

## General Relativity (1916)

- General theory of relativity
  - An extension of “special theory of relativity”, which did not include gravity but dealt only with “inertial motion” (i.e., motion with constant velocity)
  - How do we deal with gravitational force?
- Equivalence Principle
  - Equivalence principle states that “**inertial force** and **gravitational force** are the same thing.”
  - Or, “**acceleration** and **gravity** are the same thing.”
- Flat spacetime to curved spacetime
  - Gravity causes geometry of spacetime to be **curved**.

## Equivalence Principle

- Let’s imagine that you are in an elevator.
  - What happens when the elevator begins to go up?
  - What happens when the elevator begins to go down?
  - What happens when the elevator moves at constant speed?
  - What would happen when the wire hanging the elevator is cut?
- Let’s imagine that you are an astronaut on board a space station.
  - Why do you think you are weightless?
  - What would you do to make “gravity” in flight?

## Acceleration = Gravity

- Newton’s 2<sup>nd</sup> law of motion
  - $F = m a$
- Gravitational force =  $m g$ 
  - $m a = m g$
  - Therefore,  $a = g$ , for a freely falling body
- But, wait. What do we mean by mass?
  - Mass could be measured by its acceleration, given the *inertial force*
  - Mass could be measured by its weight (or gravitational force)
  - But, who said these two “masses” are the same thing?

## Inertial Mass = Gravitational Mass

- Maybe there are two distinctive definitions of “mass”.
- Inertial mass,  $m_{\text{inertial}}$ , may be defined by
  - Inertial force =  $m_{\text{inertial}} a$
- Gravitational mass,  $m_{\text{gravity}}$ , may be defined by
  - Gravitational force =  $m_{\text{gravity}} g$
- Then, Newton’s equation becomes
  - $m_{\text{inertial}} a = m_{\text{gravity}} g$
- Now, let’s drop two balls with different weights (different  $m_{\text{gravity}}$ ) from the roof of RLM and see which one reaches the ground first.
  - (Ignoring friction by air) both balls will reach the ground at the same time. **This must imply that  $m_{\text{inertial}} = m_{\text{gravity}}$**
  - Acceleration would be different for two balls, otherwise!

## Is it surprising?

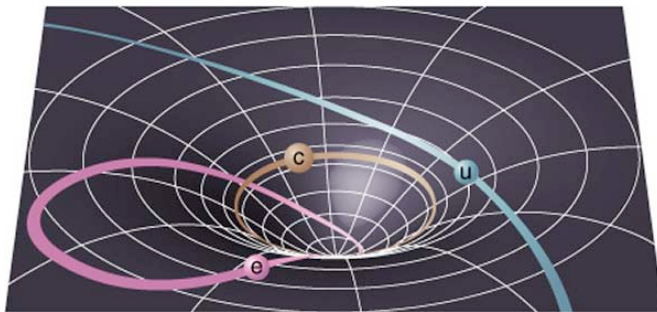
- It's surprising because equivalence principle then states that gravity can be canceled, or mimicked, by acceleration.
  - Case 1: Free-fall (gravity canceled)
  - Case 2: Constant acceleration (gravity mimicked)
- When gravity is canceled by acceleration, one can still use special relativity.
  - Free-falling frame and inertial frame are totally equivalent.

## A way to sense “gravity”

- When gravity is uniform ( $g$  is constant everywhere), its effect can be canceled by a uniform acceleration precisely.
- However, if gravity is non-uniform (which is always true), then there will be a **tidal force**
  - Case 1:  $g$  on Earth depends on altitude
  - Case 2:  $g$  in the Solar System depends on distance from the Sun
  - Case 3:  $g$  on Earth by Moon depends on locations on Earth (cause of tides)
- Let's imagine two balls falling into the center of Earth. What happens?

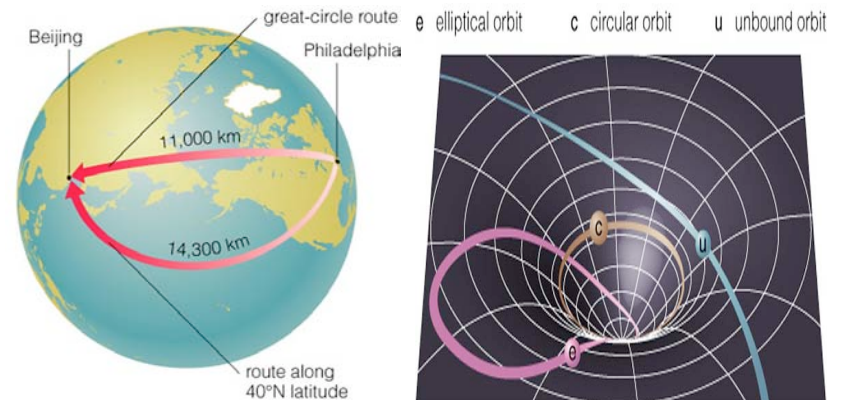
## Spacetime curvature and Gravity

e elliptical orbit    c circular orbit    u unbound orbit



- Gravity is now described by curvature of spacetime
- Spacetime is curved by the presence of energy
- Flat spacetime is equivalent to “zero gravity” = “zero curvature”

## Shortest Paths (Geodesics)

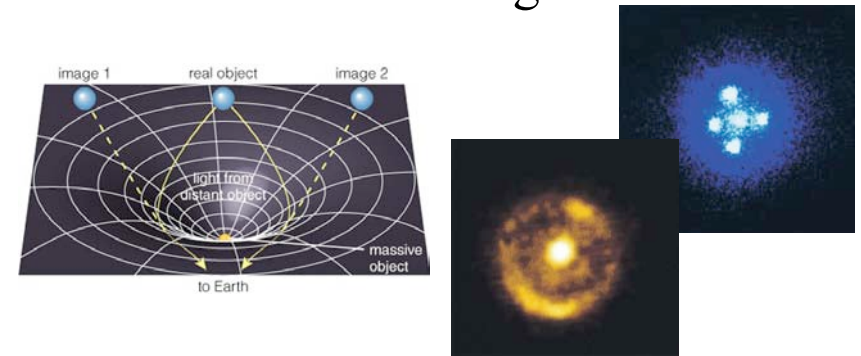


- The shortest paths, geodesics, may look “curved”, but these are the “straight” lines in curved geometry.

## Strength of Gravity

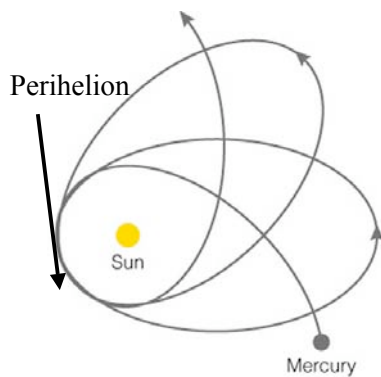
- How do we quantify the strength of gravity of an object?
- How much is the spacetime around an object curved?
  - Gravity is stronger when mass is larger
  - Gravity is stronger when distance is shorter
- Escape velocity
  - Kinetic energy = Gravitational potential energy
  - $(1/2)mv^2 = GMm/R \rightarrow v^2 = 2GM/R$
  - E.g., Sun's gravity: 42 km/s, Earth's gravity: 11 km/s

## Gravitational Lensing



- Deflection of light =  $2 (v/c)^2$  radians
- The Sun will deflect light by 1.75 arc-seconds
  - Measured by Arthur Eddington during solar eclipse in 1919.

## Perihelion Shift



Note: The amount of precession with each orbit is highly exaggerated in this picture.

- Mercury's perihelion was measured to be shifting over years
  - Subtracting precession of perihelion due to gravity from other planets remains the shift of **43 arc-second per century**
- General Relativity accounts for this naturally: perihelion shift =  $(v/c)^2$  radians per revolution
  - $v$  is the escape velocity from the Sun's gravity at Mercury's orbit.

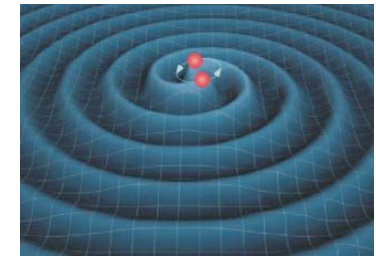
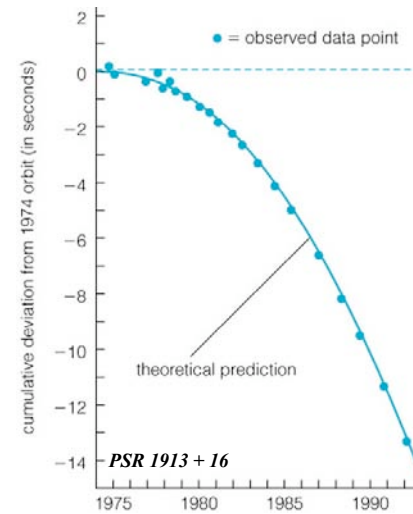
## Gravitational Redshift

- As light escapes from gravitational potential, it loses energy.
- Light with smaller energy has a longer wavelength → Color of light gets redder: **Redshift**
- **1+Redshift**
  - = **Wavelength received/Original Wavelength**
- When gravity is weak ( $v/c < 1$ ), gravitational redshift is
  - Gravitational Redshift  $\sim (1/2)(v/c)^2$
- This result can also be interpreted as “**gravitational time dilation**”
  - $1+\text{Redshift} = \text{Time dilation}$
  - When redshift  $\sim 0.1$  (rather strong gravity!), clock ticks more slowly by 10%

## Gravitational Radiation

- Newtonian gravity propagates instantly.
  - This is in conflict with relativity: nothing can travel faster than light.
- Distortion in spacetime propagates at the speed of light, just like waves in ocean!
  - Distortion propagates as waves (or ripples), just like waves of light
  - Particle-wave duality: There are **gravitons**, which propagate through spacetime at the speed of light, exchanging gravitational force.

## Example: Neutron Star Binary

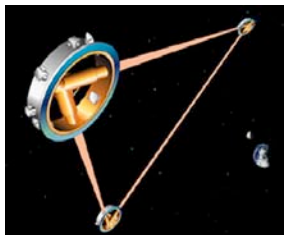


- As neutron stars orbit each other, spacetime is distorted periodically.
- Angular momentum is lost as gravitational radiation is emitted.
- This effect *has been* observed!

## Toward direct detection



LIGO (in operation)



LISA (to be launched)

- As gravitational radiation passes through an object, the shape of the object is distorted.
- “Interferometers” are used to detect such distortion; however, distortion is tiny.
- $h$ =fractional distortion
- $h \sim 10^{-20}$  is typically expected
  - Distortion of Earth is only 1/1000 of the size of hydrogen atom!
  - Distance between Earth and Moon changes by 1/30 of the size of hydrogen atom.
  - Scientists are trying to detect such a tiny distortion.