 The pressure at the center of a. nuclear reactions. the rate at which photons can c. the fact that pressure balance d. the temperature and pressure 	the sun is determined diffuse through the o s the weight (due to gr	by verlying layers.	·	ndy have the review sheet)
 How can we study the interia. Gamma rays from the sun's c Analyze spectral lines from t Interpret the complex vibration There is no other way. 	ore. he photosphere to estim			nt elements.
3. If we know the apparent brig a) distance b) ra	htness of an object, it dial velocity c) surfac		be calculated if we al d) mass	lso know its
4. If a star has a parallax of 0.05 a) 0.05 b) 5		its distance in par d) 20	secs is: e) 200	
5. The radius of a star can be es information is available.)a) parallax and spectral typec) mass and temperature	b) tempe	and erature and lumino and luminosity		known. (Assume no other
6. A certain star is observed to l From this we can infer that the a) more massive than the sun.c) probably a young star.[Can you explain this in words?]	star is	ature of about 20,0 b) less massive th d) probably a wh	nan the sun.	y equal to the sun's luminosity.
7. In the spectrum of a spectral a. neutral metals b. io		d expect to find s c. molecules	trong lines caused by d. neutral hydrogen	
8. The importance of the Hippaa. obtain spectroscopic parallaxb. obtain stellar distances muchc. measure the diameters of a finance	es of stars. larger than was possil	ole previously.	tars.	
9. What is the approximate lifeta) a few million yearsc) a few billion yearsd) 10		ve main sequence b) 100 million ye e) several trillion	ars	
10. Most spectral lines of inters a) ultraviolet b) vi	tellar molecules are in sible	the c) infrared	part of the spectrum. d) radio	[WHY?]
11. Emission nebulae like the Ca) red supergiantb) red ma[Try to explain "to someone" w	in sequence	c) massive	d) old	
12. Interstellar molecules are ob a) electron recombination b) vi			y onal changes	 d) spin-flip transition
13. The most abundant molecul a) NH ₃ (ammonia) b) C	e in a molecular cloud O (carbon monoxide)		d) H_2 (molecular h	ydrogen)