



# **The Doppler Shift in Astronomy**

# Wavelength & Frequency

- any wave pattern has wavelength and frequency
  - wavelength is distance before pattern repeats
  - frequency is how many patterns per given time
- multiply the wavelength and frequency and you get the speed.
  - units: [distance]\*[1/time]=speed
  - large wavelength means small frequency

# Light & Sound

- both light and sound are wave phenomena
  - light is made of electromagnetic waves
  - sound is made of pressure waves
- for sound, high frequency means high pitch
  - middle C=261.6 Hz
  - A above middle C=440 Hz
- for light, high frequency means "bluer" light

# Source Moving toward Observer

- wave crests are closer so shorter wavelength (see more crests per second so higher frequency)
  - waves are blue-shifted (higher pitch)
- Observer moving toward source gives same result

# Source Moving away from observer

- wave crests are farther so longer wavelength (see fewer crests per second so lower frequency)
  - waves are red-shifted (lower pitch)
- Observer moving away from source gives same result

# Doppler Football

- Ingredients:

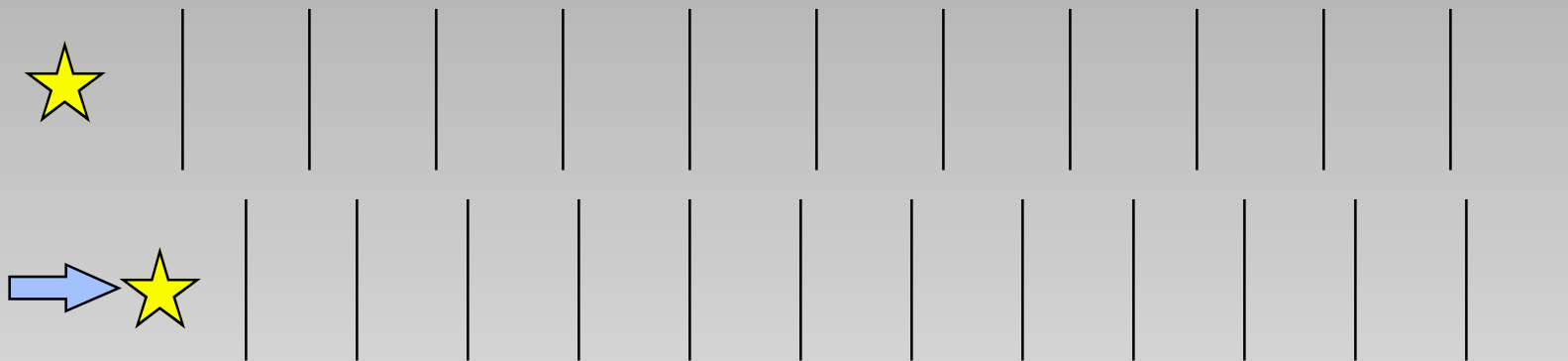
- ❑ Nerf type football (\$4)
- ❑ piezo buzzer (\$5)
  - pure tone 2.8 kHz
  - pulsing or continuous
  - no drive circuit needed
- ❑ 9V battery (?)
- ❑ 9V battery snap connection (\$1.40 for 5)
- ❑ Flat lever switch - single throw, 2 leads (\$2.80)
- ❑ 2 rubberbands

# Doppler Football

- Tools
  - ❑ wire strippers
  - ❑ Swiss army knife saw blade (cutting football)
  - ❑ electrical tape (to be safe)
- Connect one lead from buzzer to 9V (same color; strip wire as needed)
- Connect other lead from buzzer to switch
- Connect other lead from 9V to switch
  - ❑ tape as necessary
- Cut football enough to insert electronics
- Core hole in football for opening in buzzer
- Use rubberbands to “seal” football

# Doppler Shift

- Shift in observed wavelength/frequency is called the Doppler Shift



$$\lambda = c/f \quad \Rightarrow \quad \lambda c = 1/f$$

$$\lambda' = c/f + v/f$$

$$\lambda' = \lambda(1 + v/c)$$

$\lambda \Rightarrow$  original wavelength

$\lambda' \Rightarrow$  observed wavelength

$f \Rightarrow$  frequency

$c \Rightarrow$  speed of waves

$v \Rightarrow$  (vector) speed of object

# Source/Observer moving perpendicular to each other

- Distance between wave crests is unaffected
- No effect
  - only velocity along line of sight matters

# Uses of Doppler Shift

- Radar guns
  - ❑ send out radar waves
  - ❑ measure shift in return waves
- Stellar motion
  - ❑ take the spectrum of a star
  - ❑ compare observed wavelengths of absorption lines to lab values (H, Fe, Na, etc.)
  - ❑ calculate star's radial motion (need distance and tangential angular motion to get space motion)

# More uses of Doppler Shift

- Detecting planets around other stars
  - All objects orbit about the center of mass
    - stars with large planets have larger orbits
    - Jupiter-Sun center of mass is just outside Sun's radius
  - Watch for periodic Doppler shift as star orbits center of mass
    - Takes Jupiter ~12 years to orbit Sun
    - Sun's velocity is ~15 m/s
    - Difficult observation
    - Discoveries so far are massive planets close to the star

# Still More Uses of Doppler Shift

- The Expanding Universe
  - Because universe is expanding, galaxies further away are moving faster
    - Doppler shifted lines
  - We know wavelengths for lines of H, He, S, O (and their ions)
  - identify patterns in distant galaxy and deduce distance
- Shift is actually expansion of space
- Shift can be so large that relativity needs to be included

# 2 Ways EXES will use Doppler Shift

- Observing H<sub>2</sub>O lines
  - ❑ Even from SOFIA, Earth's H<sub>2</sub>O will block all light at certain wavelengths
  - ❑ Earth's orbit gives 30 km/s Doppler shift at maximum (Sun and target are moving, too)
  - ❑ The shift can bring source H<sub>2</sub>O lines away from Earth's
- Disks around young stars
  - ❑ H<sub>2</sub> rotating in disk
  - ❑ one side comes toward you, one away
  - ❑ double-peaked spectrum