

Course Outline

Unit

1. Introduction
2. Fundamental Observations
 - 2.1 The Night Sky Is Dark
 - 2.2 On Large Scales, the Universe Is Isotropic and Homogeneous
 - 2.3 Galaxies Show a Redshift Proportional to Their Distance
 - 2.4 The Universe Contains Different Types of Particles
 - 2.5 The Universe Is Filled with a Cosmic Microwave Background
3. Newton Versus Einstein
 - 3.1 Equivalence Principle
 - 3.2 Describing Curvature
 - 3.3 The Robertson-Walker Metric
 - 3.4 Proper Distance
4. Cosmic Dynamics
 - 4.1 The Friedmann Equation
 - 4.2 The Fluid and Acceleration Equations
 - 4.3 Equations of State
 - 4.4 Learning to Love Lambda
5. Single-Component Universes
 - 5.1 Evolution of Energy Density
 - 5.2 Curvature Only
 - 5.3 Spatially Flat Universes
 - 5.4 Matter Only
 - 5.5 Radiation Only
 - 5.6 Lambda Only
6. Multiple-Component Universes
 - 6.1 Matter + Curvature
 - 6.2 Matter + Lambda
 - 6.3 Matter + Curvature + Lambda
 - 6.4 Radiation + Matter
 - 6.5 Benchmark Model

7. Measuring Cosmological Parameters
 - 7.1 "A Search for Two Numbers"
 - 7.2 Luminosity Distance
 - 7.3 Angular-Diameter Distance
 - 7.4 Standard Candles and the Hubble Constant
 - 7.5 Standard Candles and the Accelerating Universe

8. Dark Matter
 - 8.1 Visible Matter
 - 8.2 Dark Matter in Galaxies
 - 8.3 Dark Matter in Clusters
 - 8.4 Gravitational Lensing
 - 8.5 What's the Matter?

9. The Cosmic Microwave Background
 - 9.1 Observing the CMB
 - 9.2 Recombination and Decoupling
 - 9.3 The Physics of Recombination
 - 9.4 Temperature Fluctuations
 - 9.5 What Causes the Fluctuations?

10. Nucleosynthesis and the Early Universe
 - 10.1 Nuclear Physics and Cosmology
 - 10.2 Neutrons and Protons
 - 10.3 Deuterium Synthesis
 - 10.4 Beyond Deuterium
 - 10.5 Baryon-Antibaryon Asymmetry

11. Inflation and the Very Early Universe
 - 11.1 The Flatness Problem
 - 11.2 The Horizon Problem
 - 11.3 The Monopole Problem
 - 11.4 The Inflation Solution
 - 11.5 The Physics of Inflation

12. The Formation of Structure
 - 12.1 Gravitational Instability
 - 12.2 The Jeans Length
 - 12.3 Instability in an Expanding Universe
 - 12.4 The Power Spectrum
 - 12.5 Hot versus Cold