

# Alignment Between Flattened Protostellar Infall Envelopes and Ambient Magnetic Fields

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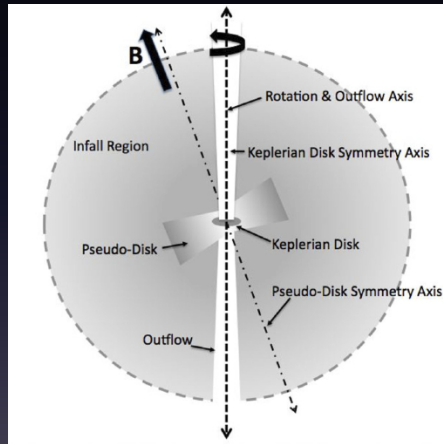
J. A. Davidson, P. F. Goldsmith, M. Houde, W. Kwon,  
Z.-Y. Li, L. W. Looney, B. Matthews, T. G. Matthews,  
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# Magnetically Regulated Core Collapse (e.g. Allen et al. (2003a,2003b))

1. Existence of Pseudodisk (flattened infall envelope)
2. Magnetic field pinched within infall region
3. Pseudodisk axis  $\parallel$  to magnetic field axis
4. (Outflow  $\parallel$  to  $B$ -field and pseudodisk)
  - ▶ cf. Joos et al. (2012), Hull et al. (2013)



Davidson et al. (2011)

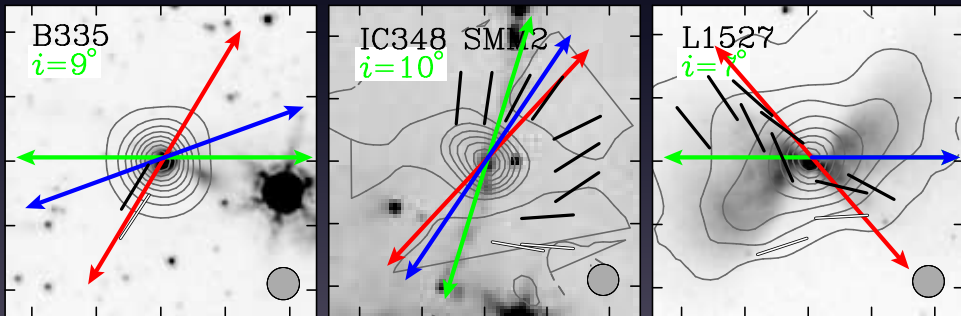
# SHARP

- SHARP operates at the CSO
- SHARP adds polarimetric capability to SHARC-II camera
  - ▶ Simultaneous Horizontal and Vertical polarization
  - ▶  $350\ \mu\text{m}$  and  $450\ \mu\text{m}$
- SHARP Survey of Low-Mass Protostellar Cores
  - ▶ We chose simple, easy-to-understand sources
  - ▶ Nearby, single, isolated, inclinations near to plane of sky(\*)



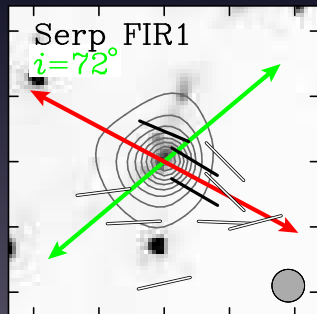
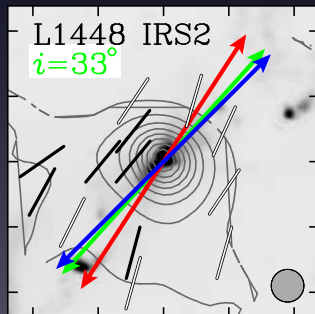
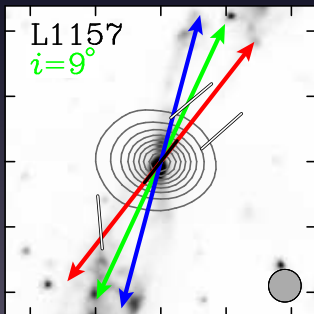
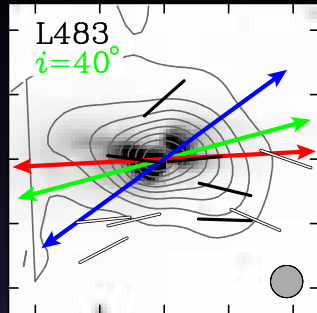
# SHARP Survey of Low-Mass Protostellar Cores

- Paper I: Davidson et al. (2011)
  - ▶ B335, IC348 SMM2, L1527
  - red - mean magnetic field
  - green - outflow axis
  - blue - pseudodisk symmetry axis



# SHARP Survey of Low-Mass Protostellar Cores

- Paper II: Chapman et al. (submitted)
    - ▶ L483, L1157, L1448-IRS2, Serp FIR1
- red - mean magnetic field  
green - outflow axis  
blue - pseudodisk symmetry axis



# SHARP Survey of Low-Mass Protostellar Cores

## Remarks

- The **blue** (pseudodisk symmetry axis) and **green** (outflow axis) are very well correlated.
  - ▶ Mean difference in angle is  $12^\circ$
- **blue** (pseudodisk) somewhat aligned with **red** (mean magnetic field direction)
  - ▶ less than  $45^\circ$  in 6/7 sources
  - ▶ 7th source (Serp FIR1) has high inclination

## Analyze

- ▶ *proxy 1*: Inclination of outflow = inclination of pseudodisk
- ▶ *proxy 2*: Outflow angle = pseudodisk angle in Serp FIR1

## Future

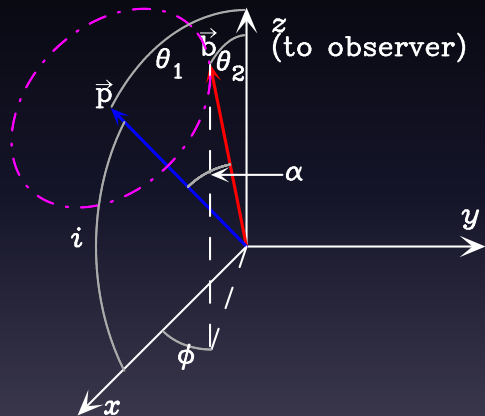
- Unreduced data: L1551-NE, HH211-MM, B1-C

# Importance of Inclination Angle

- Projected Separation  $\phi$  is a function of:

- Inclination  $i$
- 3D separation  $\alpha$
- Rotation  $\phi'$

$$\tan \phi = \frac{\tan \alpha \sin \phi'}{\sin \theta_1 + \cos \theta_1 \tan \alpha \cos \phi'}$$

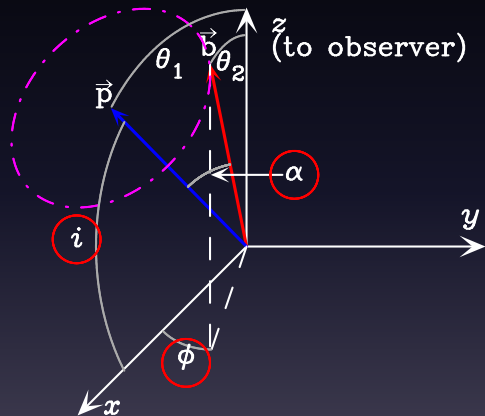


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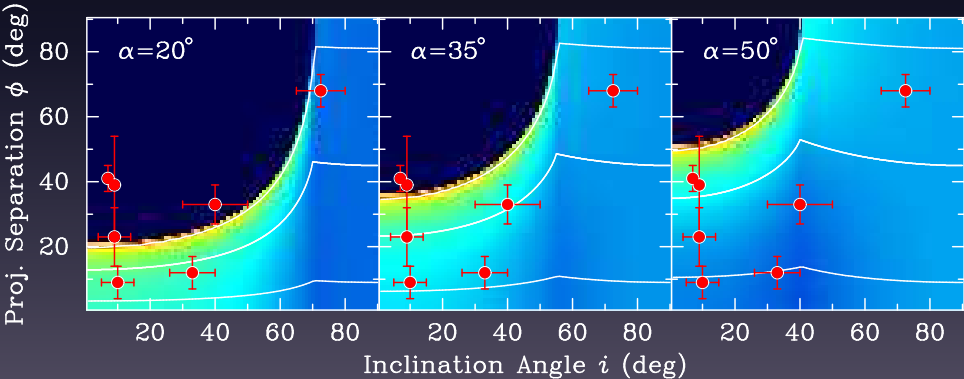
$$\tan \phi = \frac{\tan \alpha \sin \phi'}{\sin \theta_1 + \cos \theta_1 \tan \alpha \cos \phi'}$$





# Results

- Compute projected separation between magnetic field and pseudodisk for all 7 sources
- From simple chi-squared minimization best-fit  $\alpha = 36.1^\circ$
- We estimate the probability of obtaining  $\alpha \leq 36.1^\circ$  by pure chance is 5%
- Folding in errorbars,  $\alpha < 40^\circ$



# Combining Sources

- Each source has only a few vectors  $\geq 2\sigma$
- Each source has other vectors of lesser significance

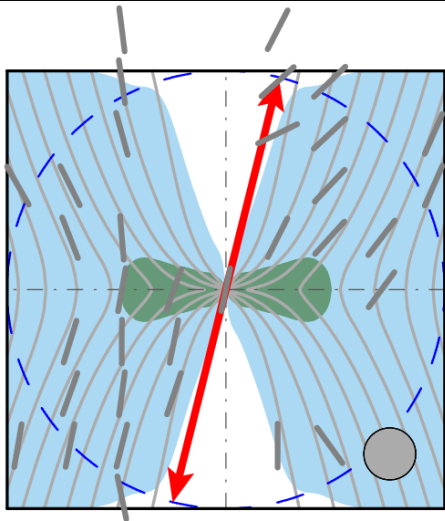
Why not utilize *ALL* the information on each source?

## Compare with magnetically regulated core collapse model

- Ignored Serp FIR1 (high outflow inclination)
- Rotated each source so pseudodisk was horizontal
- Scaled each source by its infall radius
- Computed the source average map  
(weighted average of rotated, scaled maps)

# Source Average Map

- Mean field angle  $166^\circ$
- Hints of a pinch
- 35 vectors
  - ▶  $13 \geq 3\sigma$
  - ▶  $22 \geq 2\sigma$

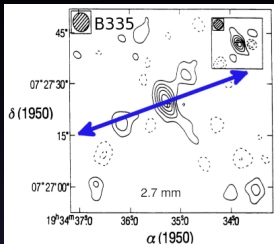


Allen et al. (2003b)

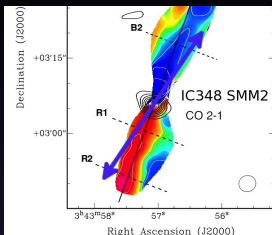
# Summary

- We have now observed 7 low-mass, isolated sources
- The outflow axis is well aligned with pseudodisk axis
  - ▶ mean difference is  $12^\circ$
- Pseudodisk preferentially aligned with  $B$ -field
  - ▶  $\alpha < 40^\circ$ ; 5% probability due to chance
- Combining sources into a single map improves S/N and supports results of magnetic field alignment with pseudodisk

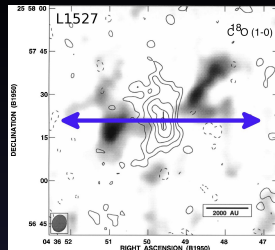
# Pseudodisks



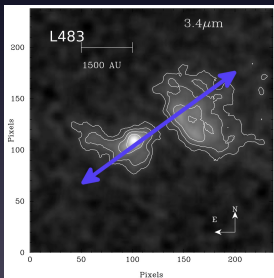
Chandler & Sargent (1993)



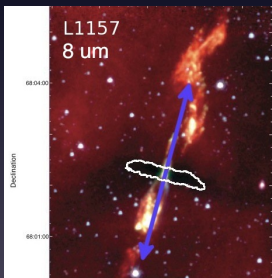
Pech et al. (2012)



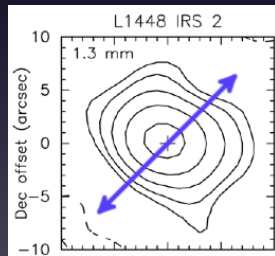
Ohashi et al. (1997)



Fuller & Wooten (2000)



Looney et al. (2007)



Kwon et al. (2009)

# Vector Separation

