

Wednesday, November 20, 2013

Exam 4 grades

I mismarked the master on the question about normal or timelike space around a non-rotating black hole, 26(A) or 25 (B).

Everyone who took the regular exam will get 3.03 more points.

We will fix the key.

Chapter 11 - all except Section 11.6; Chapter 12 – all; Chapter 13 – all; Chapter 14 - all

Astronomy in the news?

Maven launched toward Mars.

Goal:

To understand what the “Dark Ages” of the Universe were, why they came to an end, and what gamma-ray bursts have to do with that.

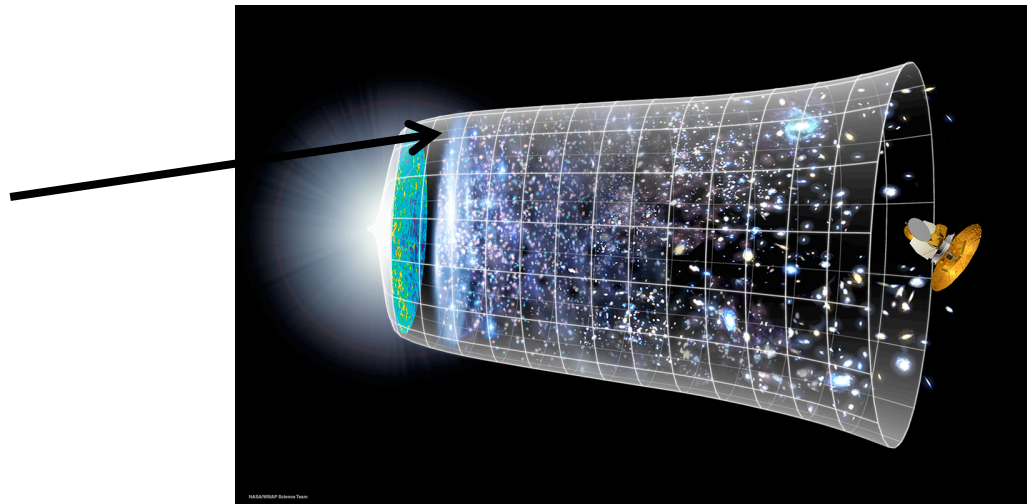
Gamma-ray bursts are intensely bright lights

Can be seen at great distance

Probe cosmology, the early Universe

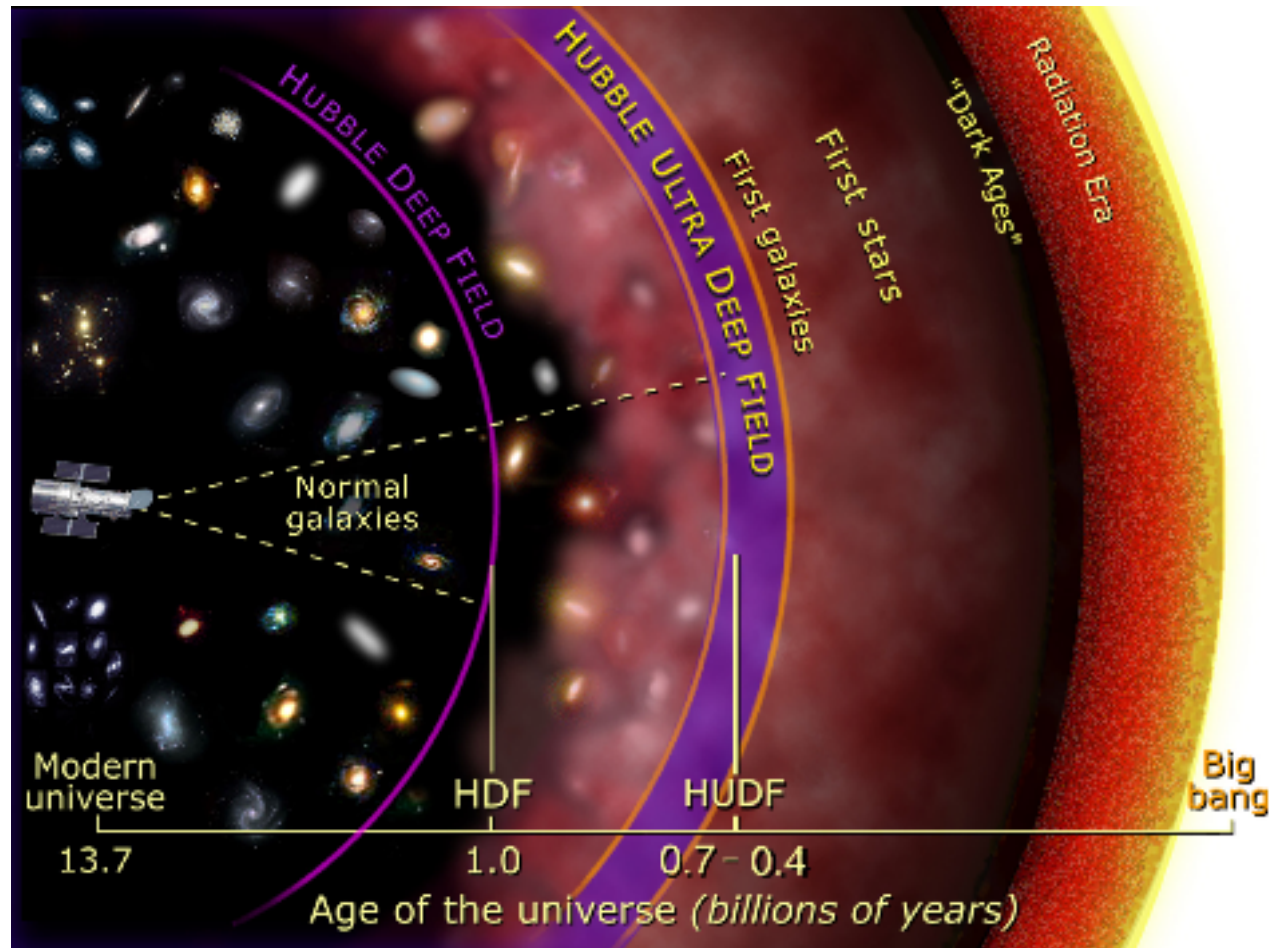
Dark Ages, after the Universe cooled off a million years after the Big Bang, before stars and Galaxies first formed half a billion years later

Dark Ages



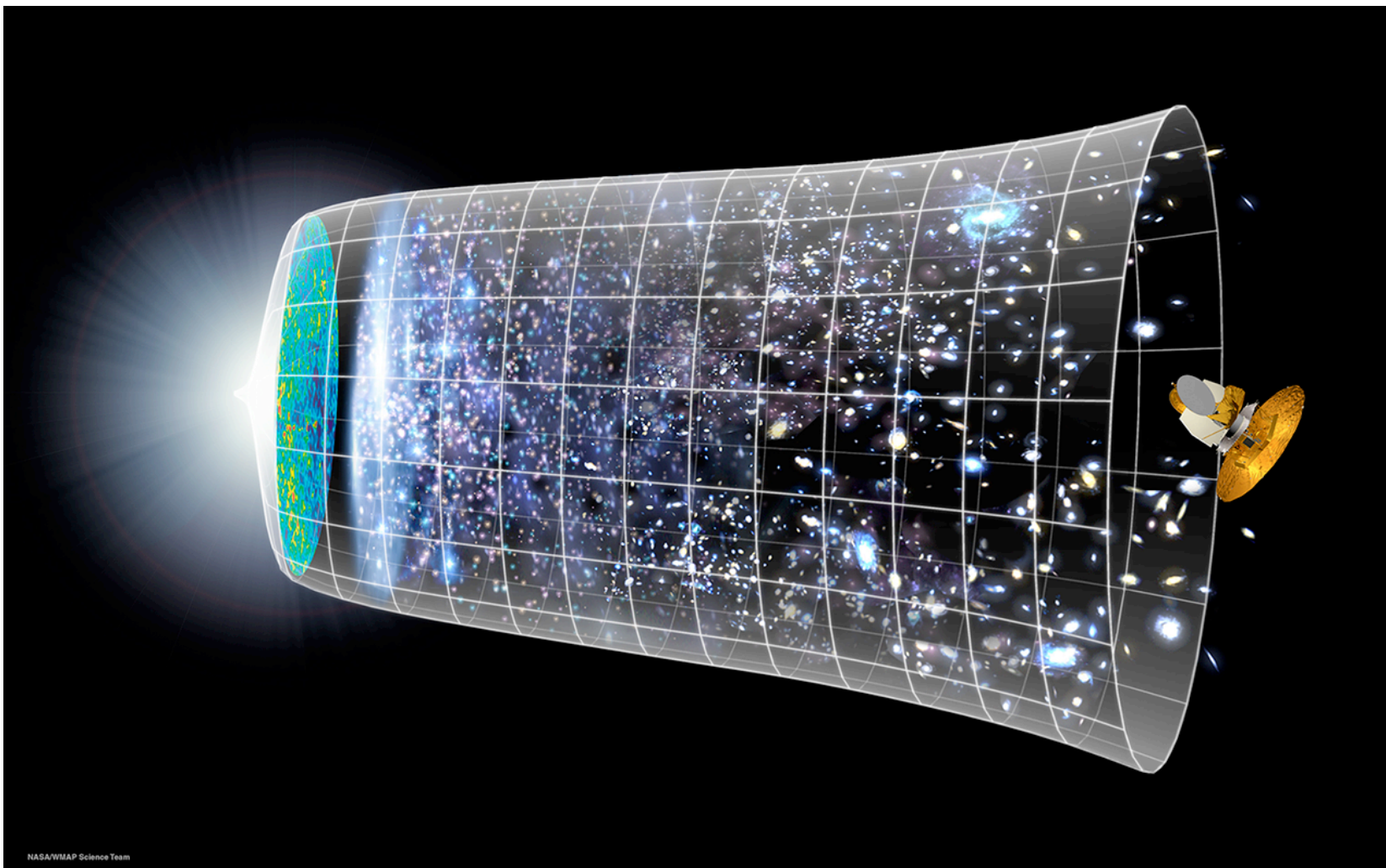
Point toward the Big Bang.

The past is all around us, in every direction, out in space, back in time.

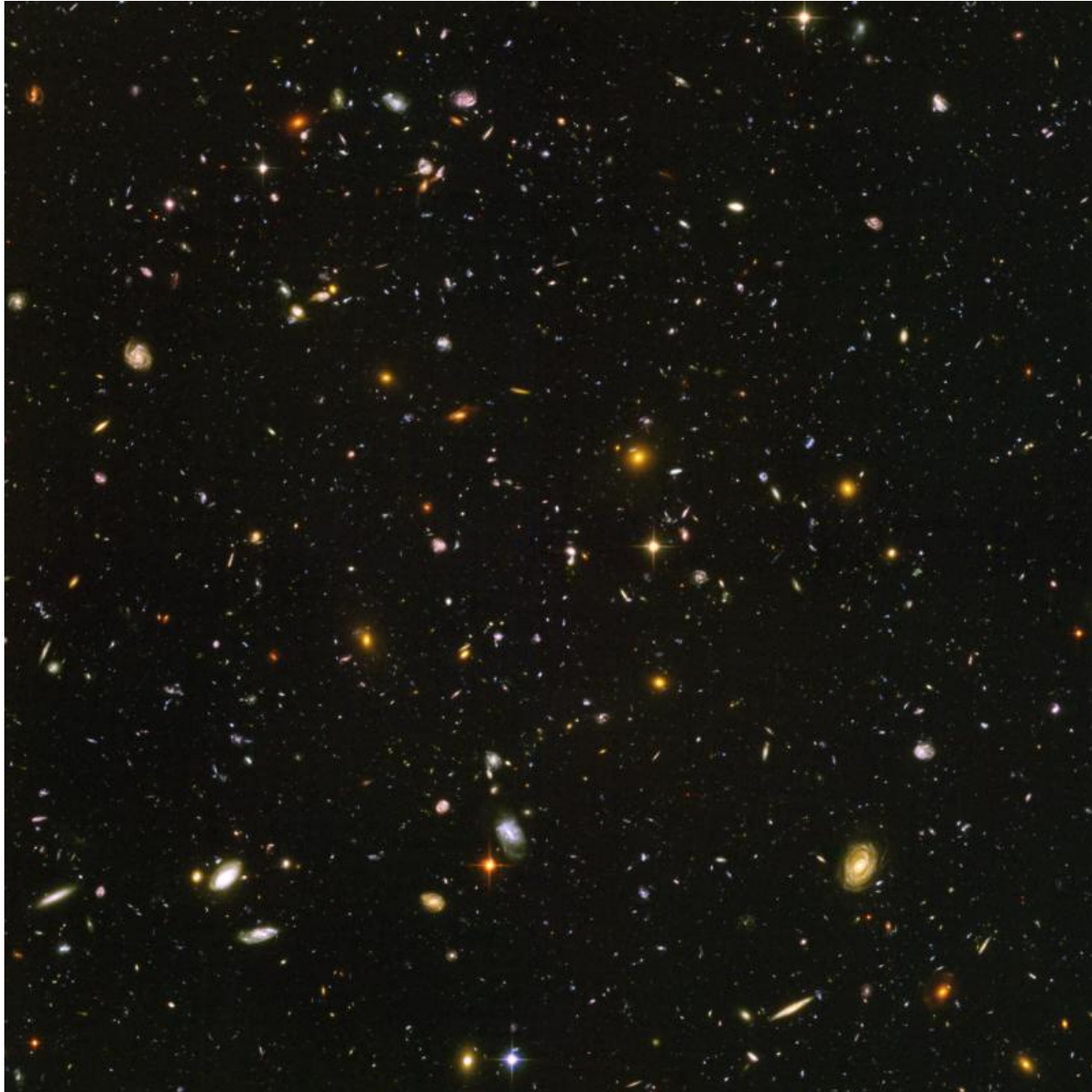


Gamma-ray bursts could be among the first objects seen at the end of the Dark Ages as the first stars are born and die, over 13 billion years ago. GRB 090423 is the first example.

From the Big Bang to Now



Hubble Deep Field – every speck a galaxy



Goal:

To understand the origin, shape, and fate of the whole Universe and how Type Ia supernovae have helped to revolutionize that understanding.

Chapter 12 Supernovae and the Universe

Expanding Universe - we observe all distant galaxies (so far away we cannot sense their individual gravity) moving away from us with speed proportional to distance: as if we were in the center of an explosion.

Our Universe is not a bomb in pre-existing empty 3-D space!

Lesson from Einstein - *space itself can expand carrying the* (almost motionless) *galaxies*

All distant galaxies move away from all other distant galaxies.

No galaxy, certainly not us, is in the center.

The result: speed proportional to distance

Review:

What is hyperspace?

What is an embedding diagram?

Our Expanding Universe

Expanding *surface* of a balloon as an example

Expansion takes every point on the surface further from adjacent points.

2D embedding diagram of a 3D expanding Universe

No 2D center, no 2D edge, no 2D outside to the 2D surface

There is a 3D center, a 3D edge, a 3D outside, in 3D hyperspace

Our Expanding Universe

All 3D space expands - carrying essentially motionless matter
(galaxies)

No 3D center, no 3D edge, no 3D outside

As 3D astronomers, we don't have to ask what the Universe is expanding into, but if anything it is a 4 (or more) D hyperspace, just as a 2D balloon expands into 3D hyperspace.

Infinite flat rubber sheet could expand without expanding into any hyperspace (2D embedding diagram example of how expansion does not necessarily mean intrusion into hyperspace)

Einstein's theory says that for a Universe that is the same, on average, everywhere, there are only three basic shapes it can have

The 3D analog of a spherical surface - *Closed Universe*

The 3D analog of a “saddle” or “Pringle” shape - *Open Universe*

The 3D analog of a flat plane - *Flat Universe*

The *2D embedding diagrams* of these 3D Universes are, respectively, *a sphere, an infinite saddle or Pringle, and an infinite flat plane.*

A closed universe is finite in space and time, the other two are infinite in space and time, but all must have started 13.8 billion years ago in the Big Bang.

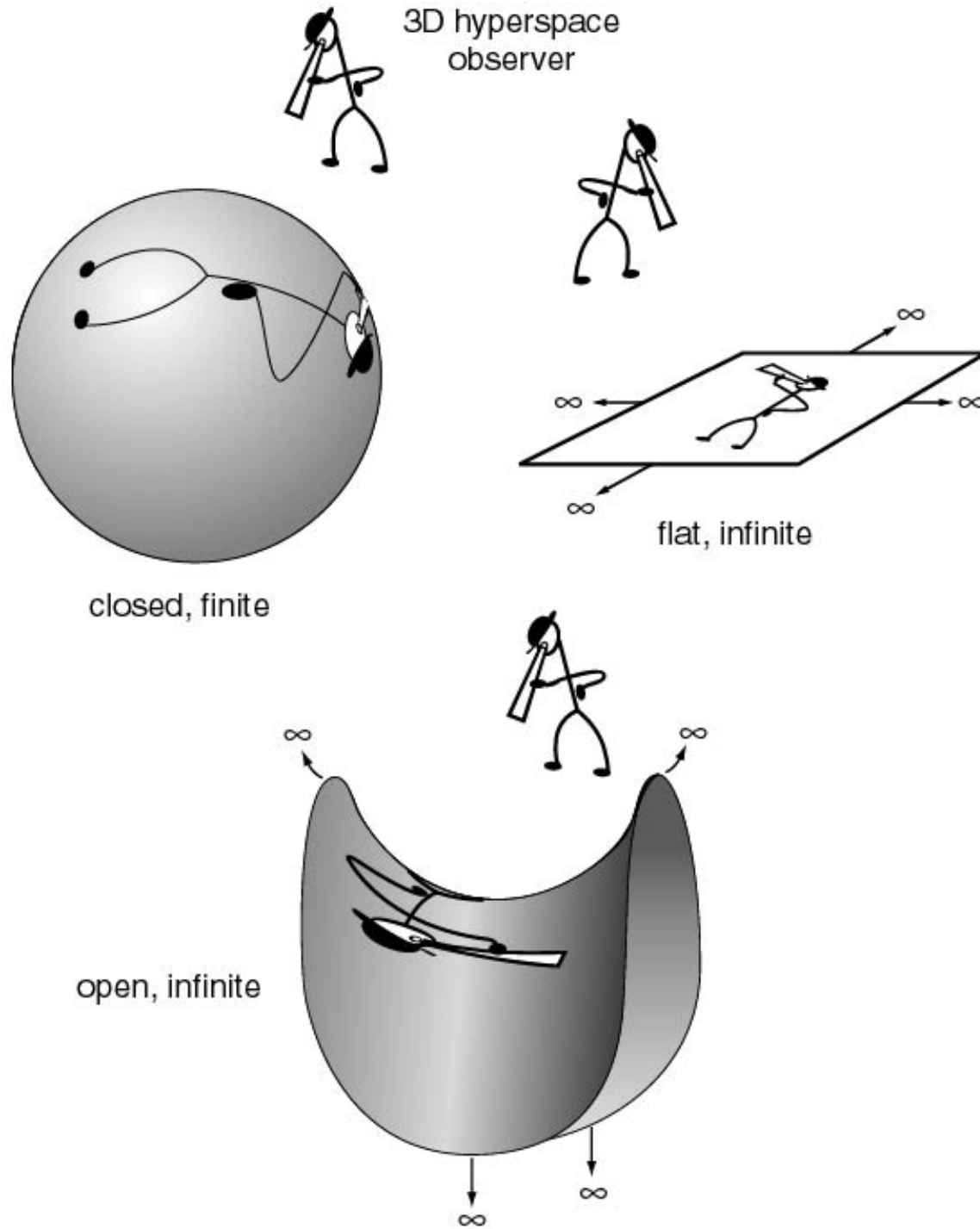


Fig. 11.1

2D
embedding
diagrams of
possible
shapes of our
3D Universe

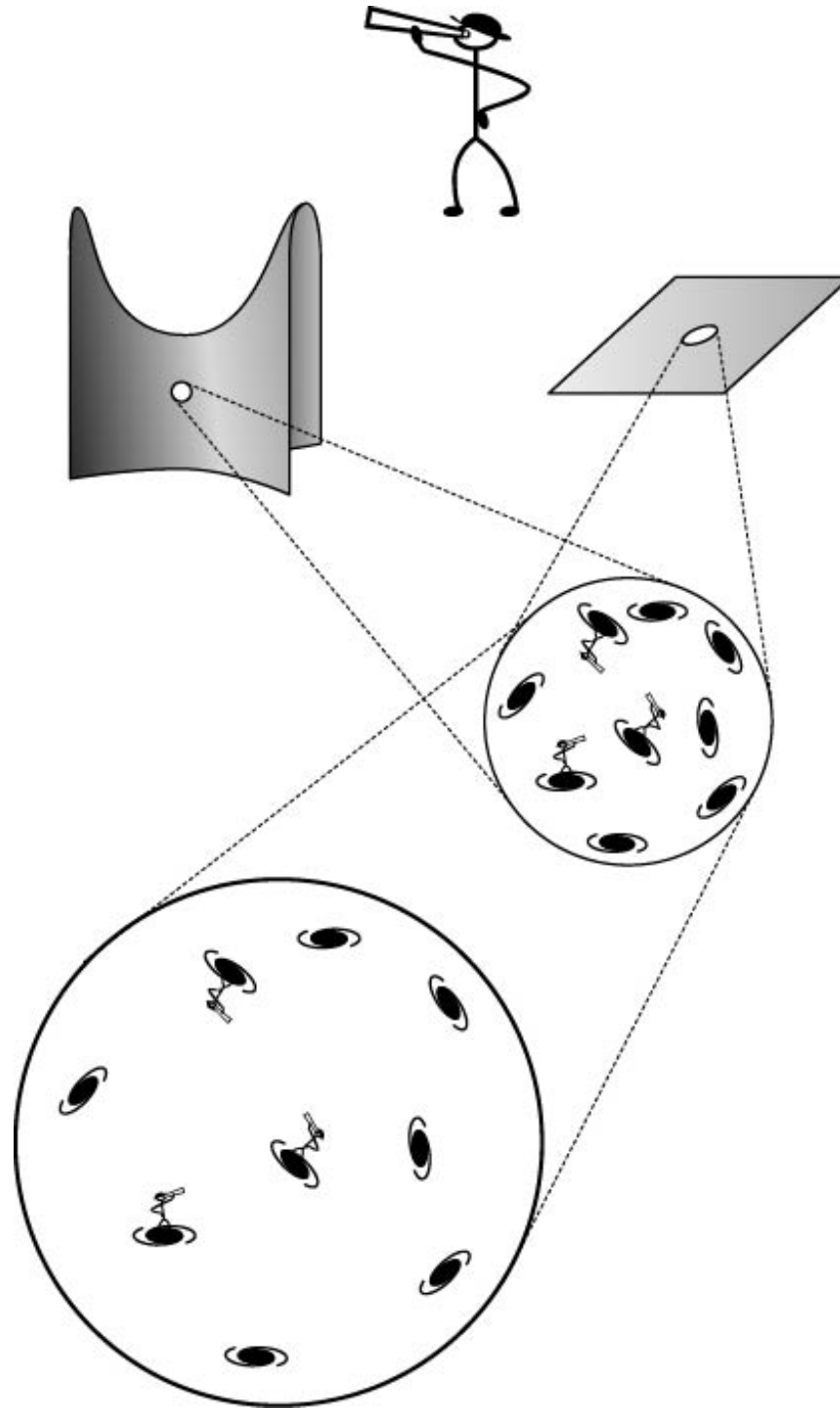


Fig 11.2

A patch of the space in a universe expands, drawing all galaxies away from all others, independent of the overall shape of the curvature of the universe.