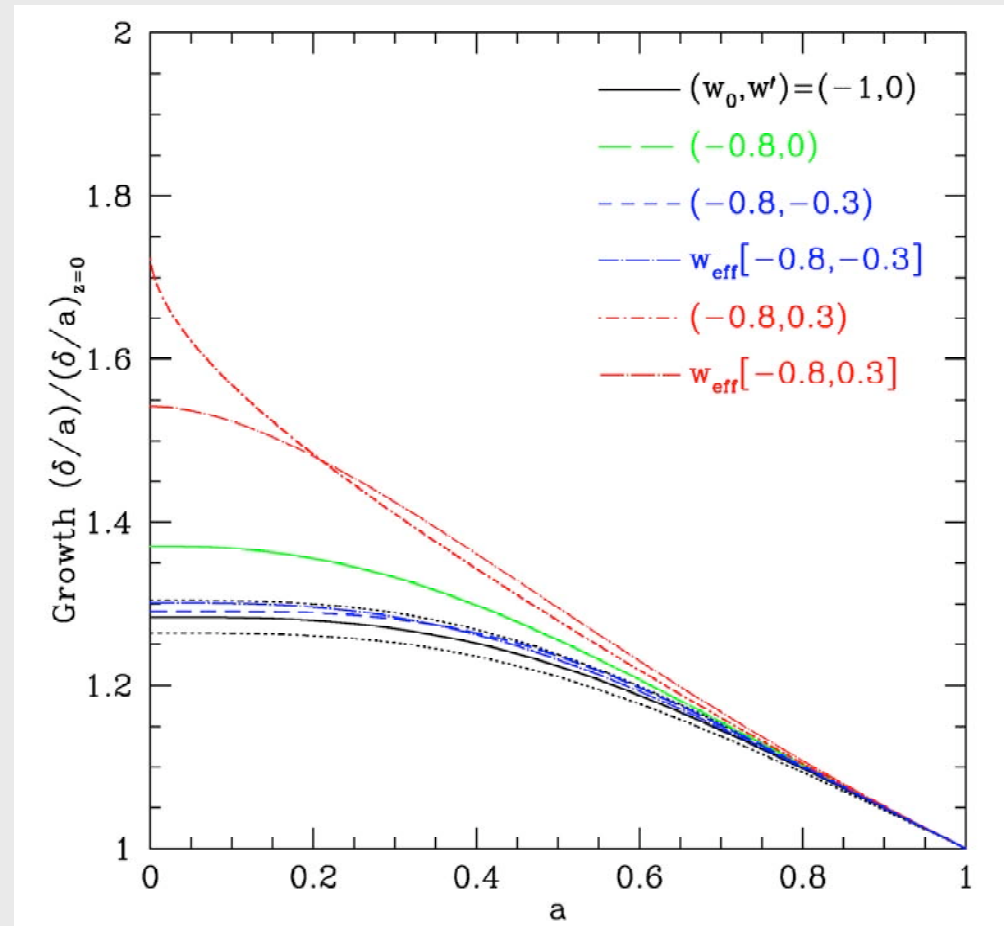




Smith, Peacock et al. 02

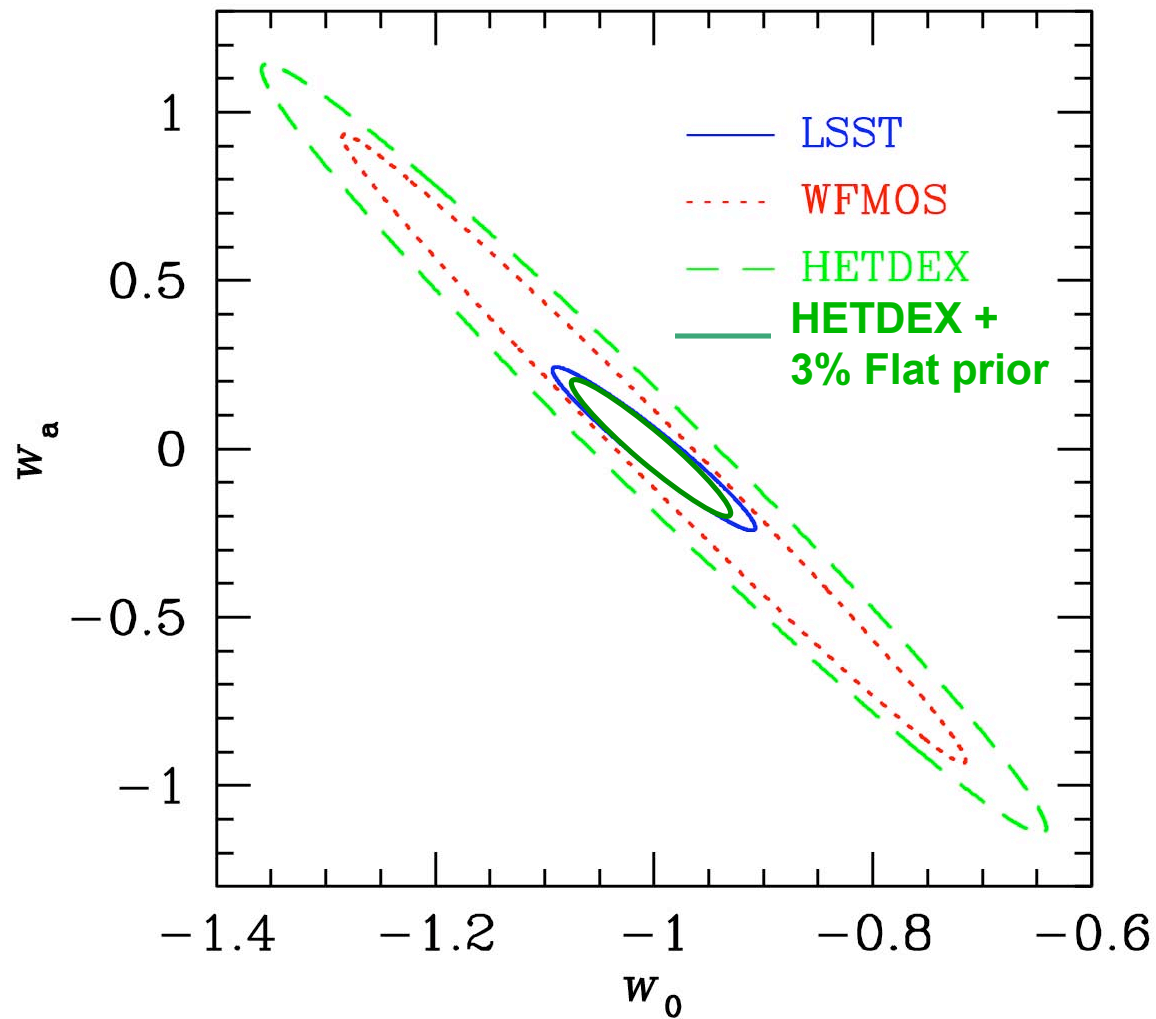
Power spectra analysis has 5 measures to exploit:

1. Phases of the oscillations (only one used now): geometric
2. Amplitude of oscillations: structure growth
3. Amplitude of $P(k)$: structure growth
4. Linear/non-linear transition: geometric
5. Turn-over transition: geometric



- Linder & Jenkins show the change in growth rate versus redshift.
- HETDEX is at $a=0.25$, where the effect is very strong (from $z=0$)
- WL growth estimates measures only integrated effect of $g(z)$, whereas HETDEX measures it a particular z (evolution of w)
- HETDEX will get $P(k)$ normalization to $<1\%$, and $g(z)$ to 2%

Comparison of various DE projects (for $w=w_0+wa[1-a]$)
Curvature assumption is very important for HETDEX (high- z)



Systematics, Systematics, Systematics

- Koehler, Schuecker, Gebhardt (2006) show the power of BAO using an essentially assumption-free approach
- No other technique can compare with BAO in terms of understanding systematics: SN are problematic at the few percent level, CL needs mass estimates to 5%, WL who knows?
- A great advantage is that the accuracies predicted from BAO are very well understood: ratio of optimistic to pessimistic is 1.05 for BAO, 1.8 for SN, 2.0 for CL, over 6 for WL
- Thus, we can confidently predict what it will take to reach 3x and 10x current accuracies.

Advantages of HETDEX are:

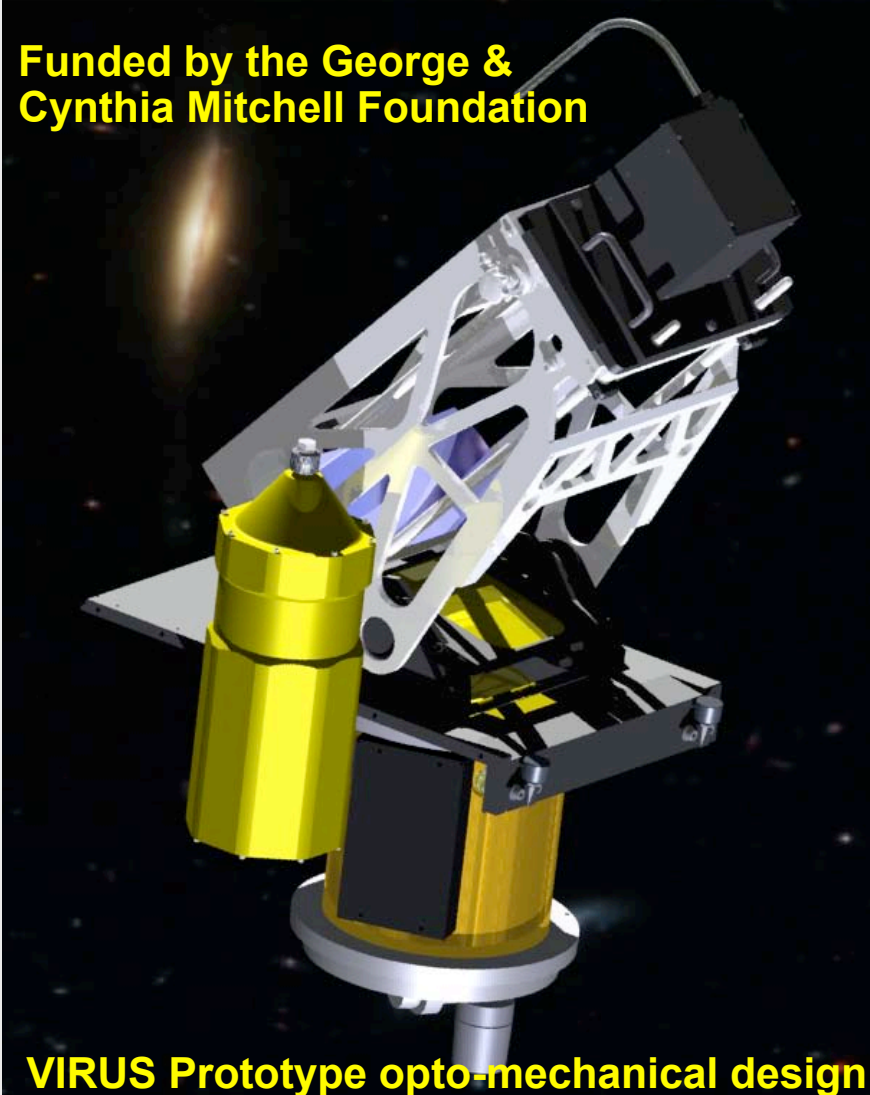
- **fastest Stage III program (needed for Stage IV planning)**
- **lowest systematics, because of BAO and high-z**
- **only experiment with constraints on early-time evolution, both from high-z estimate of distance and structure growth**
- **can be designed in phases**

HETDEX experimental requirements

- A LAE DE survey reaching 1% precision requires the following
 - large volume to average over cosmic variance
 - » 200-500 sq. degrees and $\Delta z \sim 2$
 - » this is 6-15 Gpc³ at $z \sim 2-4$
 - ~ 1 million galaxies tracing LSS in the volume
 - » surface density ~3000 per sq. degree per $\Delta z=1$
 - » LAEs have 18,000 /sq. deg./ $\Delta z=1$ at line flux $\sim 1e-17$ erg/cm²/s
 - lowest possible minimum redshift (bluest wavelength coverage)
 - » $z = 1.8$ at $\lambda 3400$ Å is a practical limit
 - » ties in well with high redshift limit of SNAP and other experiments
- These requirements can be met by a wide field IFU spectrograph
 - IFS obtains a spectrum of all spatial elements in the field of view, simultaneously
 - such an instrument is perfectly suited to the HET

HETDEX technical requirements

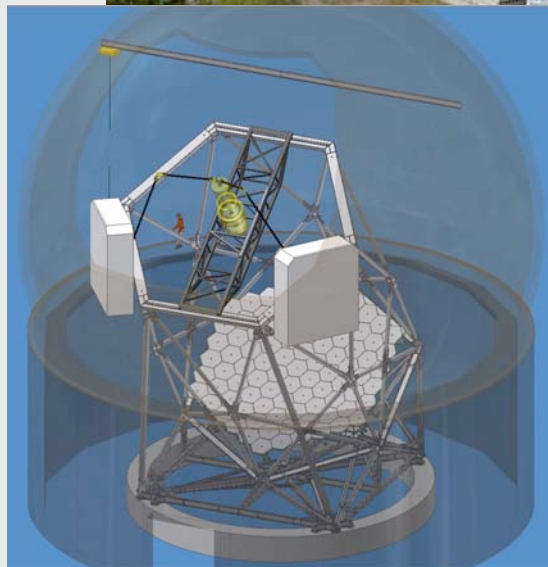
- **Visible Integral-field Replicable Unit Spectrograph**
 - Find LAEs and measure redshifts in one observation
- VIRUS is the prototype of the industrial replication concept
 - Each unit spectrograph is well within the state of the art
 - » each gives 0.25 sq. arcmin. and 340-570 nm wavelength range, $R=850$
 - ~145 VIRUS will cover
 - » 30 sq. arcminutes per observation
 - » Detect 14 million independent resolution elements per exposure
 - An order of magnitude more powerful than any existing spectrograph
- Prototype nearly ready for observing
 - Will be used for large pilot survey starting in November



VIRUS on the HET

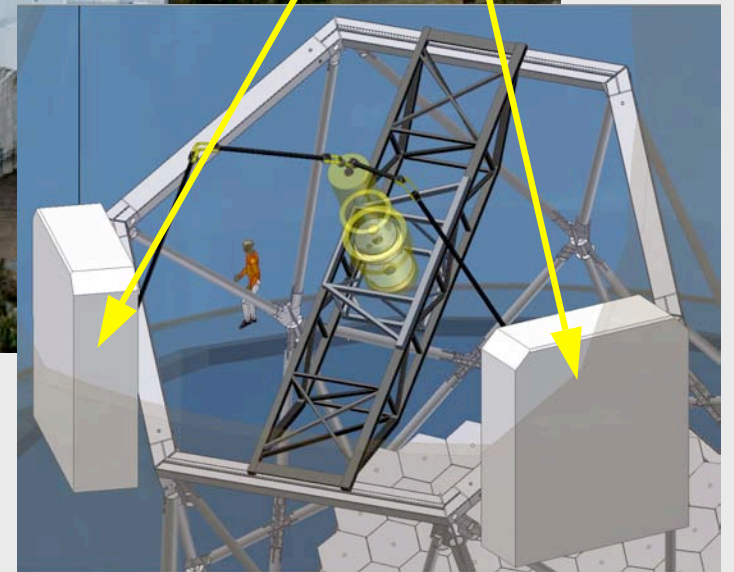


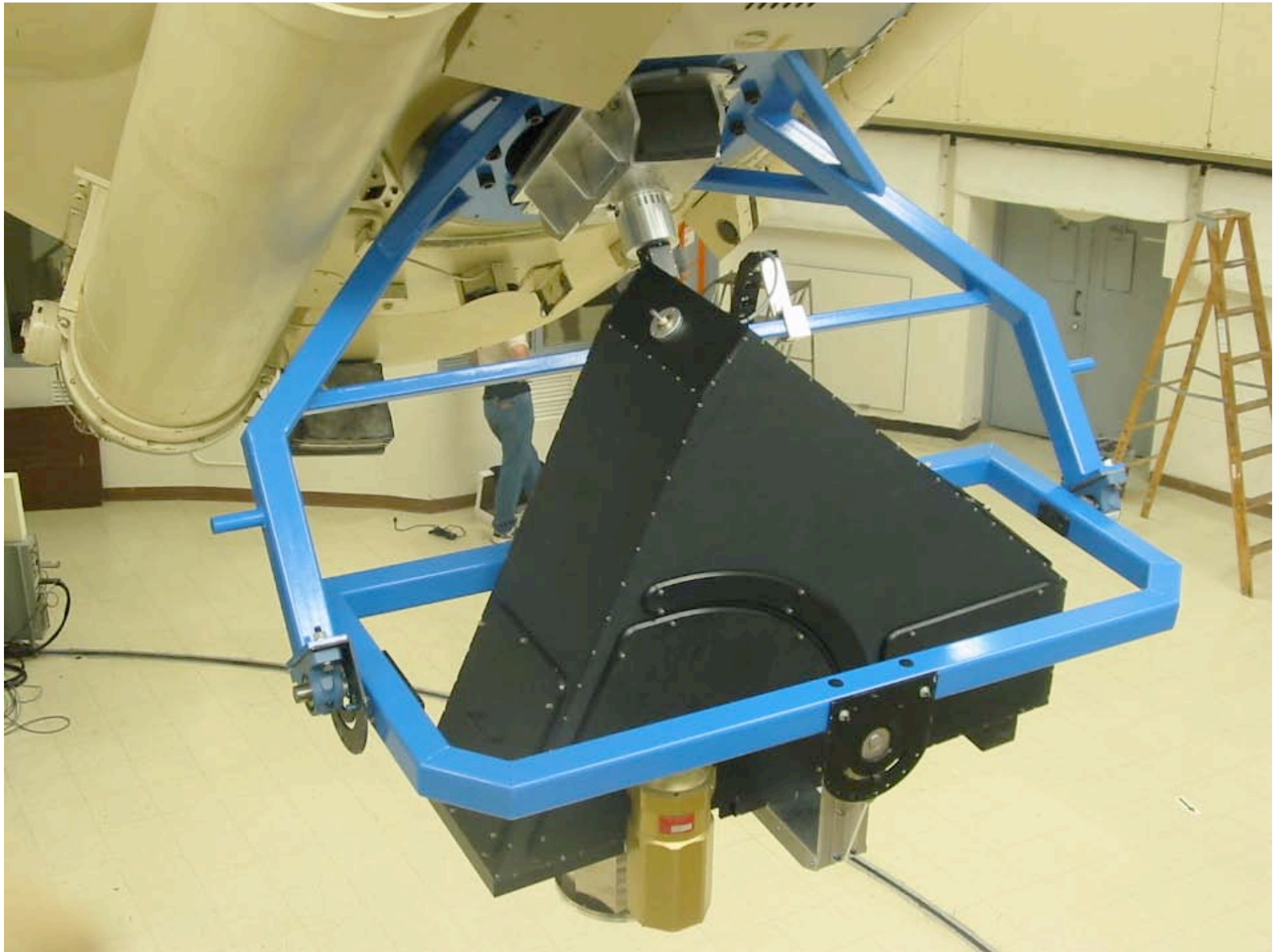
VIRUS consists of 145 units mounted on HET

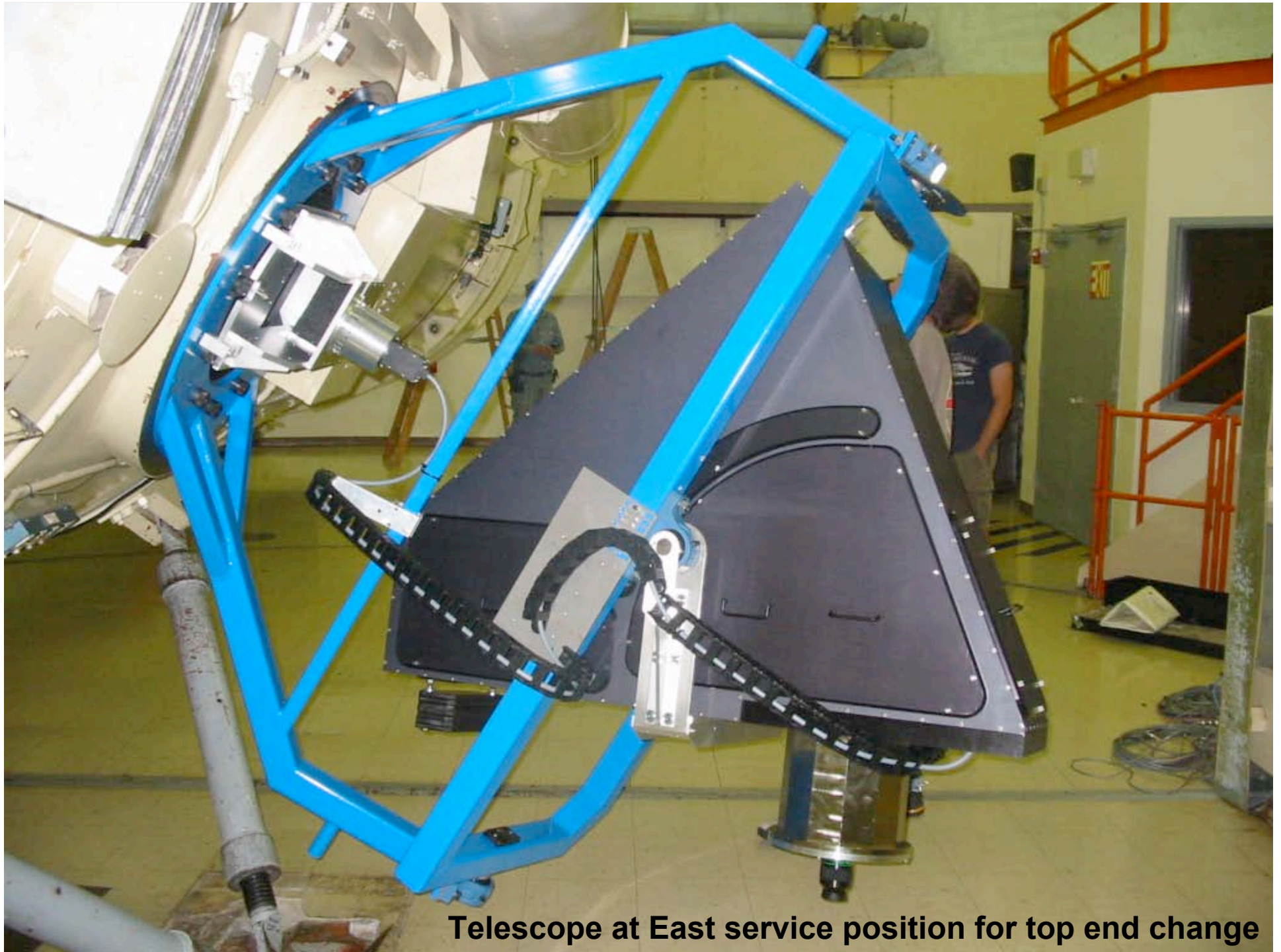


9.2 m HET
Mt. Fowlkes west Texas

New wide field corrector
New Tracker







Telescope at East service position for top end change

HETDEX Status

- **VIRUS-P pilot survey**
 - to verify performance, test the engineering
 - to tell us the remaining information we need to plan HETDEX
- **Total cost ~\$33M**
 - Upgrading HET to 22 arcminute diameter field of view
 - Building VIRUS
 - Observing and analyzing the data
 - ~\$15M raised and identified
- **Timeframe is 3 years construction, 3 years observing**
 - ~100 clear dark nights
 - Completion in 2012/2013



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