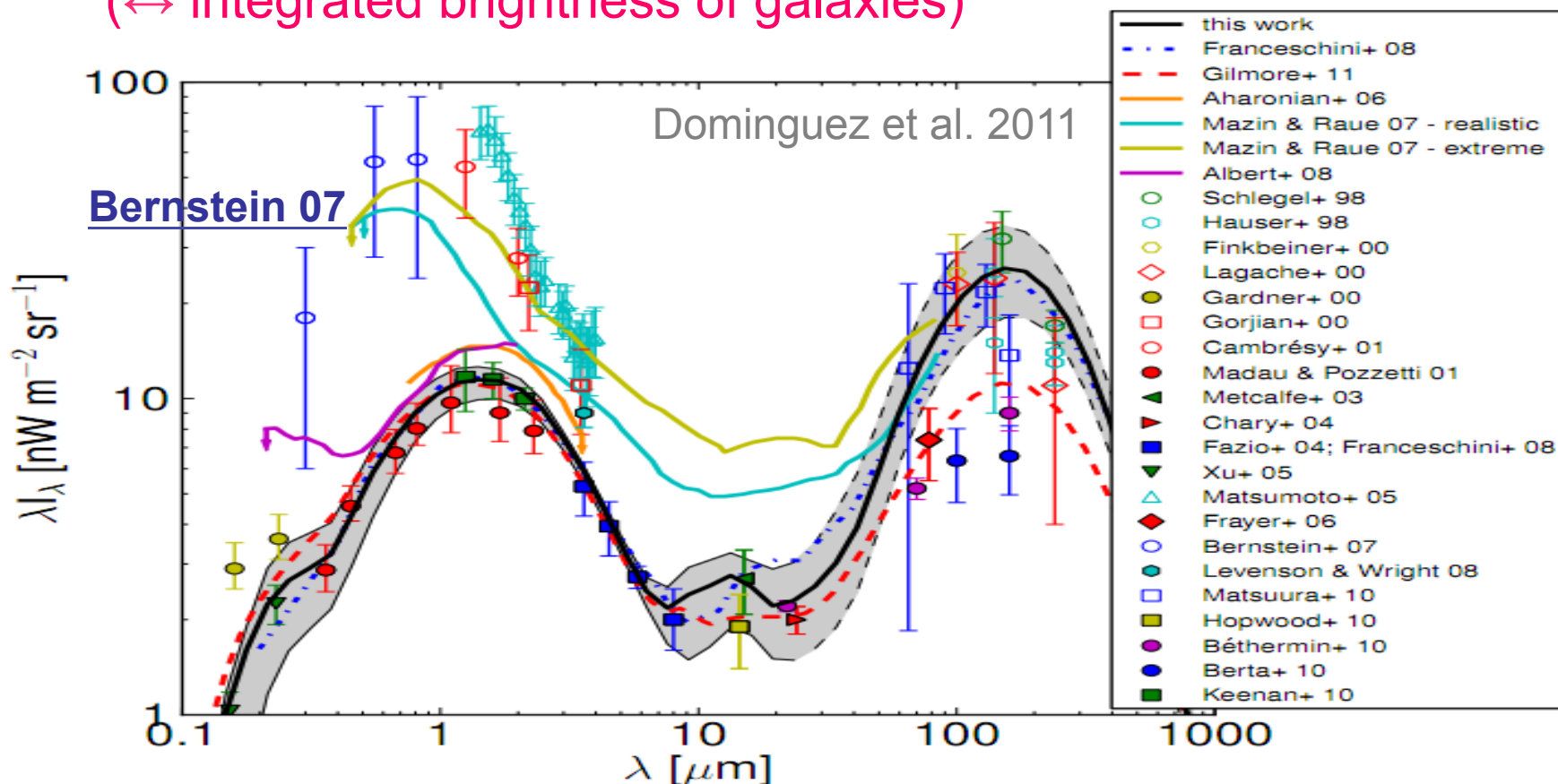


Abstract

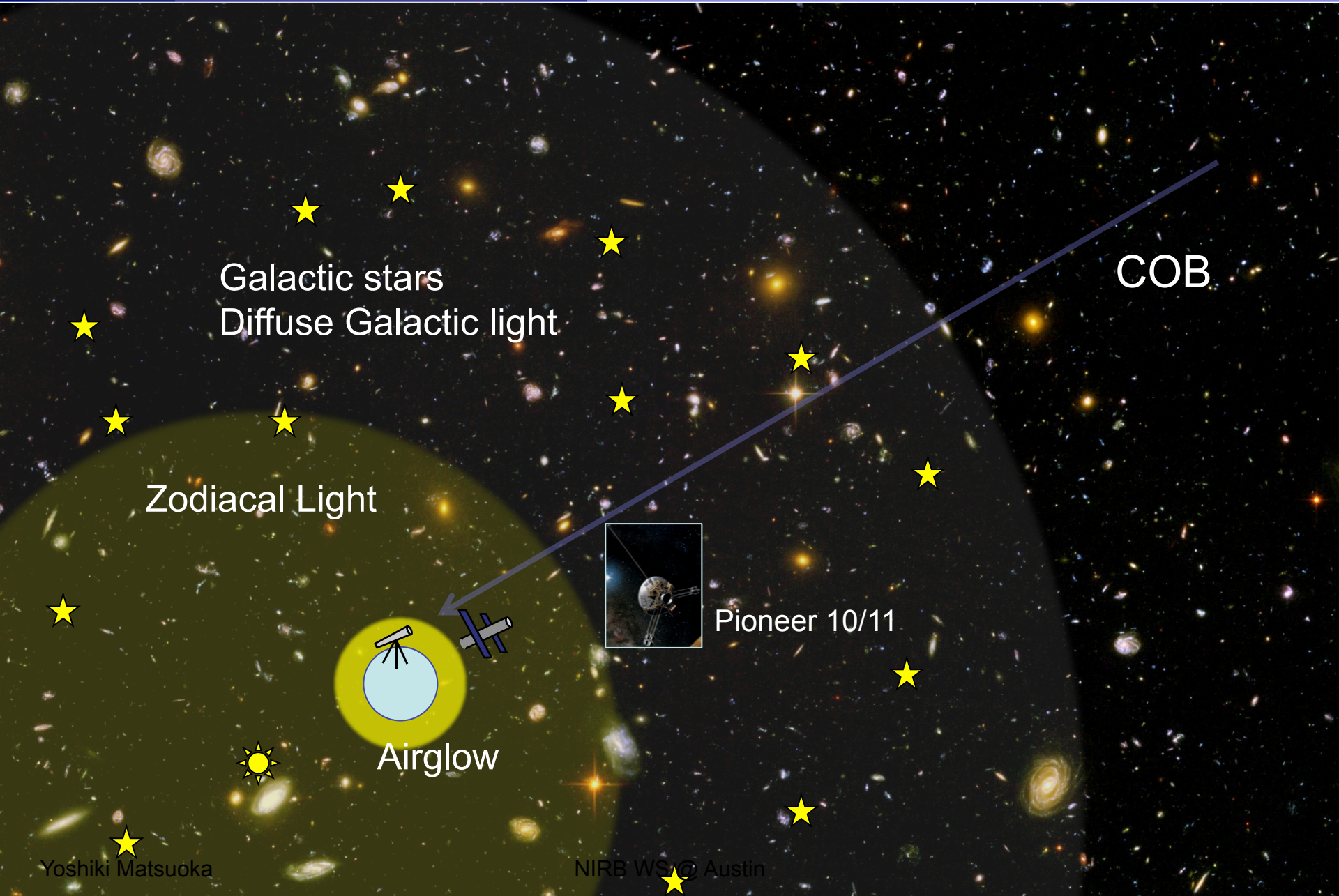
- We have re-analyzed the all-sky imaging data taken by the IPPs on board the Pioneer 10/11 spacecrafts.
- The new constraints on the COB are obtained, which are
 - $7.9 \pm 4.0 \text{ nW m}^{-2} \text{ sr}^{-1}$
at $0.39 - 0.50 \text{ }\mu\text{m}$,
 - $7.7 \pm 5.8 \text{ nW m}^{-2} \text{ sr}^{-1}$
at $0.60 - 0.72 \text{ }\mu\text{m}$.
- The derived COB is consistent with the integrated brightness of galaxies in the Hubble deep field.

COB: The View from Pioneer 10/11

- Cosmic optical background (COB) = Optical component of the extragalactic background light (EBL)
- UV and optical light of all radiation sources in the Universe
(\leftrightarrow integrated brightness of galaxies)

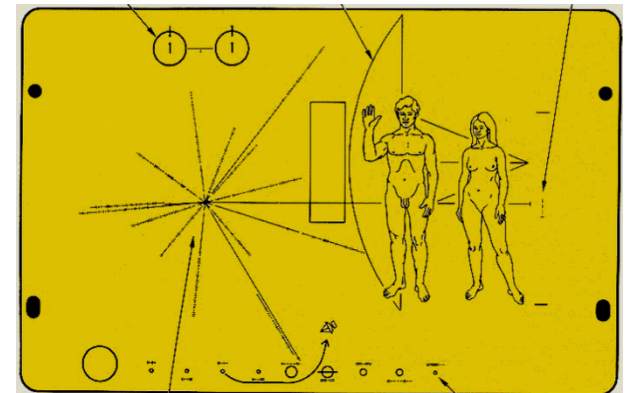


COB: The View from Pioneer 10/11



COB: The View from Pioneer 10/11

- NASA's Pioneer 10/11 spacecrafts
- “the first to be sent to the outer solar system and the first to investigate the planet Jupiter, after which followed an escape trajectory from the solar system”
- Launched: May 1972 (Pioneer 10)
Apr 1973 (Pioneer 11)
- Comm. Stop: Jan 2003 (P10; D = 82 AU)
Nov 1995 (P11; D = 45 AU)
- Scientific instruments:
magnetometer; plasma analyzer; charged particle detector; ionizing detector; non-imaging telescopes; sealed pressurized cells of Ar and N gas; UV photometer; IR radiometer; **imaging photopolarimeter**



COB: The View from Pioneer 10/11

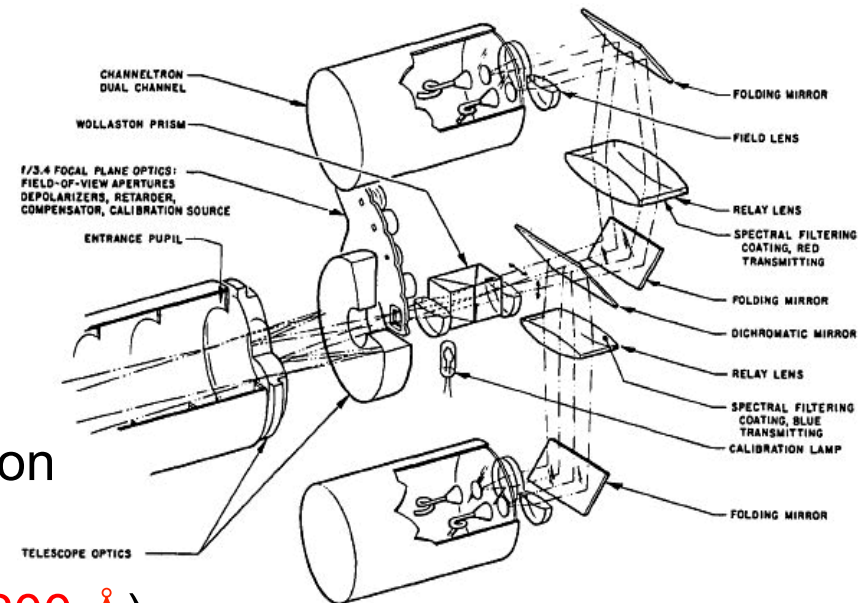
- Imaging Photopolarimeter (IPP)

- 2.5-cm Maksutov telescope
- Wollaston prism
- multilayer filters
- dual-channel Bendix channeltrons

- measures two orthogonal polarization components in the two wave bands

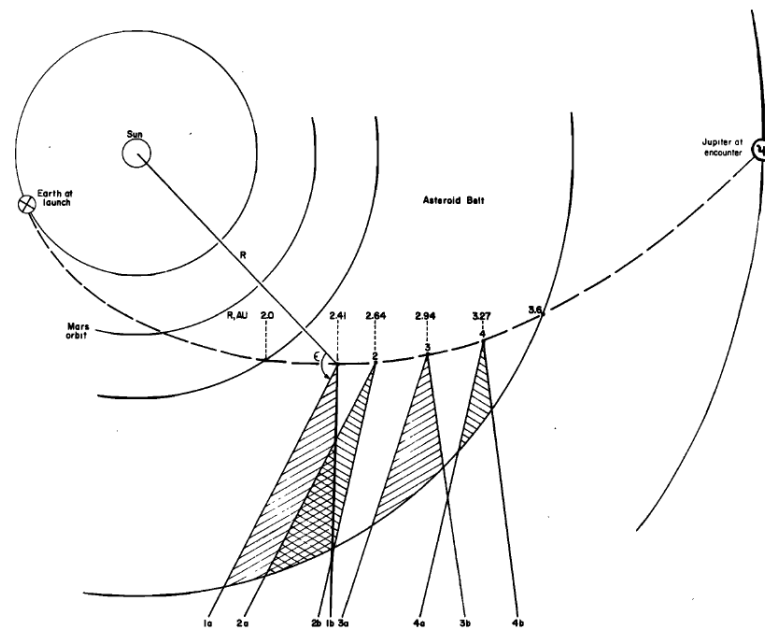
(blue; 3900 – 5000 Å, red; 5950 – 7200 Å)

- Instantaneous FOV: $2.29^\circ \times 2.29^\circ$
- takes 64 exposures per one spacecraft spin (12.5 sec)
 - effective FOV: $2.29^\circ \times (2.29^\circ + 5.625^\circ \sin L) \sim 10 \text{ deg}^2$
- L : “look angle” between the IPP pointing and spacecraft spin axis
- 1 data cycle = 10 rolls (8 for sky measurements, 1 for photometric calibration, 1 for dark-current and offset measurement)



COB: The View from Pioneer 10/11

- ZL and COB measurements by the IPP
- Hanner et al. (1974)
 - monitored the sky brightness during the cruise phase of the Pioneer 10 at the heliocentric distances 2.4 – 4.8 AU.
 - ZL @ 2.4 AU < 10% ZL @ 1 AU
 - ZL undetectable @ > 3.3 AU
- Toller (1983)
 - attempted to detect the COB in the Pioneer 10 IPP data taken at the heliocentric distances > 3.3 AU (i.e., outside the ZL clouds)
 - COB < $4.5 \times 10^{-9} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ sr}^{-1} \text{ \AA}^{-1}$ at 4400 \AA
 - ... comparable to the HST results by Bernstein et al. (2002-2007)
 - BUT the starlight subtraction is the fatal problem in his analysis.



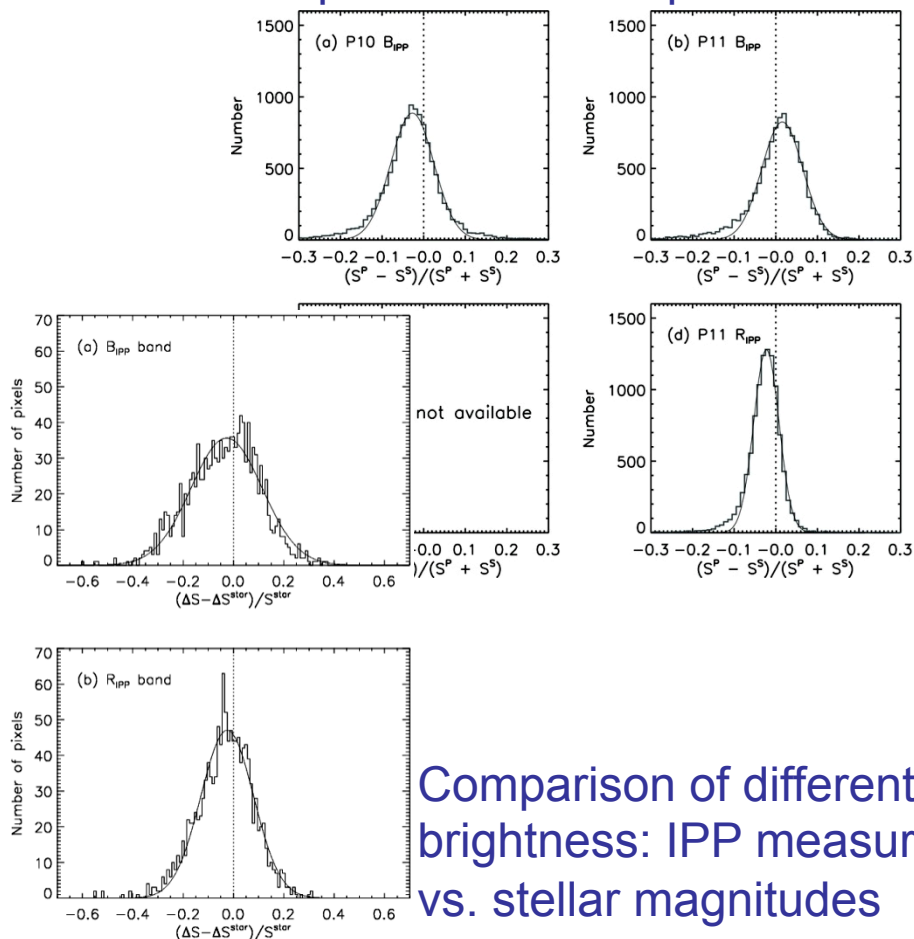
COB: The View from Pioneer 10/11

space craft	date	R (AU)	all #	good Q	corrupt	scatter. sunlight	abnorm. (global)	abnorm. (local)
Pio 10	354/'72	3.26	5696	60 %	25 %	5 %	7 %	3 %
Pio 10	093/'73	3.92	1344	75 %	5 %	0 %	14 %	6 %
Pio 10	149/'73	4.22	5504	61 %	19 %	10 %	5 %	6 %
Pio 10	216/'73	4.54	2816	51 %	3 %	0 %	2 %	45 %
Pio 10	237/'73	4.64	5312	57 %	24 %	5 %	12 %	3 %
Pio 10	279/'73	4.81	5248	57 %	21 %	8 %	10 %	4 %
Pio 10	021/'74	5.08	5504	59 %	27 %	4 %	6 %	4 %
Pio 10	068/'74	5.15	5376	67 %	29 %	1 %	1 %	2 %
Pio 11	057/'74	3.50	4544	73 %	11 %	1 %	10 %	4 %
Pio 11	106/'74	3.81	4672	74 %	5 %	6 %	10 %	4 %
Pio 11	148/'74	4.06	4672	77 %	3 %	9 %	9 %	2 %
Pio 11	178/'74	4.22	4608	73 %	3 %	9 %	11 %	4 %
Pio 11	236/'74	4.51	4544	82 %	10 %	3 %	1 %	5 %
Pio 11	267/'74	4.66	4416	71 %	9 %	3 %	6 %	12 %

COB: The View from Pioneer 10/11

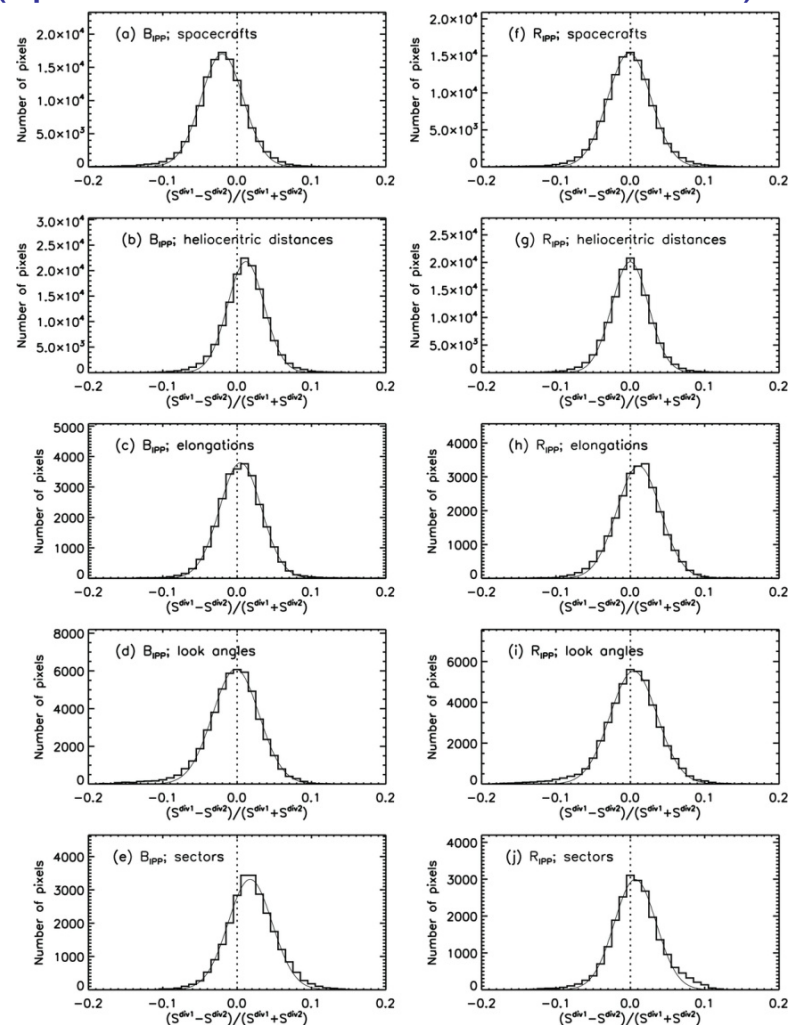
- Quality Assessment

Comparison between two polarization components



Comparison of differential brightness: IPP measured vs. stellar magnitudes

Comparison between data subgroups (spacecraft, distance to the Sun, ...)



COB: The View from Pioneer 10/11

S: Brightness measured by the IPP

S^{star} : Brightness from the independent star catalogs

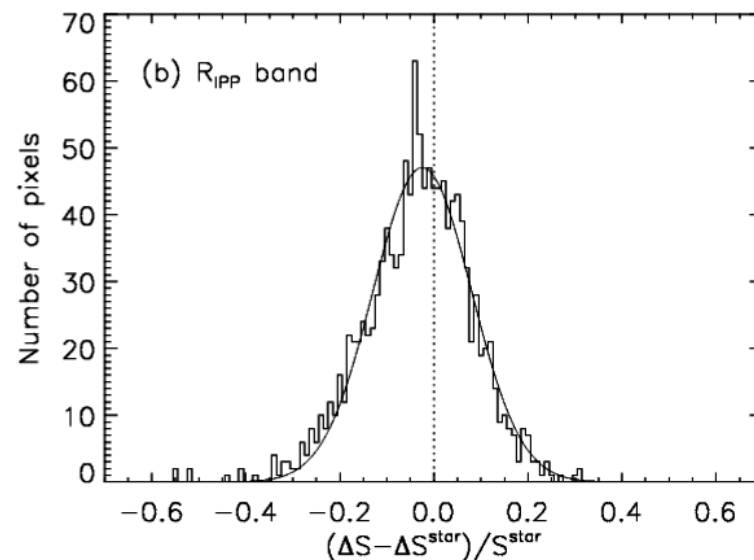
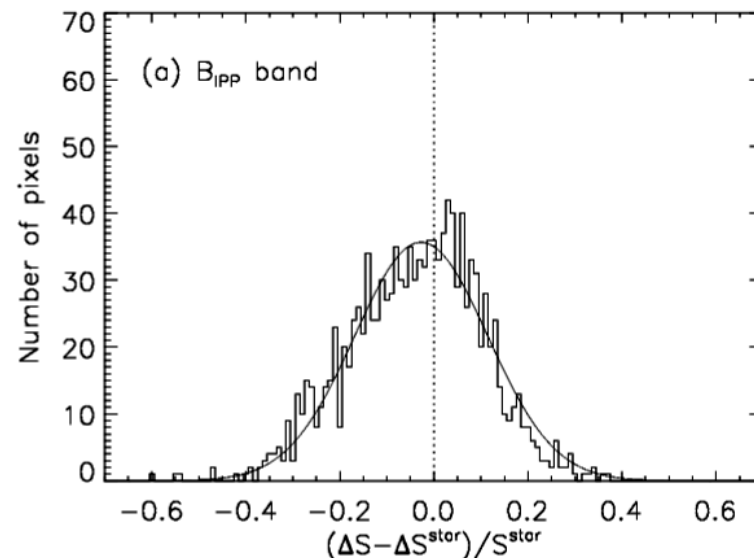
- S and S^{star} cannot be directly compared because (only) S includes the COB component.



ΔS : Brightness difference between any two IPP FOVs.

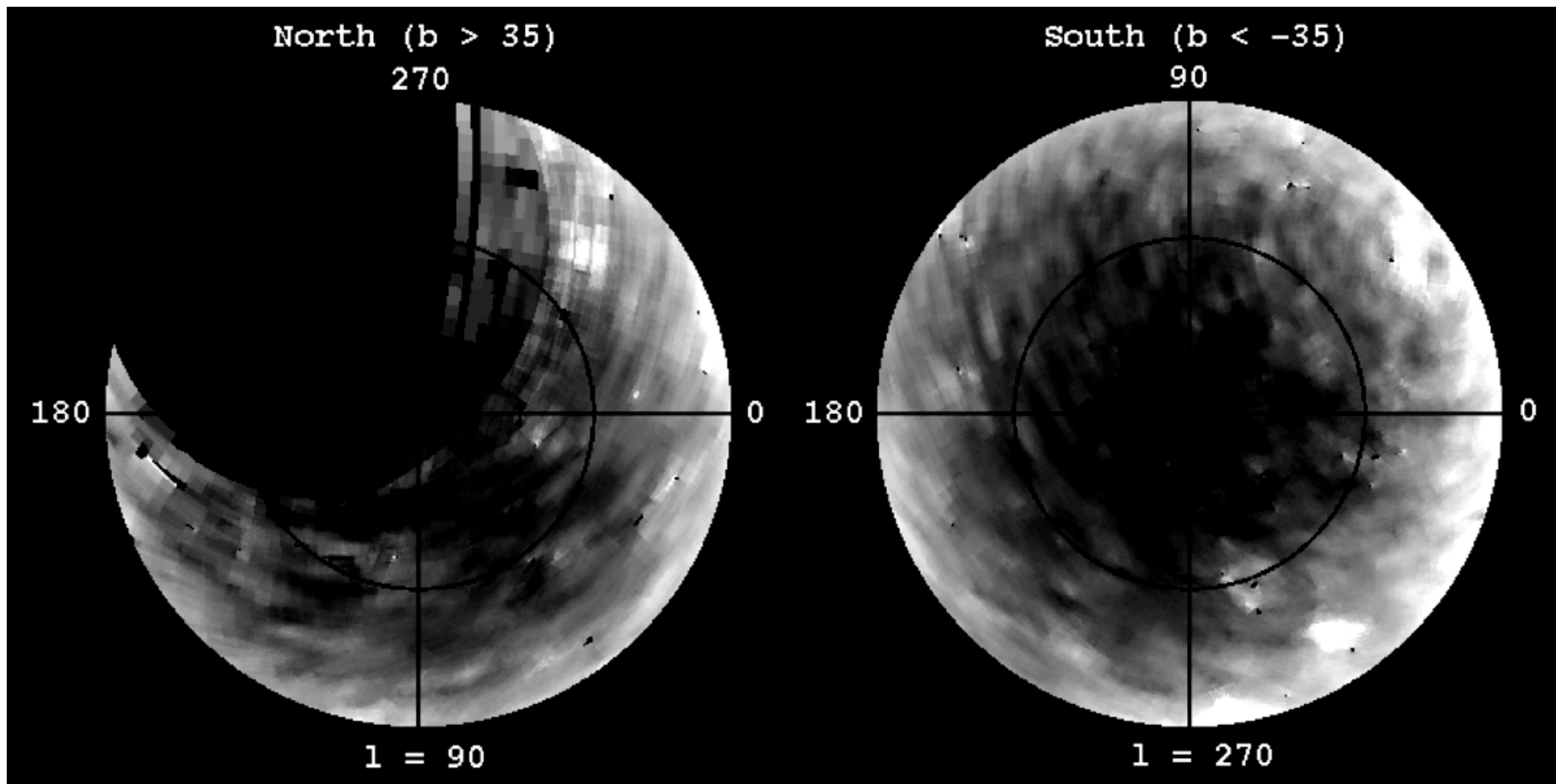
ΔS^{star} : Corresponding difference from the star catalogs

.. a few % systematic difference



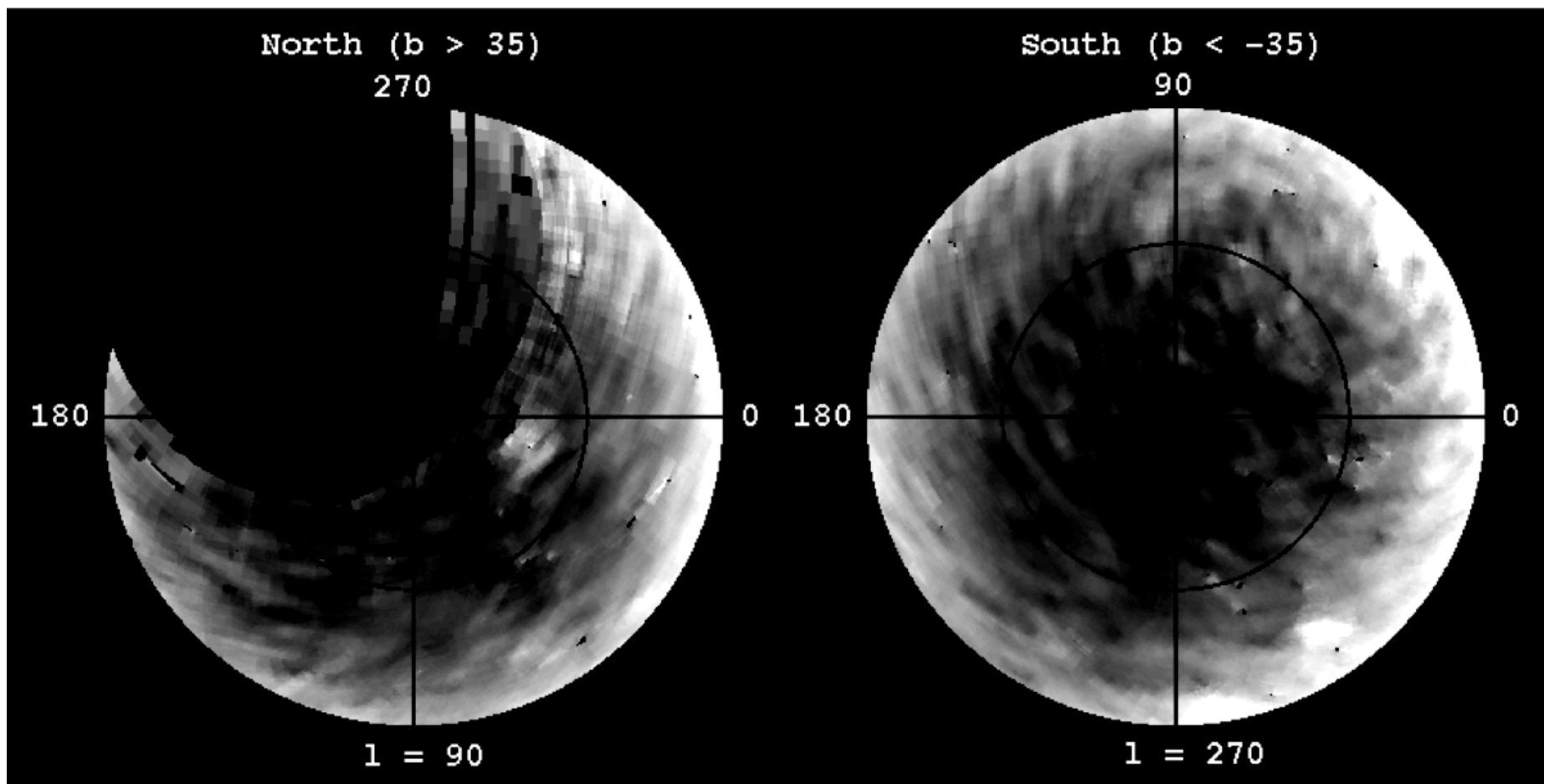
COB: The View from Pioneer 10/11

- IPP Blue-band brightness maps (north/south Galactic hemispheres at $|b| > 35^\circ$)



COB: The View from Pioneer 10/11

- IPP Red-band brightness map (north/south Galactic hemispheres at $|b| > 35^\circ$)



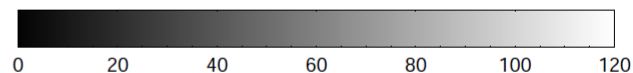
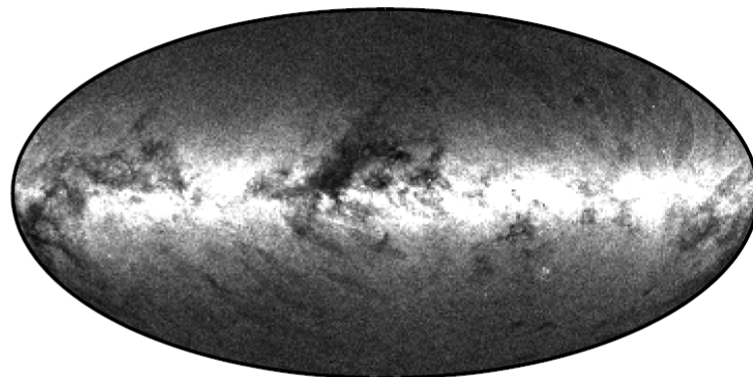
COB: The View from Pioneer 10/11

- Contribution of faint stars
 - Bright ($m_v < 6.5$) stars in the Yale Bright Star Catalog and the USNO Photoelectric Catalog have already been subtracted.
- Fainter stars... all-sky catalogs:
 - Tycho-2 Catalog (6 – 10 mag), - GSC-II Catalog (9 – 20 mag) and Galactic star-count model: - TRILEGAL (> 20 mag) are used to derive the starlight contributing to each IPP FOV



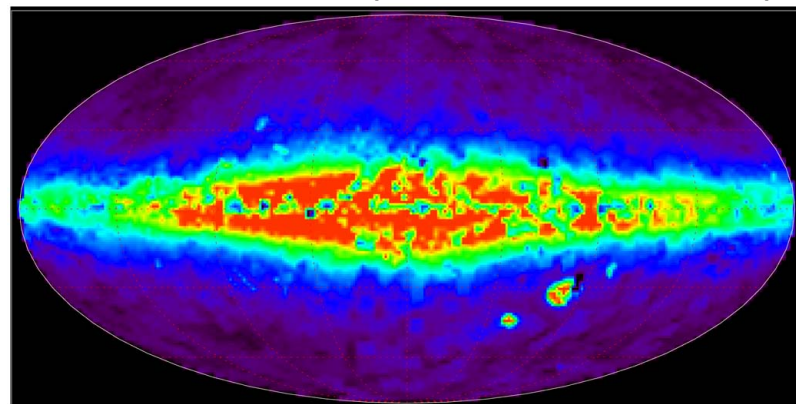
- “Diffuse emission map”

Tycho-2 stars (Høg et al. 2000)



stars per square degree

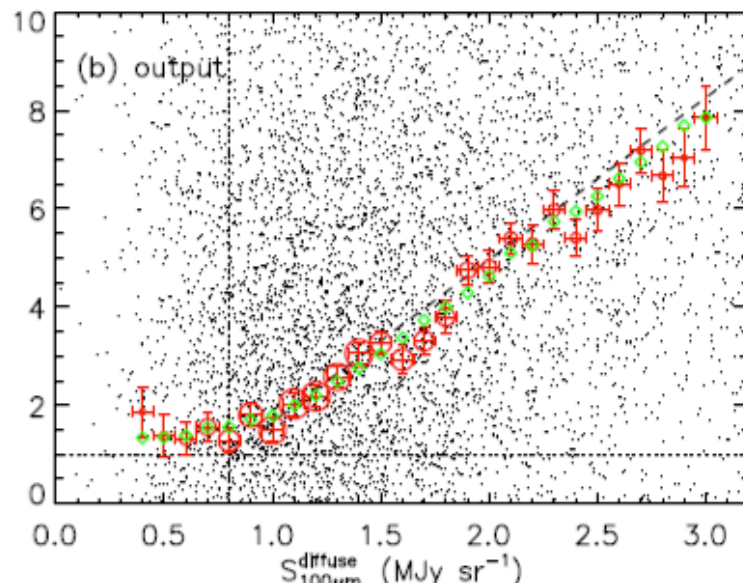
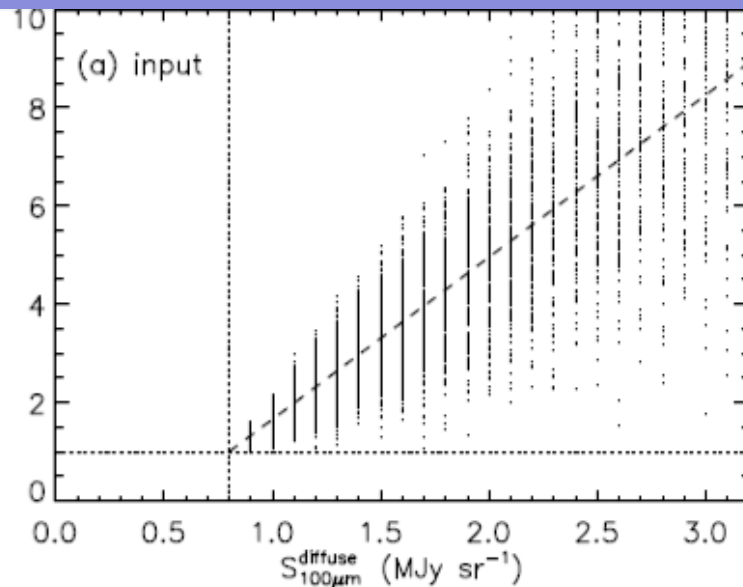
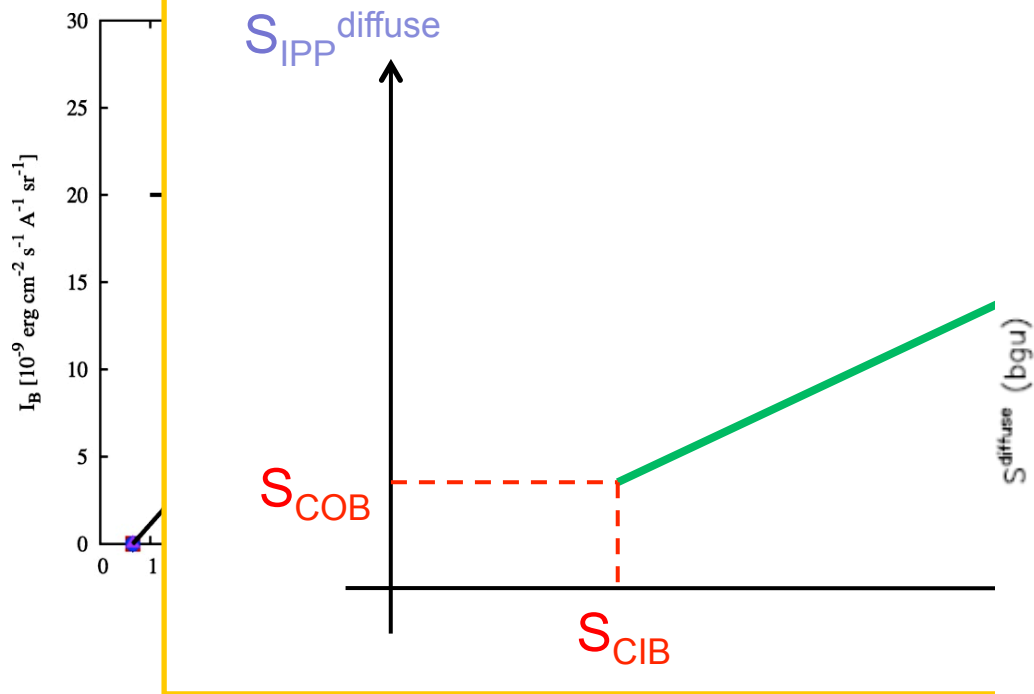
GSC-II stars (Lasker et al. 2008)



COB: The View from Pioneer 10/11

- F
- “S
- in
- W
- S

$$\begin{aligned}
 S_{\text{IPP}}^{\text{diffuse}} &= S_{\text{IPP}} - S_{\text{IPP}}^{\text{S}} \\
 &= S_{\text{DGL}} + S_{\text{COB}} \\
 &= a_{\text{DGL}} (S_{100\mu\text{m}}
 \end{aligned}$$



COB: The View from Pioneer 10/11

- DGL subtraction

IPP diffuse emission map at $|b| > 35^\circ$



IRAS/DIRBE 100 μm emission map
(Schelgel et al. 1998)

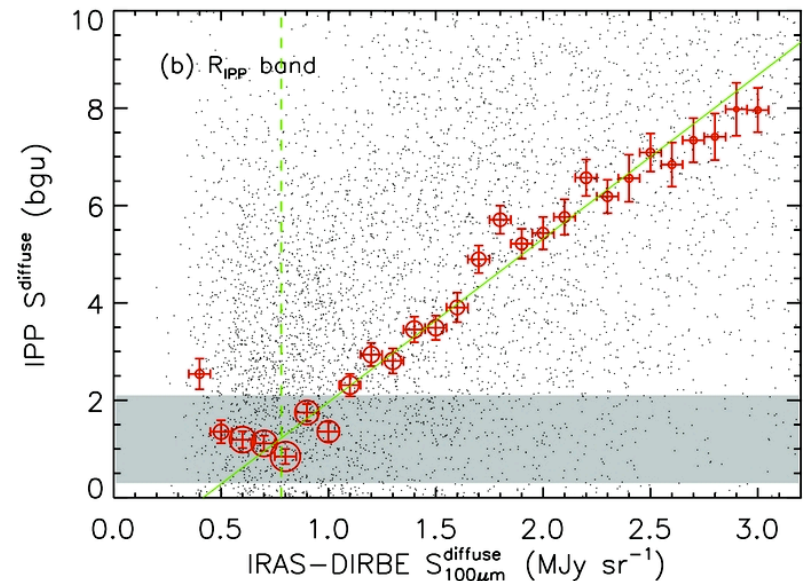
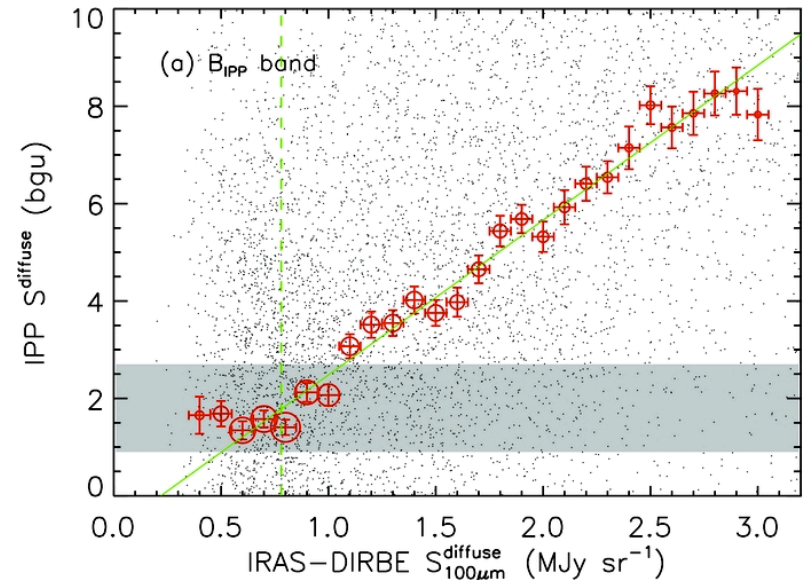
- Derived DGL/100 μm brightness ratios are in good agreement with the previous measurements.

- Residual COB

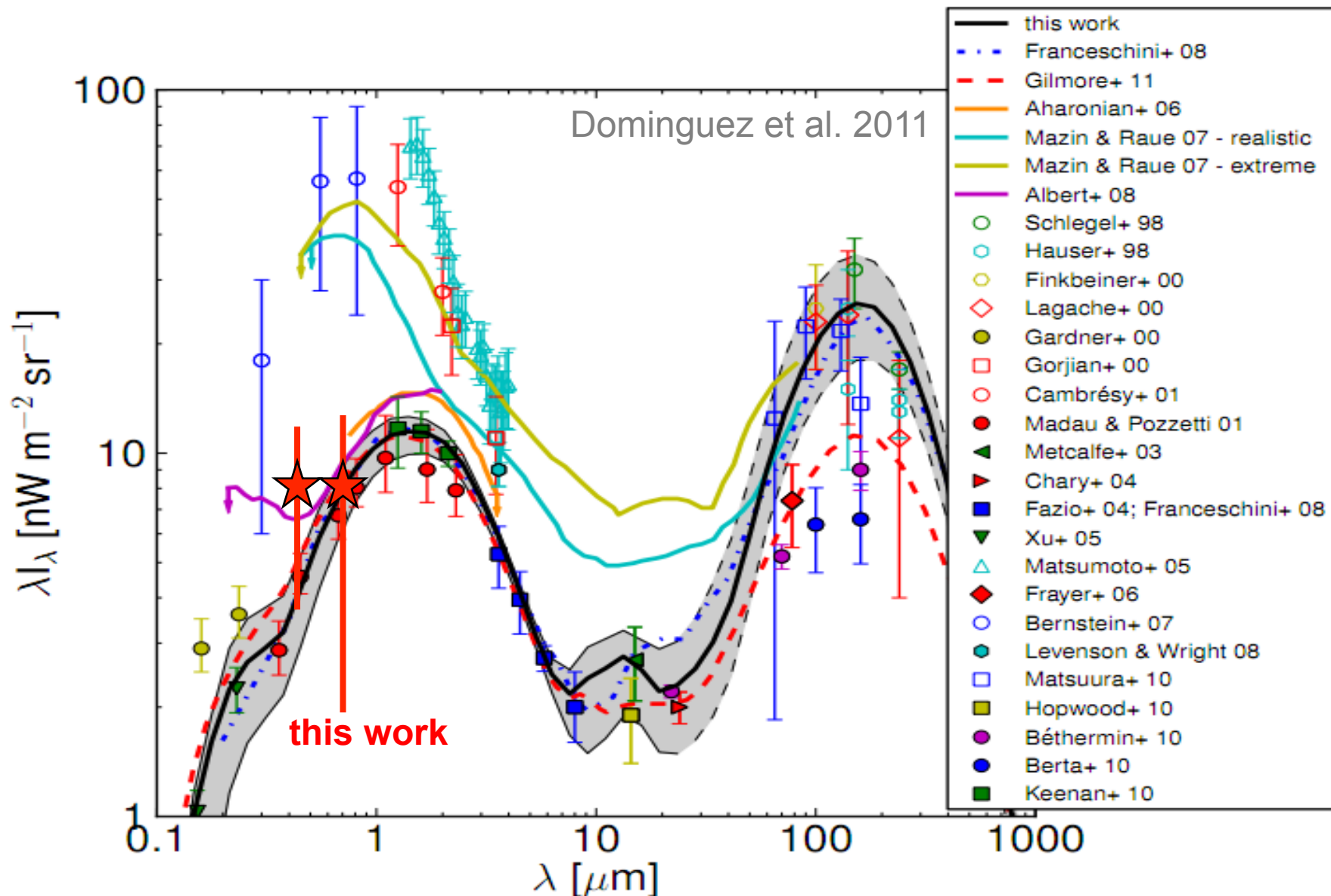
- $7.9 \pm 4.0 \text{ nW m}^{-2} \text{ sr}^{-1}$ at Blue band

- $7.7 \pm 5.8 \text{ nW m}^{-2} \text{ sr}^{-1}$ at Red band

($I_{\text{CIB}} = 0.78 \pm 0.21 \text{ MJy sr}^{-1}$
at $100\mu\text{m}$; Lagache et al. 2000)



COB: The View from Pioneer 10/11



Summary

- We have re-analyzed the all-sky imaging data taken by the IPPs on board the Pioneer 10/11 spacecrafts.
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 - $7.9 \pm 4.0 \text{ nW m}^{-2} \text{ sr}^{-1}$
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at $0.60 - 0.72 \text{ }\mu\text{m}$.
- The derived COB is consistent with the integrated brightness of galaxies in the Hubble deep field.
- Reference: Matsuoka et al. 2011, ApJ, 736, 119