

# Life in the Outer Solar System

# Jupiter



Big

$$R = 11R_{\oplus}$$

Massive

$$M = 300 M_{\oplus}$$

= 2.5 x all the rest

Day about 10 Earth hours

Year about 12 Earth years

Thick Atmosphere, mostly  $H_2$ , He  
But also more complex molecules

Colors, storms

Like Miller - Urey

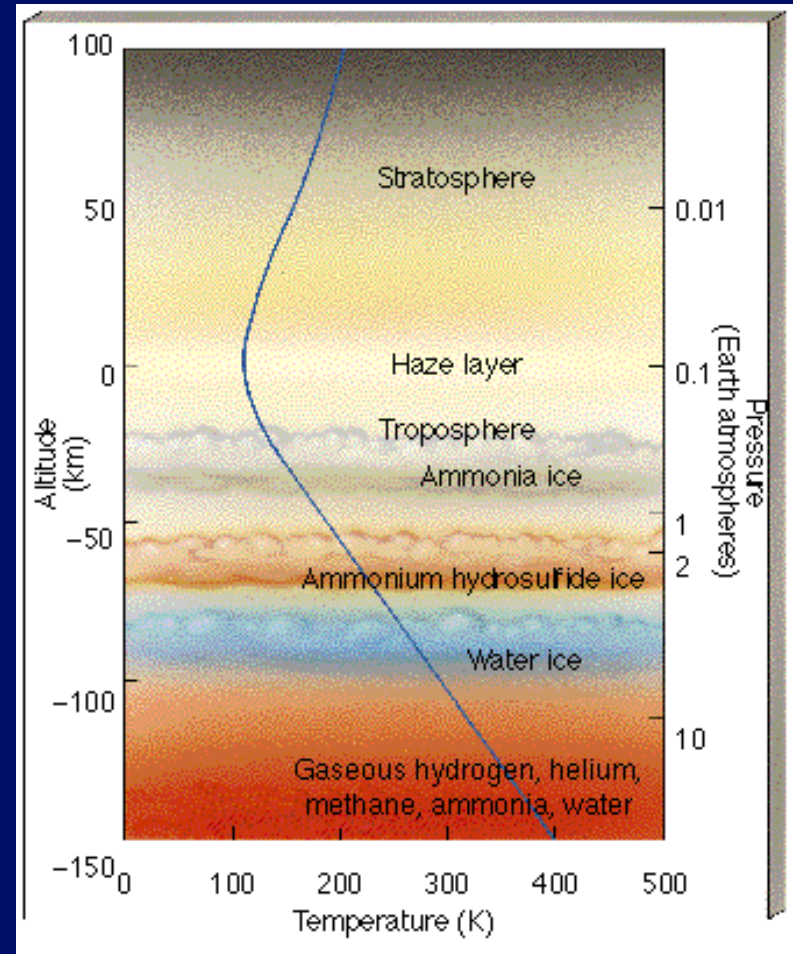
# Life in Jupiter Atmosphere?

Sagan-Salpeter, etc.

Sinkers (Plankton)

Floaters (Fish)

Hunters (Fish)



# Galileo Results on Jupiter

Reached Jupiter Dec. 1995 Sent probe into Jupiter's atmosphere at 100,000 mile/hour

Decelerated at 230 g Lasted for 57 min.

Found: Strong winds

Turbulence, little lightning

Surprise: Little or no H<sub>2</sub>O

May have entered in an unusual place (fewer clouds)

Life less likely?

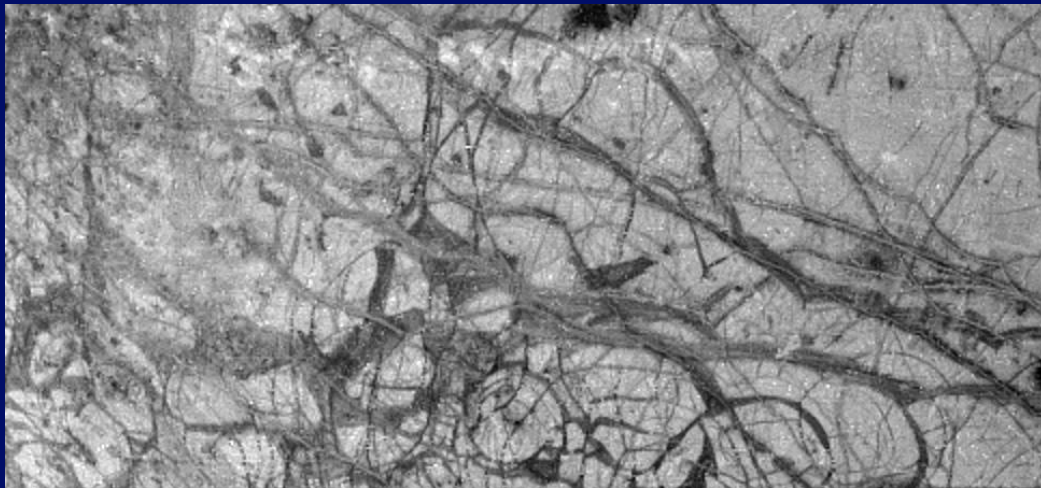


# Europa (Moon of Jupiter)

Surface: Fractured Ice

Subsurface Oceans?

(Heated from Inside)

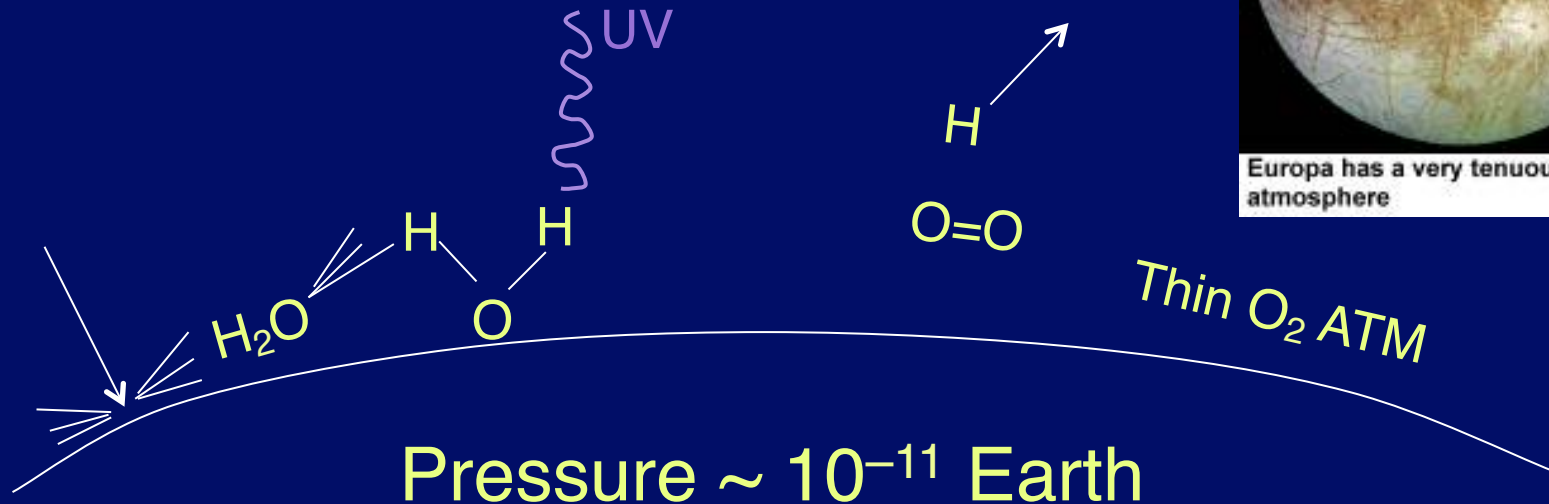
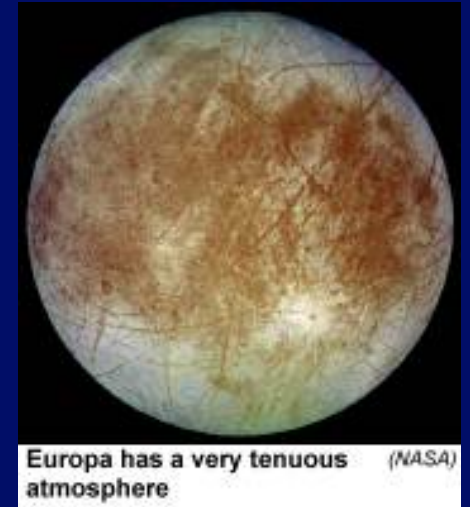


Close-up of "ice floes"

# Galileo - Jupiter's Moons

<http://www.jpl.nasa.gov/galileo/index.html>

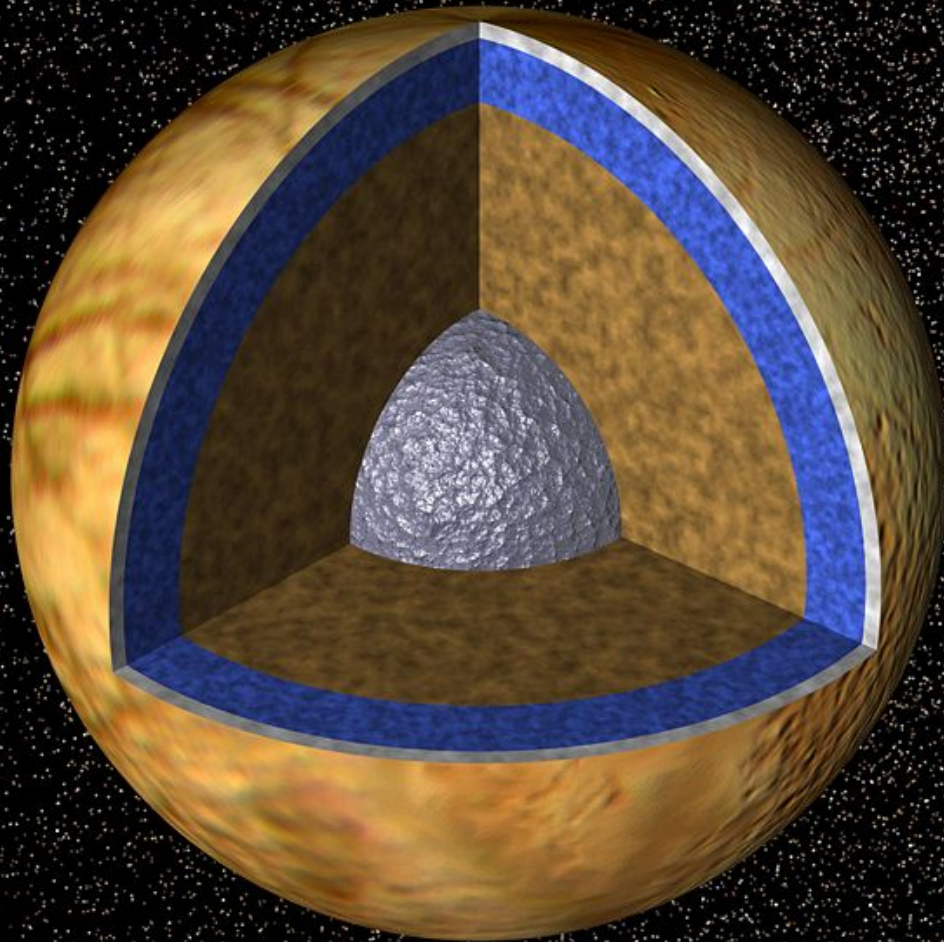
Europa has a (THIN!) atmosphere



More evidence for resurfacing along cracks by  
“ice geysers” → fluid ice or liquid water

Organic molecules on Callisto & Ganymede, maybe Europa?

# Model of Europa's Interior



Ice crust may be 10-30 km thick.

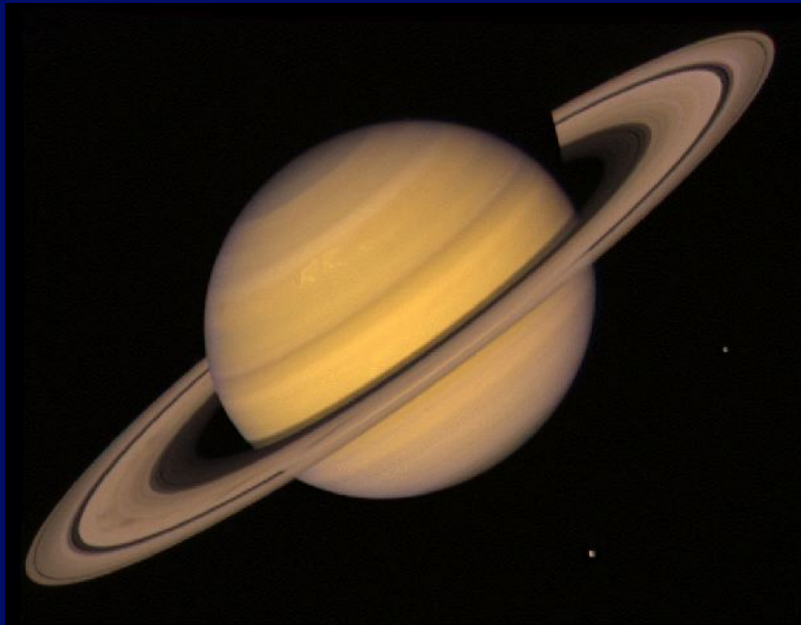
Ocean may be 90 km deep.

## Future Missions

- Juno, launched in 2011, arrive 2016
  - Will go into polar orbit, map gravity field
  - Determine if Jupiter has a rocky core
- Jupiter Icy Moon Explorer (JUICE)
  - ESA Selected in 2012
  - Launch 2022, arrive 2030
  - Ganymede, Callisto, Europa orbiter
  - Look for evidence of organic molecules



# Saturn



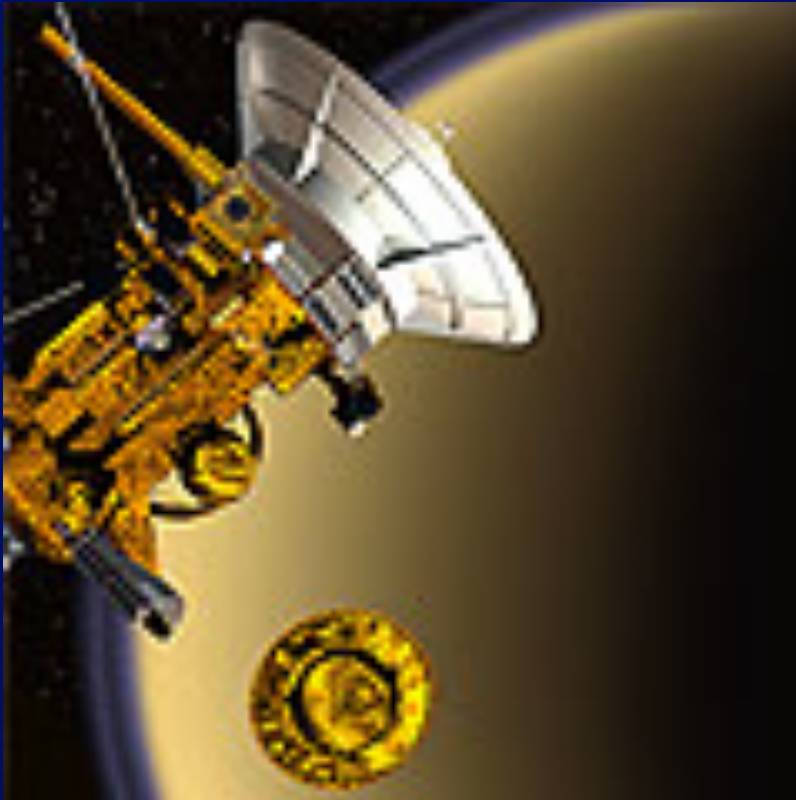
- Big ( $9.4 R_{\oplus}$ )
- Massive ( $95 M_{\oplus}$ )
- Year 29.5 earth years
- Day 0.43 earth days
- Composition similar to Jupiter

# Titan



- Moon of Saturn
- Diameter  $\sim 0.4$  Earth
- Atmospheric Pressure =  $1.5 \times$  Earth
- 85% Nitrogen **BUT**
- Cold ( $\sim 90$  K)
- Subsurface water ocean?
- Reducing atmosphere
- Haze
- Lab for prebiotic chemistry

# The Cassini-Huygens Mission



- Launched 10/13/97
- Arrived Saturn 7/2004
- Cassini studies
  - Saturn
  - Moons
- Huygens
  - Dropped onto Titan
  - Study atmosphere
  - Surface

<http://saturn.jpl.nasa.gov/>

# CASSINI SPACECRAFT



# Huygens Probe



- Released from Cassini
- Slowed by heat shield
- Parachute deployed
- Soft landing
- Sampled gases in atm.
- Results:
  - High winds
  - 430 km/hr at 120 km

# Titan Surface 10km up

- Mosaic of images
- Taken during descent
- Clearly shows features

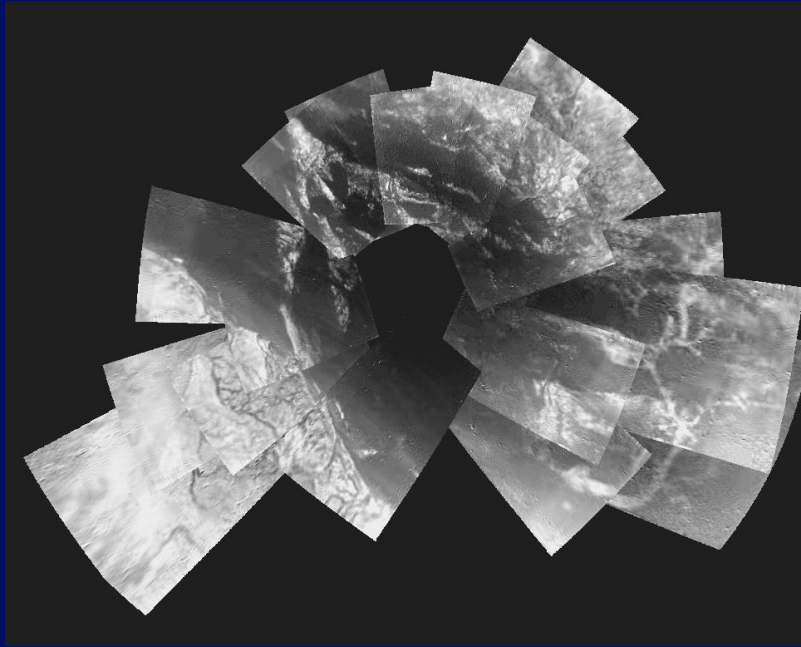
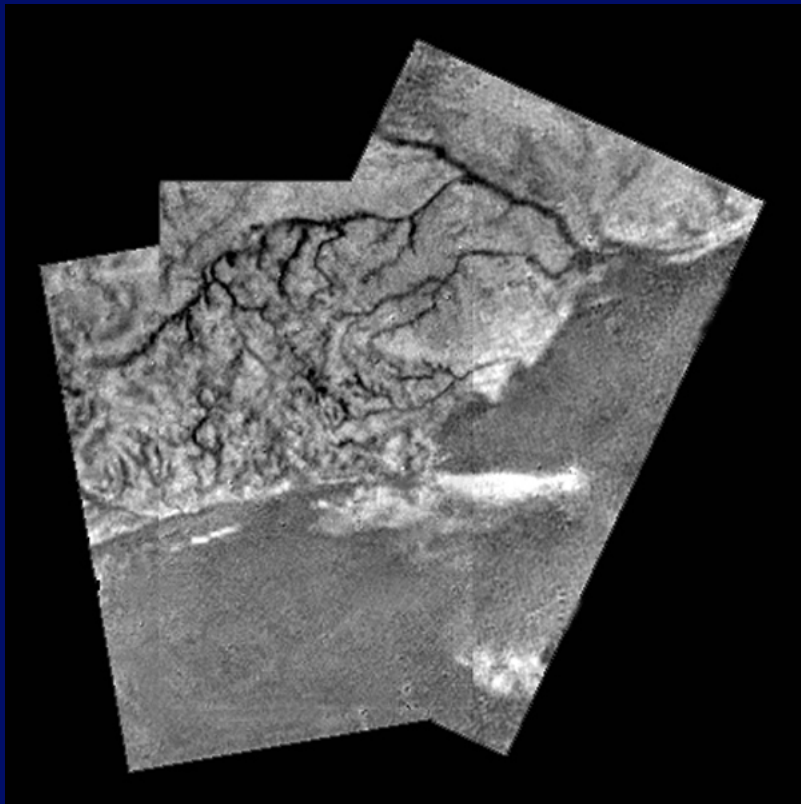


Photo: ESA

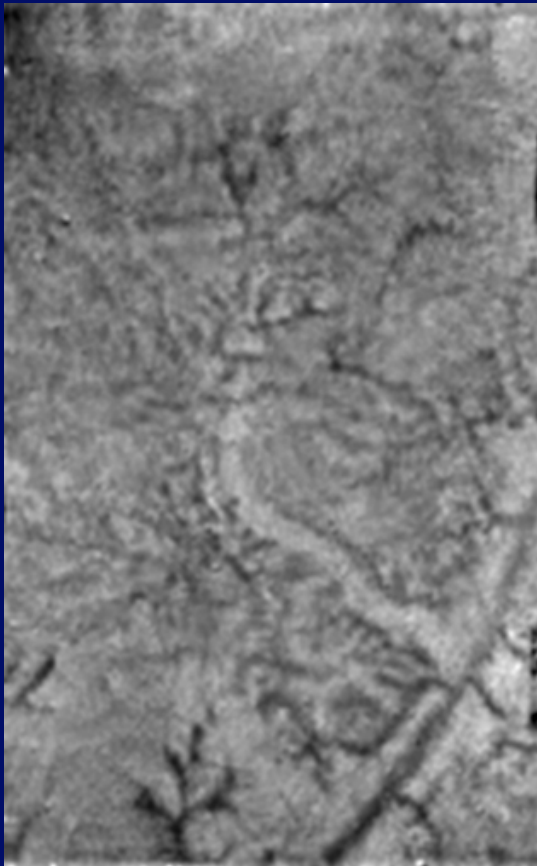
# Titan



- River channel
- Coastline
- Liquid is present
- Methane ( $\text{CH}_4$ )

Photo: ESA

# Water Rift and Methane Springs?



- Straight feature:
- Water ice extruded?
- Stubby channels:
- Methane springs?



## Lakes at northern latitudes

- Radar mapping of northern latitudes (2006)
- Strong evidence for liquid lakes
- And big cloud of ethane ( $C_2H_6$ )
- Ethane raining (or snowing) into lakes

## Lakes and Islands

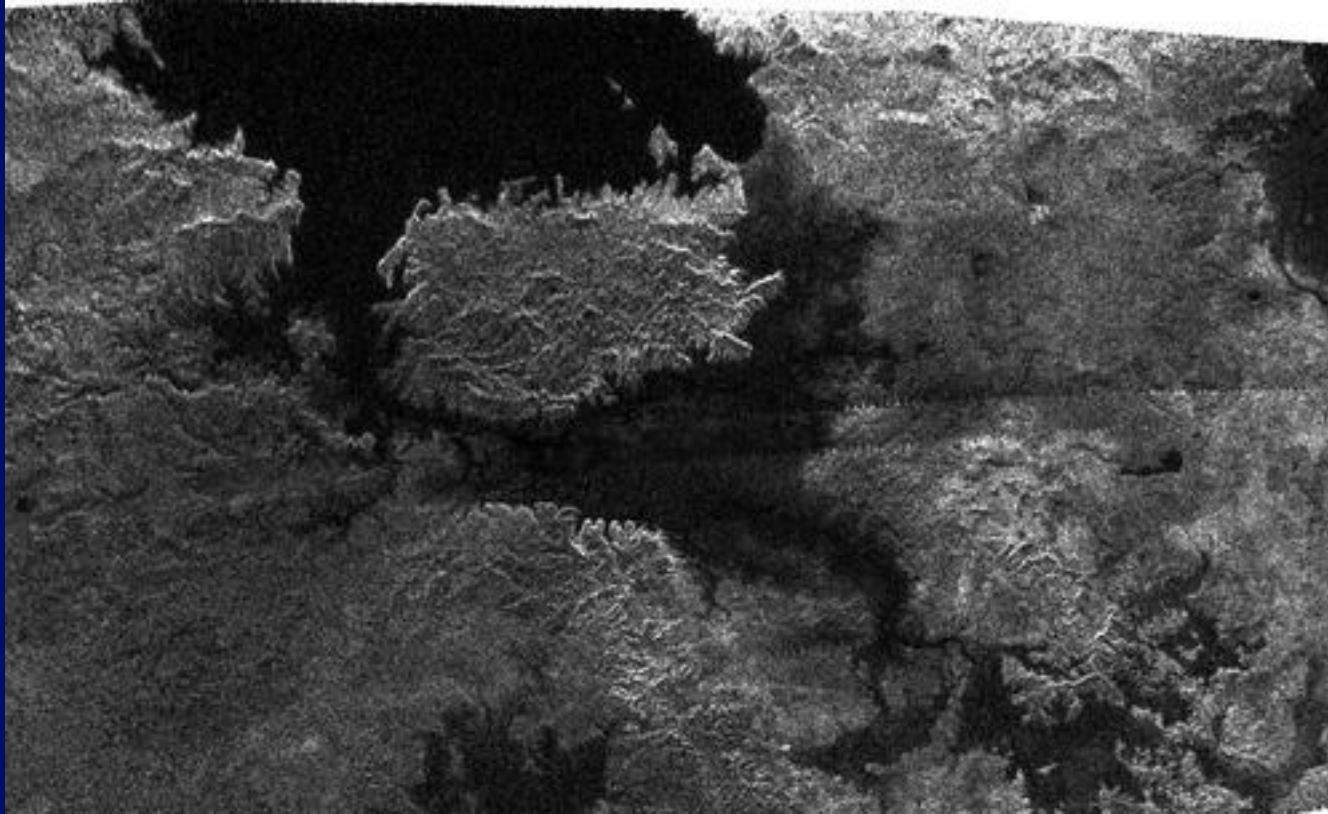


Image from Feb. 2007: based on radar.  
Large lake and island (size of Big Island, Hawaii)  
And smaller lakes

# From the surface of Titan

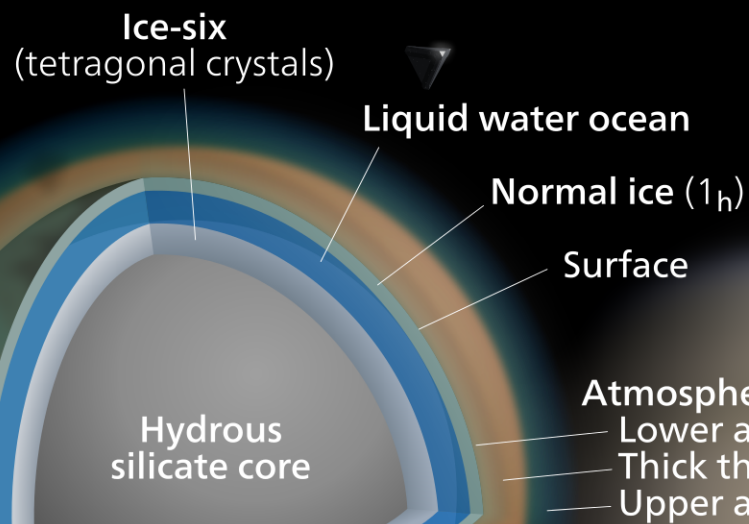


- First view of surface
- “Rocks” of water ice
  - Pebble size (15 cm)
- Surface yielding
- Mixture of ices
  - Water
  - hydrocarbons

## More Titan Results

- Hints of ammonia ( $\text{NH}_3$ )/water ( $\text{H}_2\text{O}$ ) ocean
  - About 200 km under surface
  - Outgassing of  $\text{NH}_3$  may supply  $\text{N}_2$  atm.
- Mapping by radar reveals many lakes and seas of hydrocarbons
  - Seasonal changes in size, depth of a lake
  - Total hydrocarbons on surface about 100 times total oil and gas reserves on Earth

# Model of Titan Interior



# titan

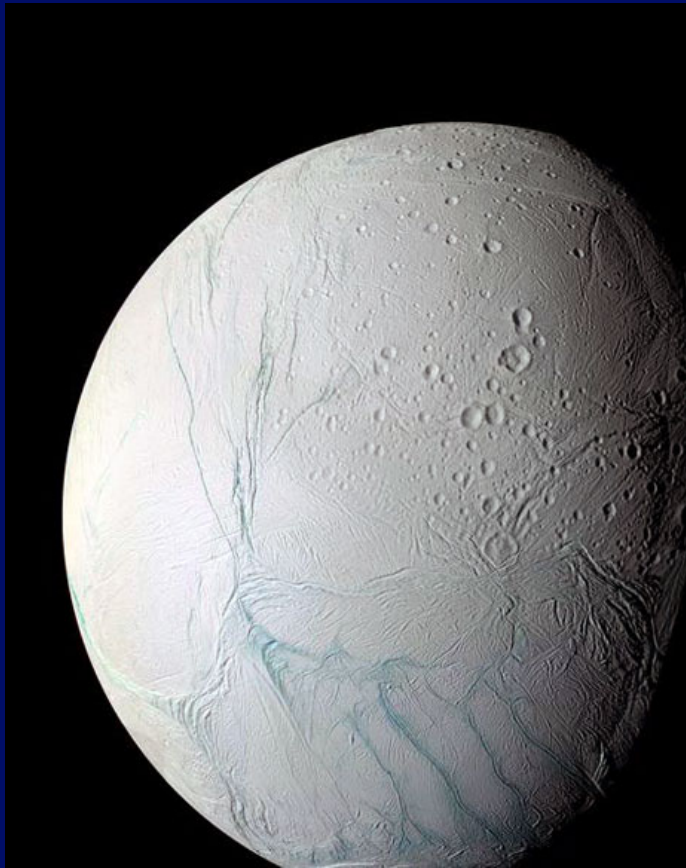
Fully differentiated dense-ocean model  
Drawn to scale

Atmosphere  
Lower atmosphere  
Thick tholin haze  
Upper atmosphere

## Possible Site for Life

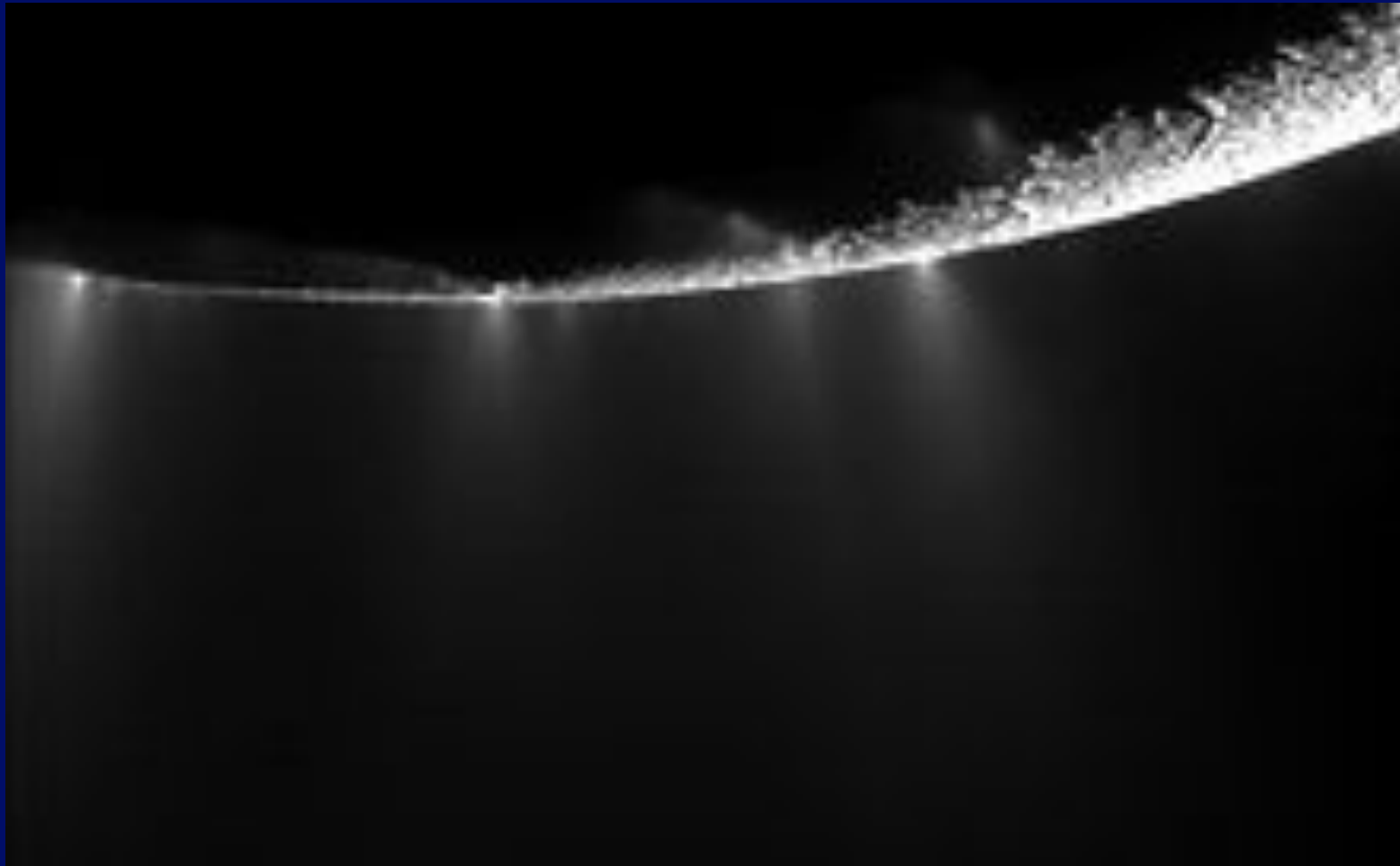
- Miller-Urey type experiments with Titan atm:
  - Formed amino acids and nucleotide bases
- Methane-based life?
- Metabolize with  $H_2$  and  $C_2H_2$ , produce  $CH_4$
- Parallel to  $O_2$  and glucose, produce  $CO_2$
- Also, could produce atmospheric nitrogen

# Enceladus



- Moon of Saturn
- Very shiny
- Part of surface old (craters)
- Part is new, with cracks
- Cassini saw ice geysers (2006)
- Subsurface liquid water
- Source of heat unclear

# Geysers on Enceladus





# How to search for life

Have to decide what test indicates life

Hard to anticipate conditions (recall Viking results)

What about finding “protolife”?

National Academy report - how to search for life

1. Delivery by comets, meteorites e.g. Mars meteorites

2. Sample return - Mars possible

3. Experiments by landers -

Viking on Mars, ...

Future: Europa probe and return?

Titan?

Issues of contamination

4. Biomarkers

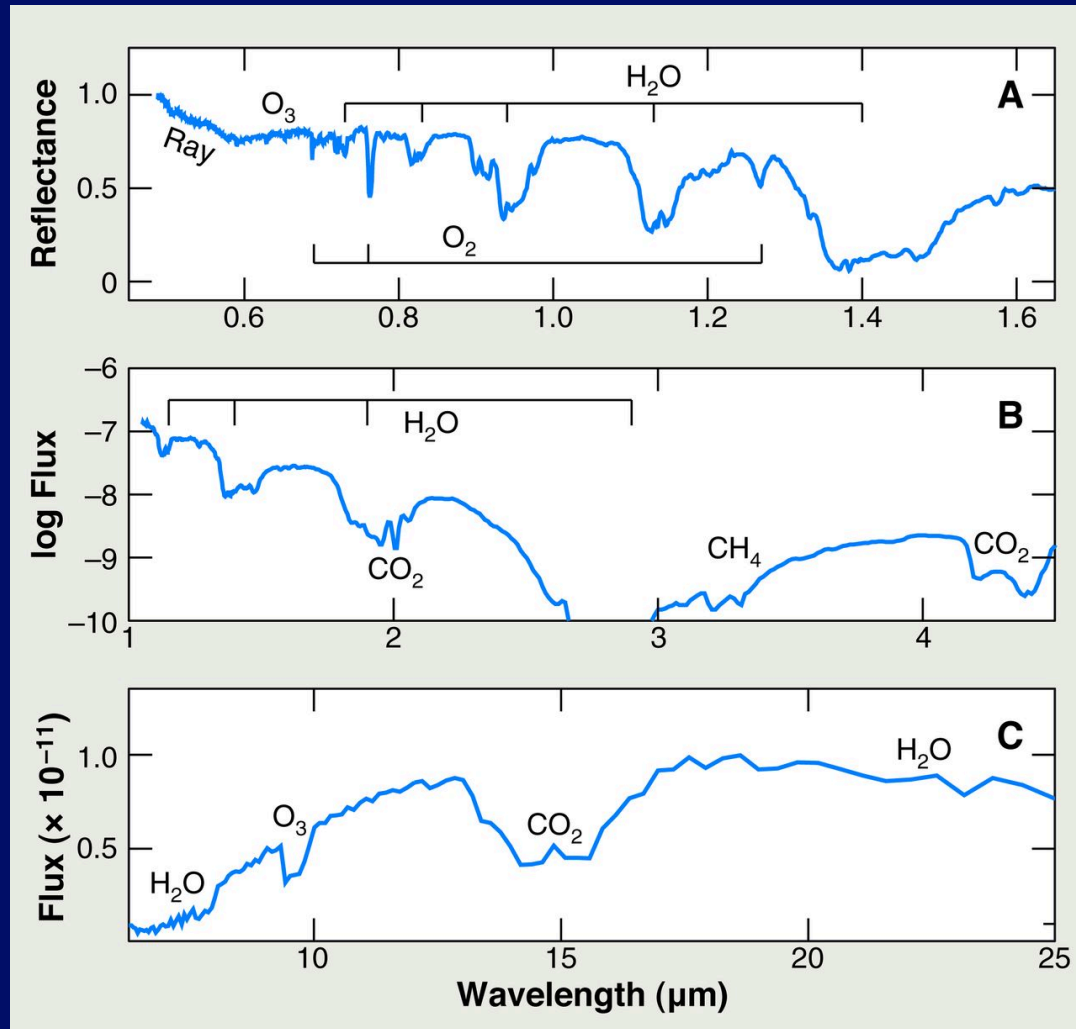
Presence of both  $O_2$  and  $CH_4$  in Earth atmosphere  
indicative of life

How convincing?

## Detecting Life

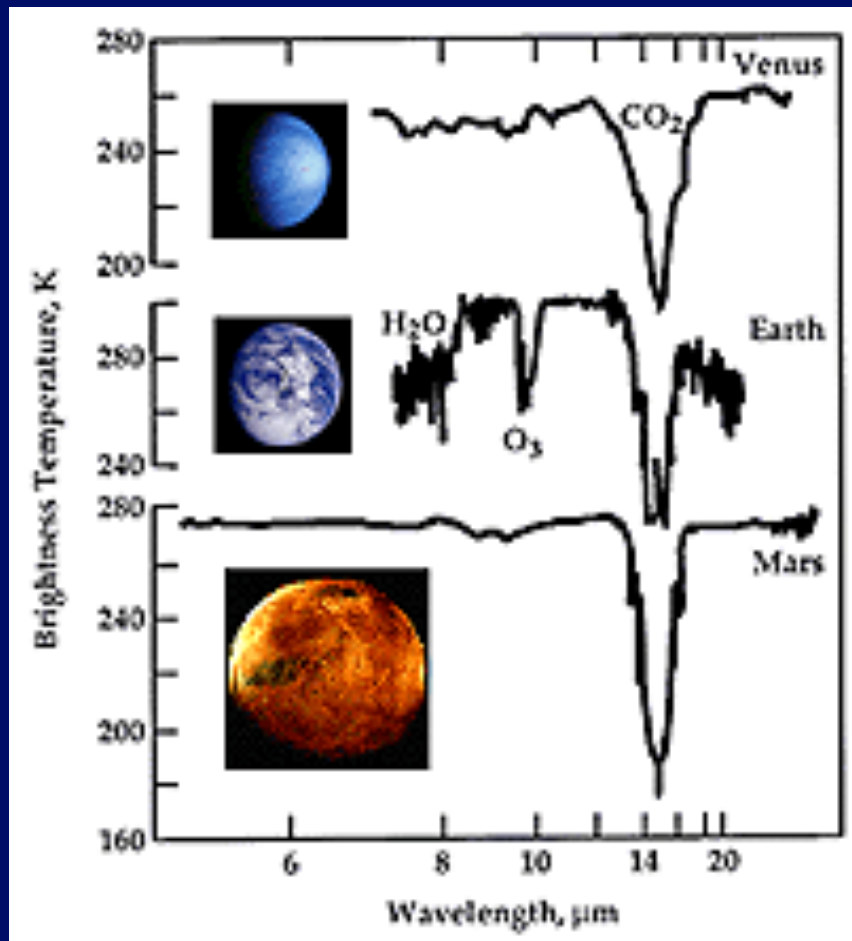
- Suppose we find a planet in habitable zone
- Can we find out if it actually HAS life?
- Radio signals from intelligent life
- But microbes more likely
- “Biosignatures”: gases in the atmosphere that are characteristic of life
  - e.g., Oxygen AND methane

Fig. 3 Earth as an exoplanet, via observed disk-integrated spectra.(A) Visible-wavelength spectrum from Earthshine measurements plotted as normalized reflectance (67).



S Seager Science 2013;340:577-581

# Spectroscopy of atmosphere: simulation



Could be detected with  
future large space  
telescope, but very  
difficult  
Need specialized  
capabilities

# Summary

- Moons of outer planets may extend HZ
  - Subsurface water oceans
  - Methane as solvent (Titan)
  - Closer in giant planets in other systems could have habitable moons
- Searching for life in solar system
  - Planning phase
- Searching for biosignatures
  - Planning phase